

DAMA/LIBRA

- DAMA

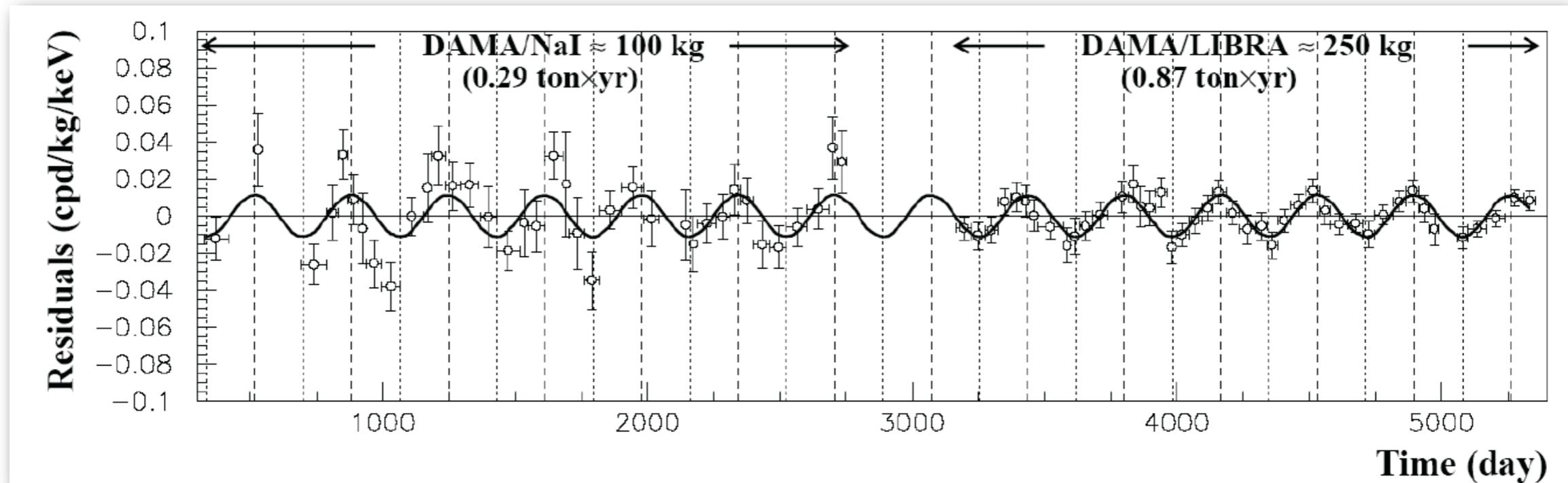
- 100 kg NaI array operated from 1996 - 2002 in Laboratori Nazionali del Gran Sasso.
- Measures scintillation from particle interactions in detectors.
- No discrimination between nuclear and electron recoils
- Positive results reported in 1998.



- LIBRA

- 250 kg array operating since 2003 with first results in 2008.

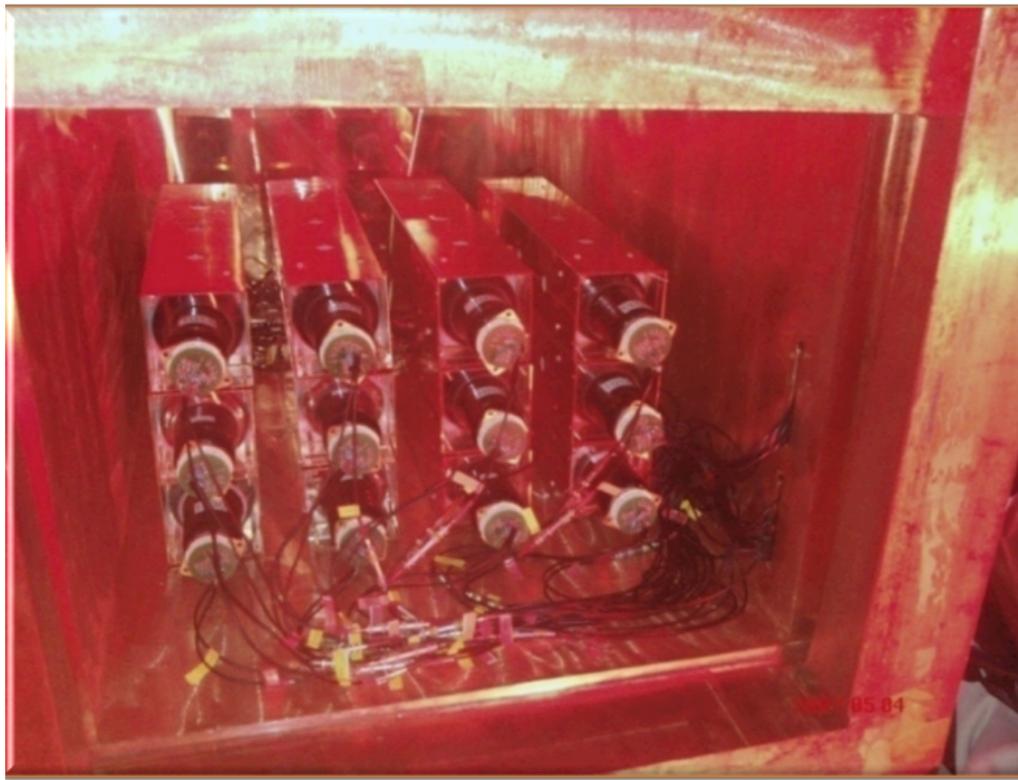
DAMA/LIBRA



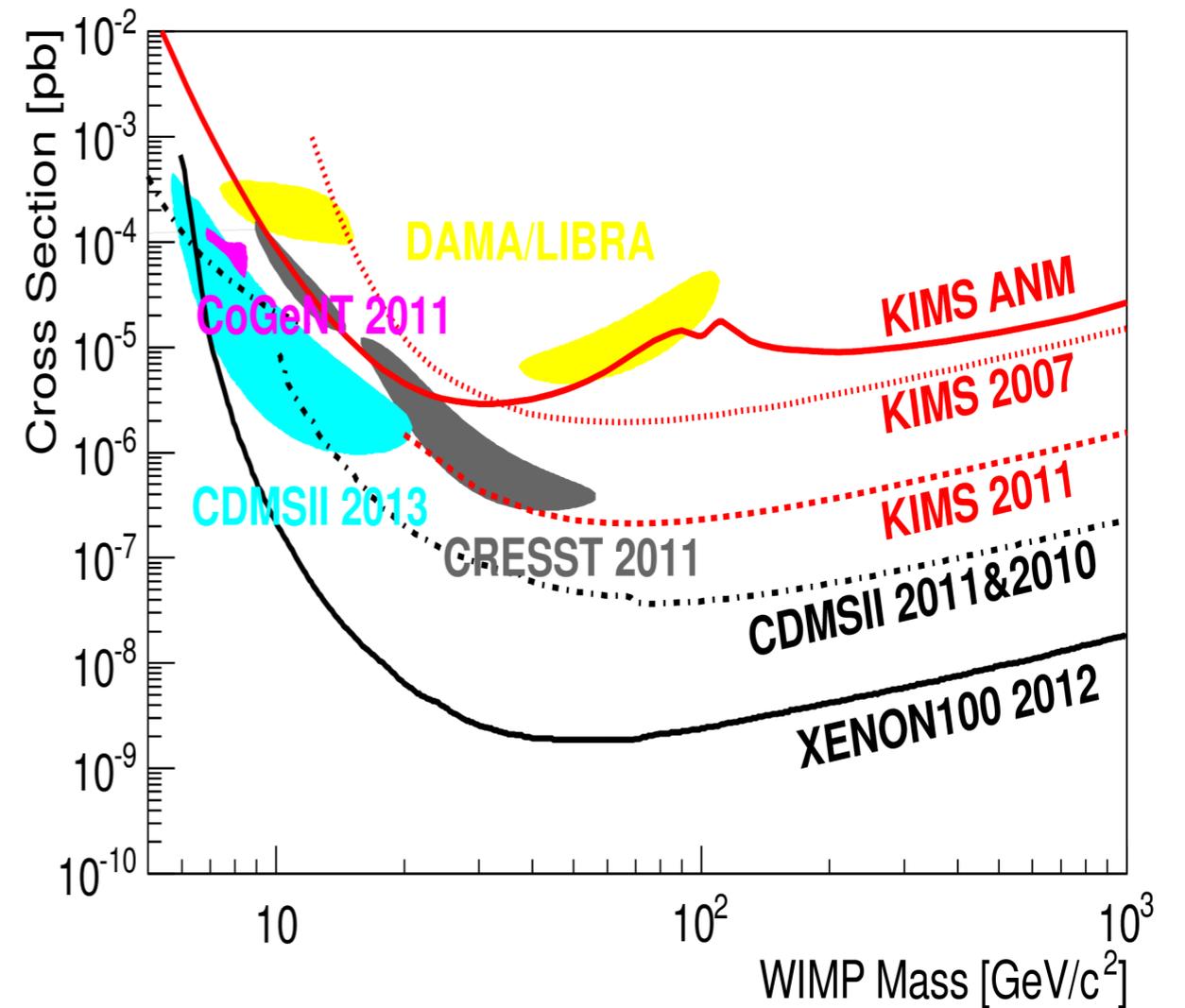
- Modulation has been observed over 13 cycles.
- Significance is 8.9σ .
- Signal is observed only in lowest energy bin.

KIMS

- Direct comparison to DAMA annual modulation signal using CsI(Tl) crystals
 - Pulse shape discrimination also possible
- 12 crystals (104.4 kg) installed
- Data taking from Sept 2009 - Feb. 2012



< SI WIMP-nucleon Cross Section >



- Pulse shape discrimination excludes DAMA/LIBRA - **PRL 108, 181301 (2012)**
- No annual modulation is observed.

KIMS

- **KIMS-CsI:** Upgrade of CsI(Tl) crystal detector

- Lower threshold $\sim 1.5\text{keV}$, $< 1\text{dru}$, counts/(keV kg day).

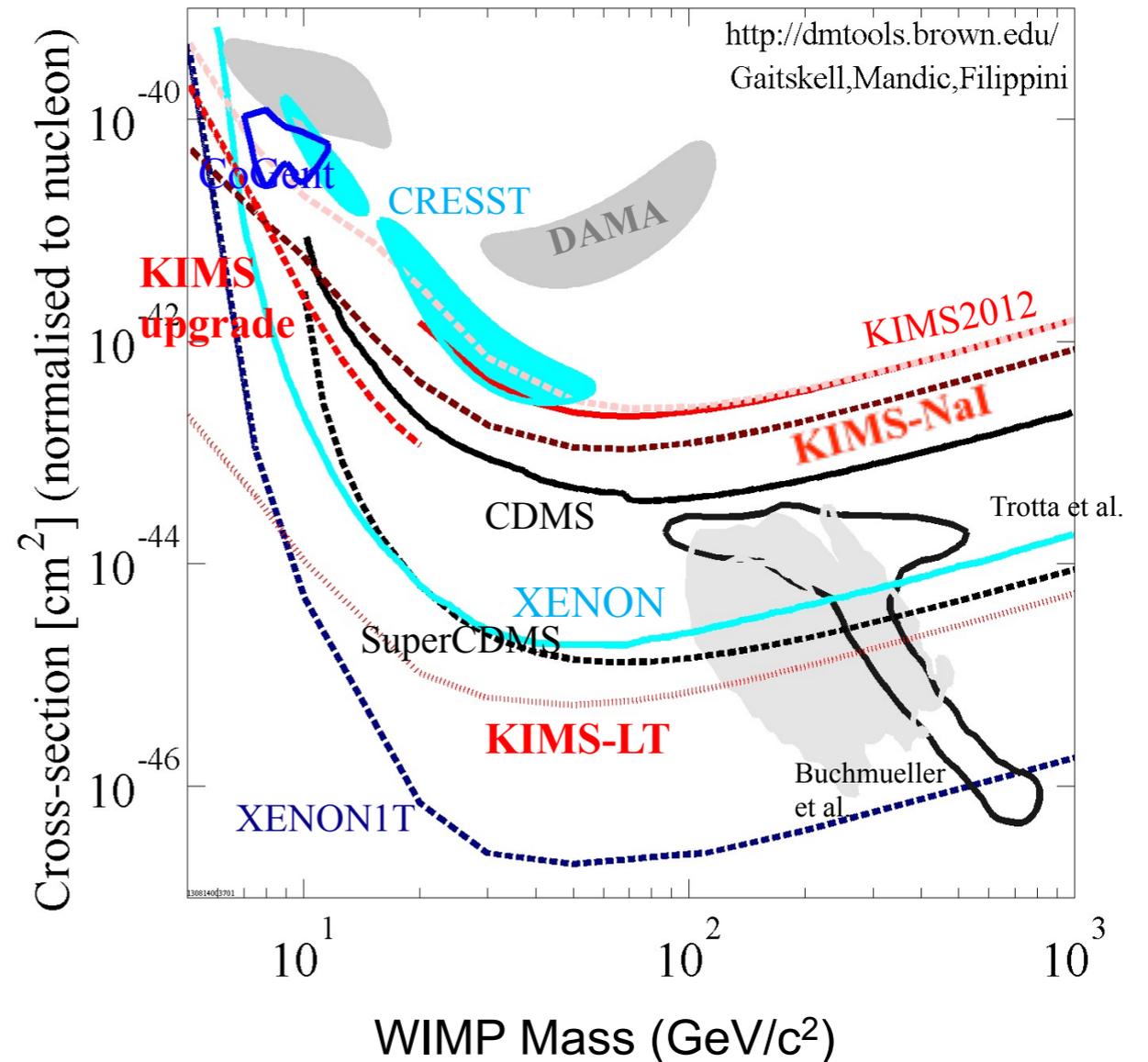
- **KIMS-NaI:** new NaI(Tl) detector

- Duplicate DAMA experiment with ultra-low background NaI(Tl) crystals.

- 200kg run in 2015-2016

- **KIMS-LT**

- Use scintillating bolometer such as $^{\text{nat}}\text{Ca}^{\text{nat}}\text{MoO}_4$ crystals $\sim 200\text{ kg year}$.
- High sensitivity to low mass WIMP.
- Operations in 2019-2022

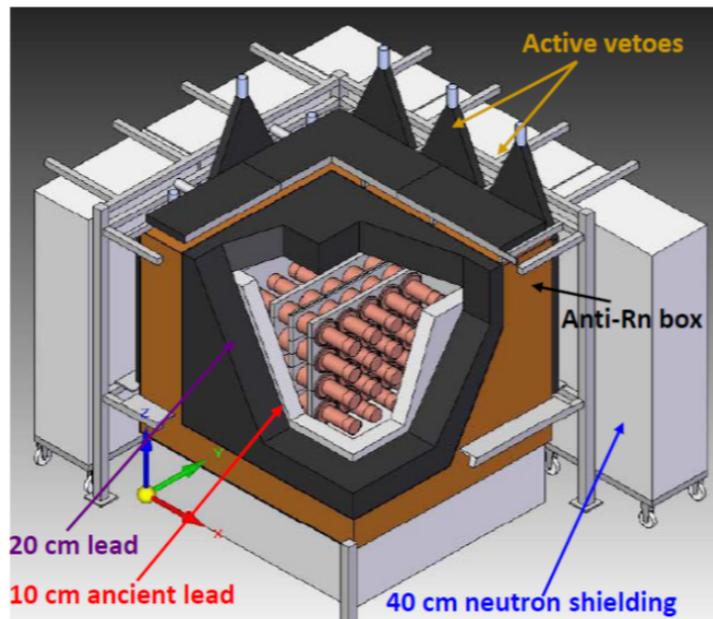


ANAIS

Goal:

250 kg of NaI(Tl) detectors to study the annual modulation effect at the Canfranc Underground Laboratory, LSC (Spain)

Same target and technique used by DAMA/LIBRA



2 prototypes taking data at LSC – 25 kg (ANAIS-25)

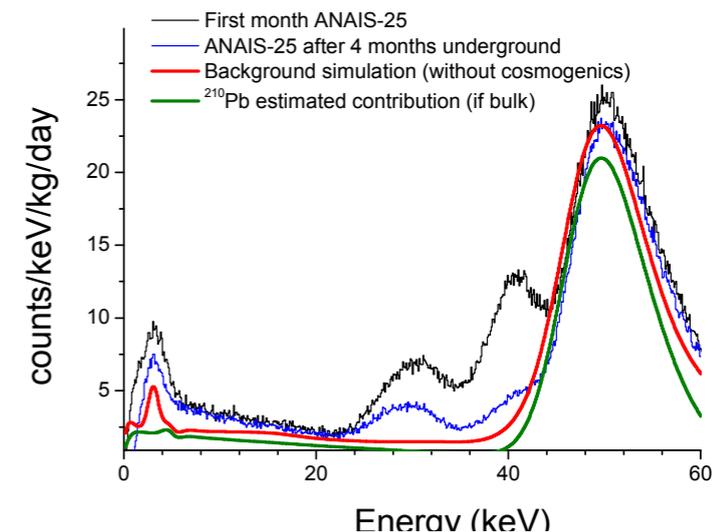


Ham PMT – R12669 SEL coupled at LSC clean room



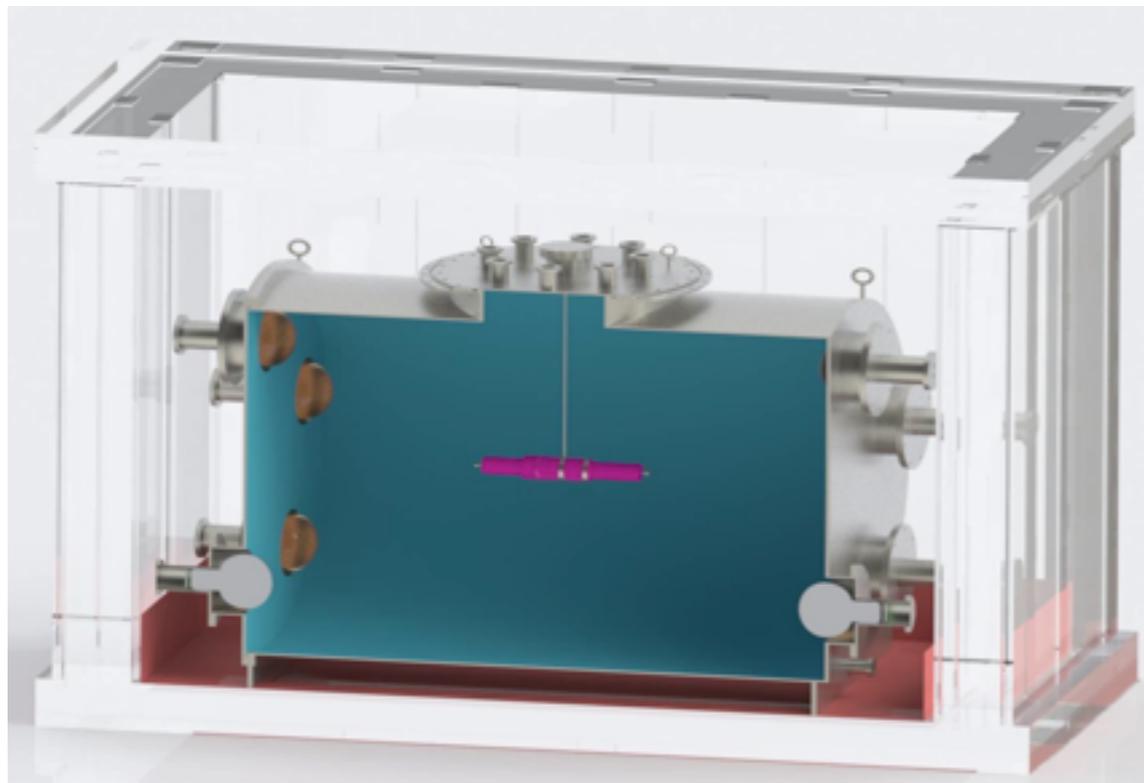
ANAIS-25 PRELIMINARY RESULTS

- Excellent light collection efficiency: 12-16 phe/keV
- ^{40}K bulk NaI content = 41.7 ± 3.7 ppb
- Problem with ^{210}Pb too high level
- Cosmogenic activation still decaying & low energy events selection not yet fully developed

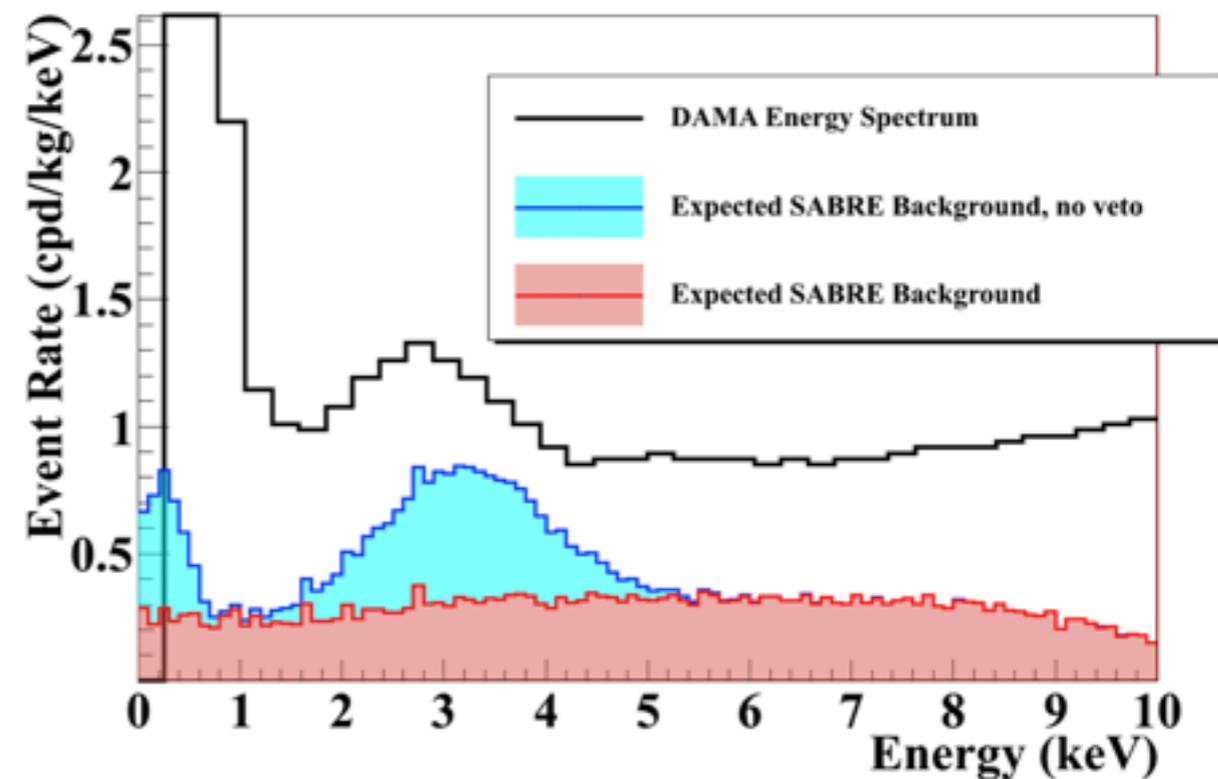


SABRE

- Sodium-iodide with **A**ctive **B**ackground **R**Ejection
- Plan to use an active muon veto to drastically reduce the number of gamma backgrounds



SABRE Expected Background with DAMA Energy Spectrum

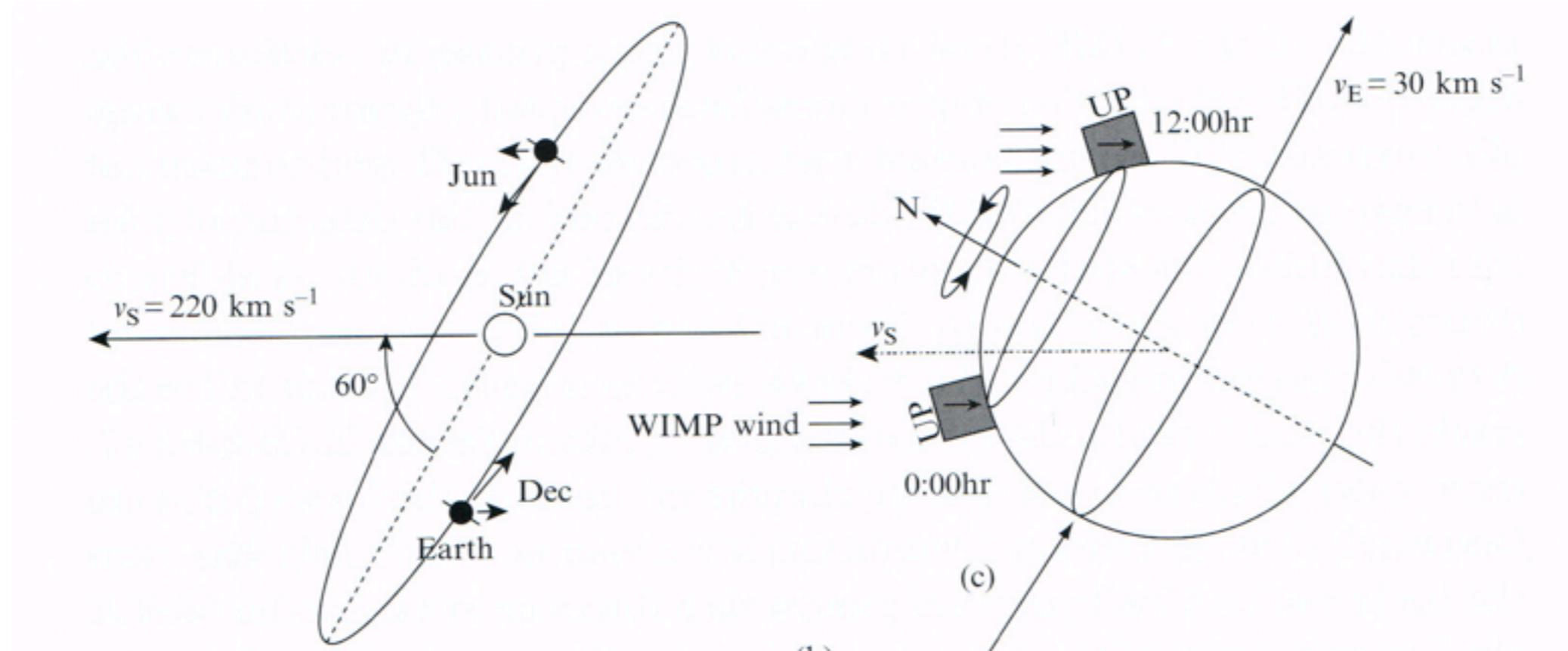


- 1460 keV gamma from decay of ^{40}K can deposit 3 keV of energy in a crystal (arXiv: 1210.5501)
- Currently in building phase.

The Experiments Part 3: The Search Continues

Directional Searches and Techniques
Involving Superheated Fluids

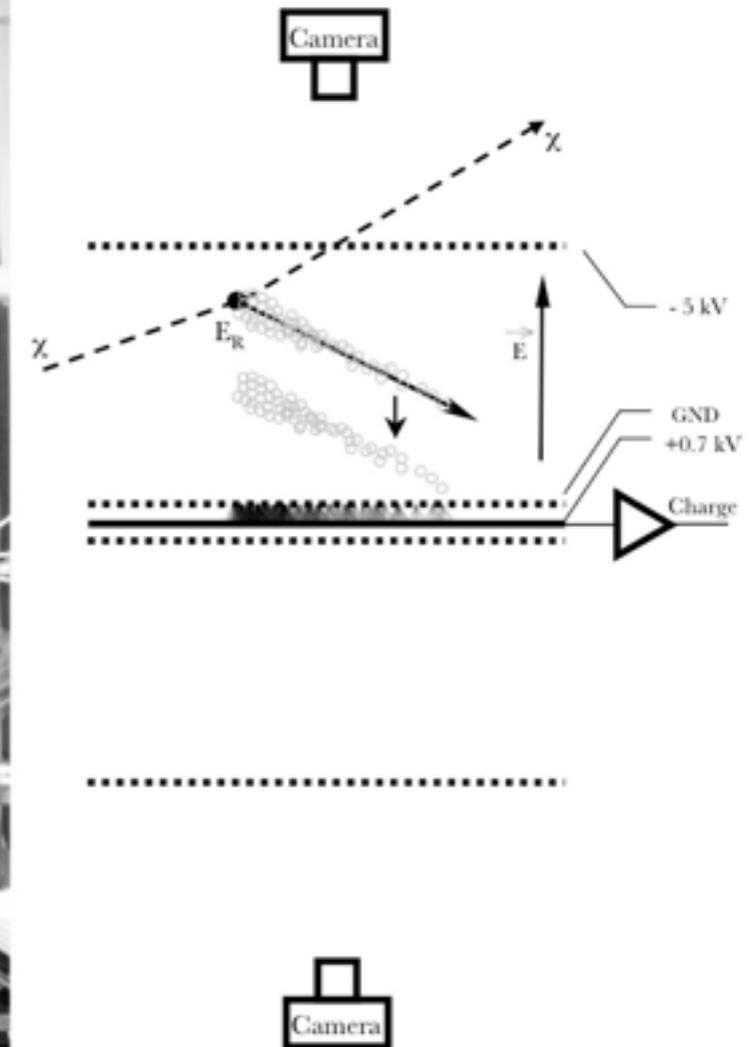
Directional Searches



- A detector at 45 degree latitude will see the dark matter wind oscillate in direction over the course of a day.
- This is a sidereal (tied to stars) effect, not diurnal (tied to sun).

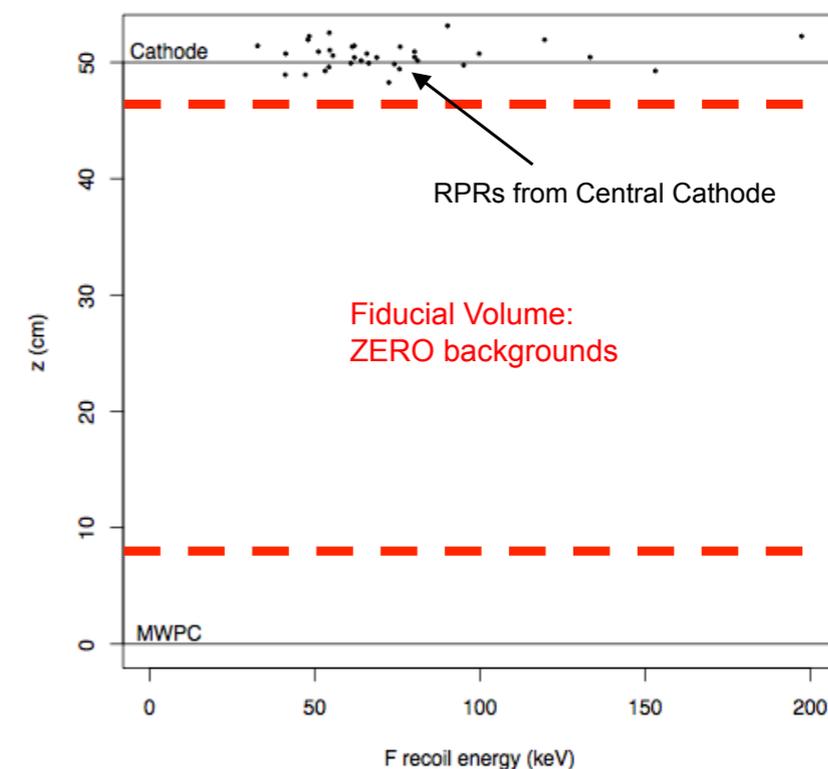
DMTPC

- 10 L prototype underground at WIPP in Carlsbad, NM, USA
- Filled with CF_4 gas to probe the WIMP- ^{19}F spin-dependent cross-section
- Dark matter is identified by directional signal.
- In addition, electron recoils can be identified by their low ionization density (i.e. stopping power).
- “4Shooter” 20 L experiment at MIT
- “DMTPCino” 1 m^3 funded by the DOE and NSF



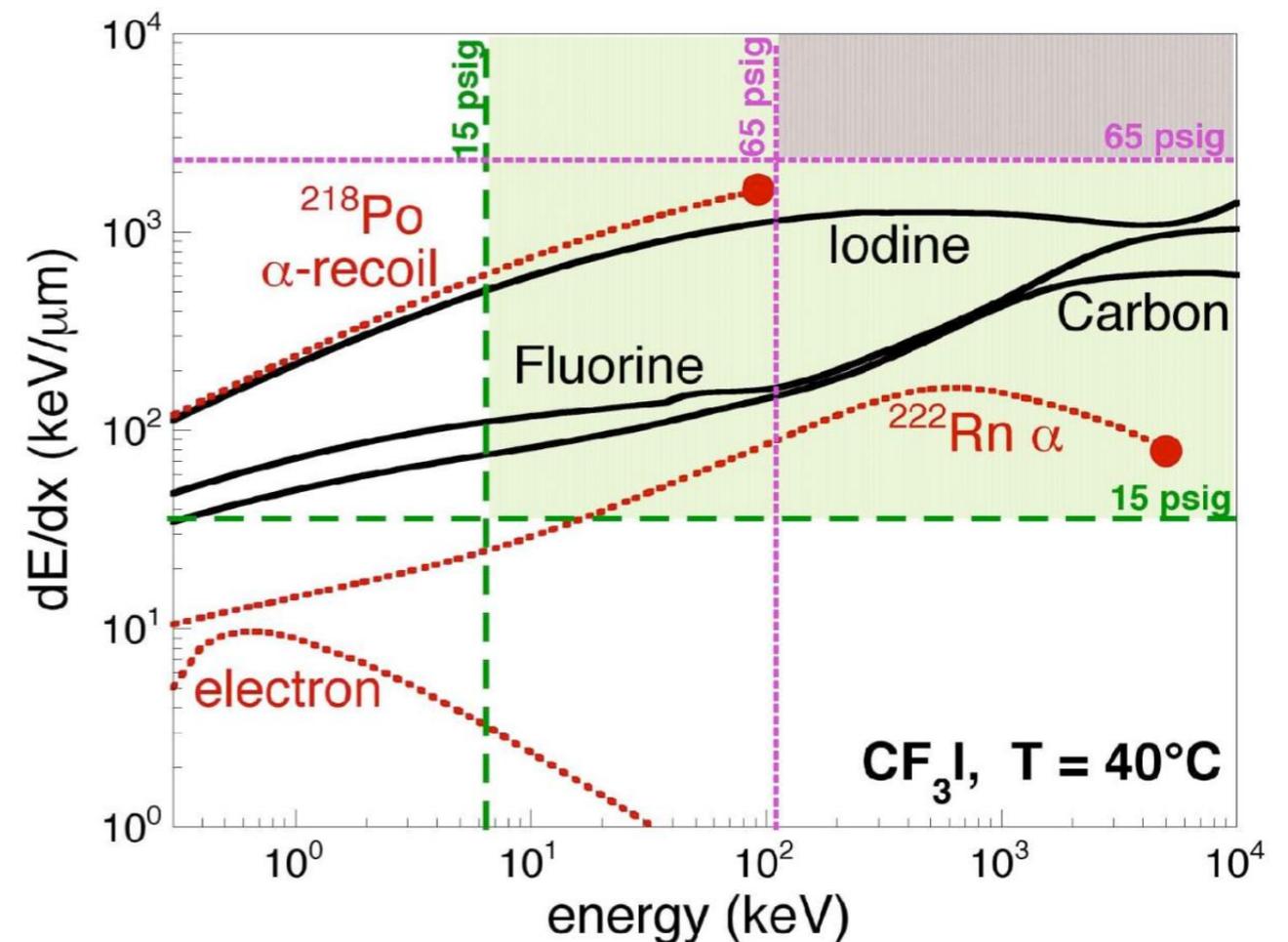
DRIFT

- DRIFT-IIId (800 L fiducial volume) operating in Boubly, UK.
 - Negative ion TPC with MWPC wire readout. with a 30 Torr CS_2 + 10 Torr CF_4 target.
 - DRIFT is now running with “zero backgrounds” and with sensitivity 10x lower than previously.
- A completely redesigned DRIFT-IIe detector will be deployed in early 2014.
 - Enhancements include “transparent”, stackable MWPCs, enhanced resolution, every-wire readout and redesigned field cage.
- Plans to submit proposal for an expandable 24 m³ experiment Fall 2013.



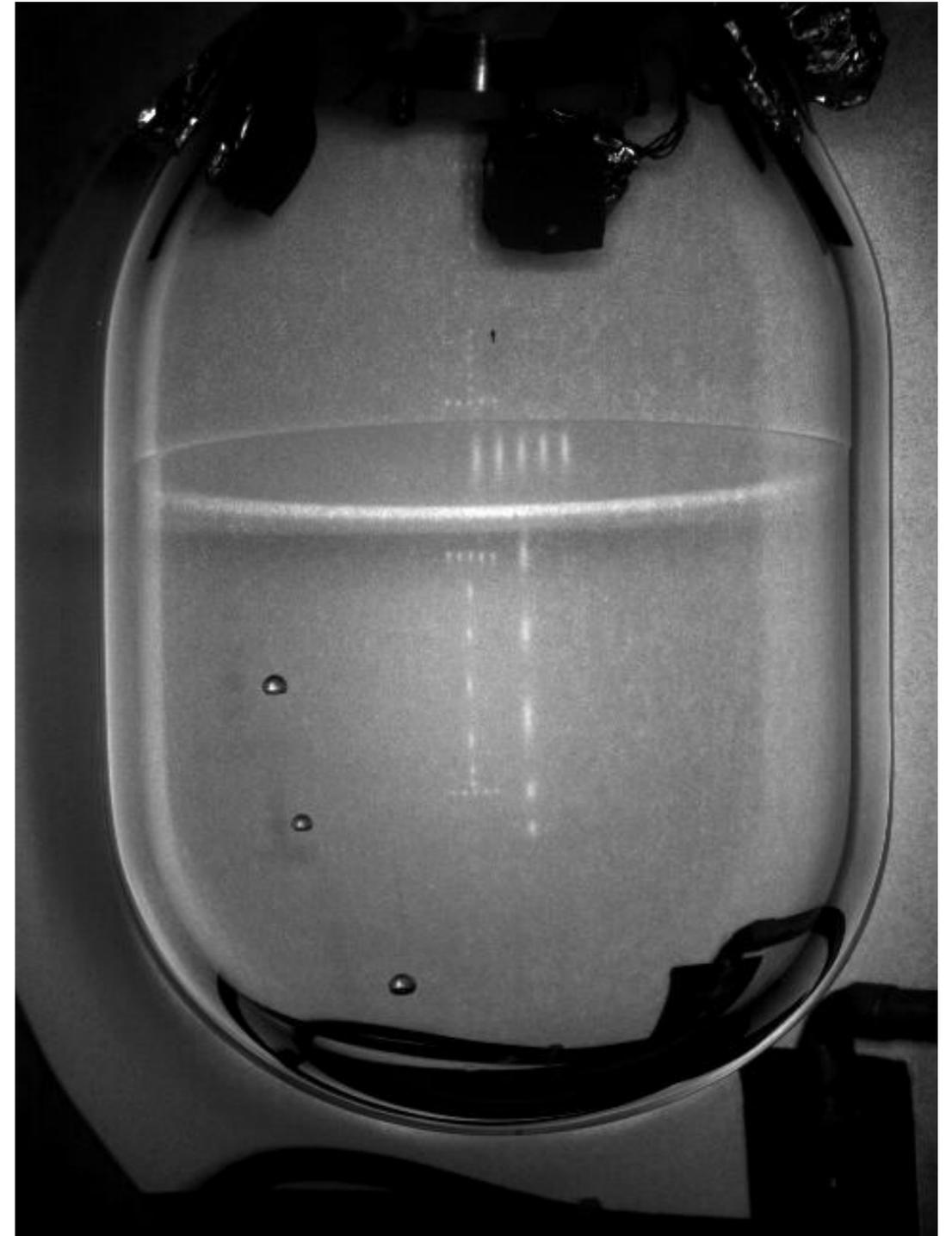
Particle Detection in Bubble Chambers

- A bubble chamber is filled with a superheated fluid in a metastable state.
- A particle interaction with energy deposition greater than E_{th} in a radius $< r_c$ results in an expanding bubble.
- A smaller or more diffuse energy deposition will result in a bubble that immediately collapses.
- You can “tune” the chamber to make bubbles for nuclear recoils and not for electron interactions.



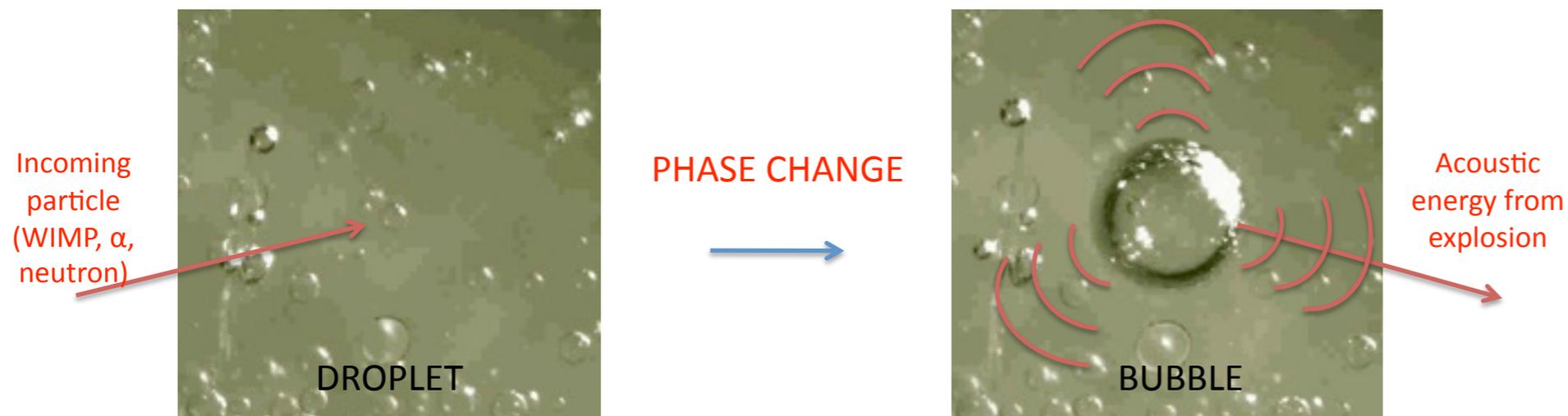
COUPP

- Superheated fluid CF_3I located in SNOLAB, Canada.
 - F for spin-dependent interactions
 - I for spin-independent interactions
 - Target can be swapped out
- Bubbles are observed by two cameras and piezo-acoustic sensors
- Better than 10^{-10} rejection of electron recoils
- Alphas can be a concern. However, they can be rejected by acoustic discrimination.



PICASSO

- A superheated liquid detector using a C_4F_{10} target.
- Location: SNOLAB, Canada
- C_4F_{10} droplets are suspended in polymerized gel in a 4.5L acrylic vessel. Experiment contains 32 modular detectors.
- Acoustic energy deposition of incoming particles is measured by 9 piezoelectric sensors.
- New results are in preparation.
- Total exposure will be ~ 800 kg-d by end of 2013.



COUPP/PICASSO/PICO

COUPP-60

- Filled with 37 kg of CF_3I on April 26, 2013
- First bubble May 1, 2013 (radon decay)
- Installation completed May 31, 2013
- Started first physics run in late June
- Increase target mass to 75 kg in fall/winter
- Ultimate goal of 3 year run (50000 kg-days exposure)

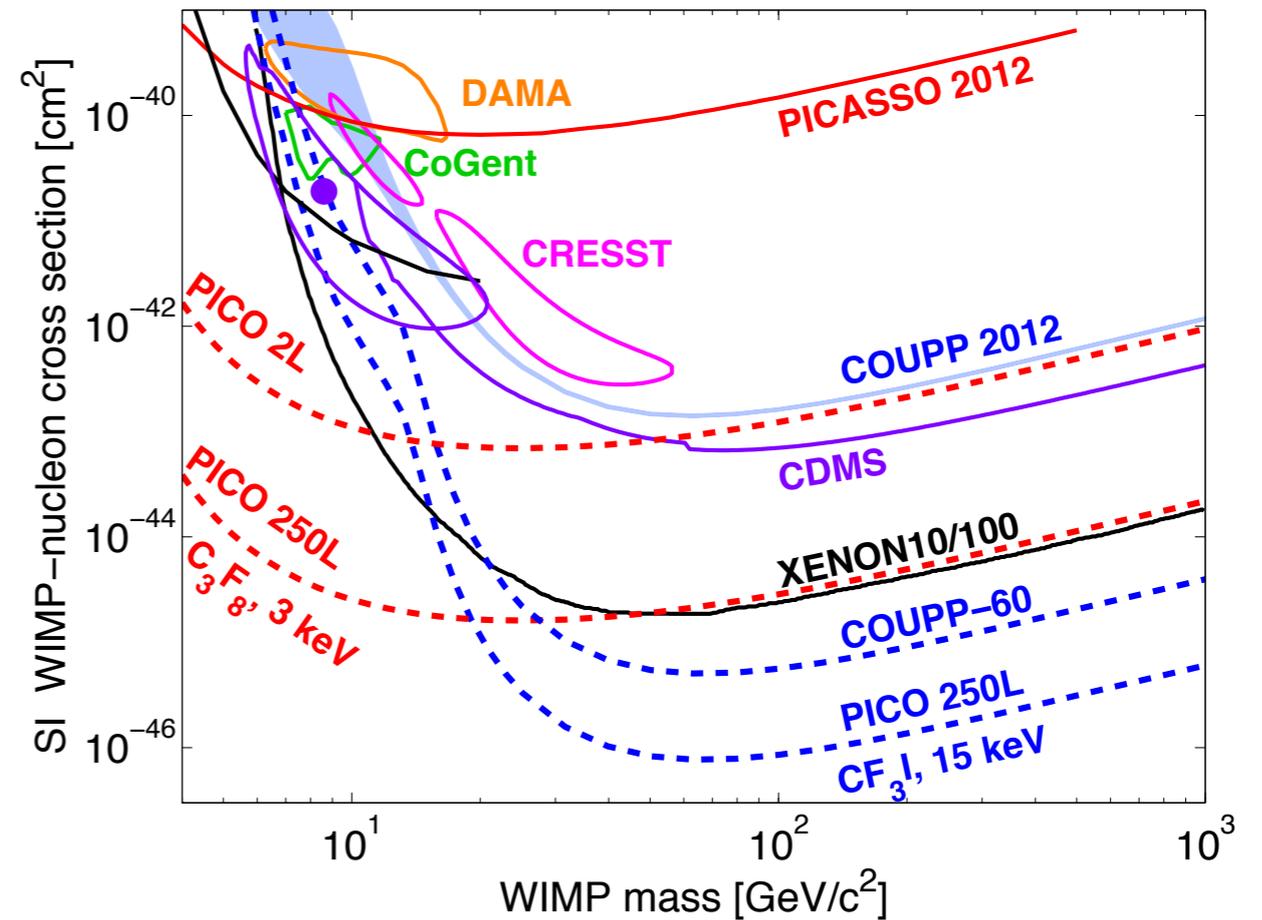
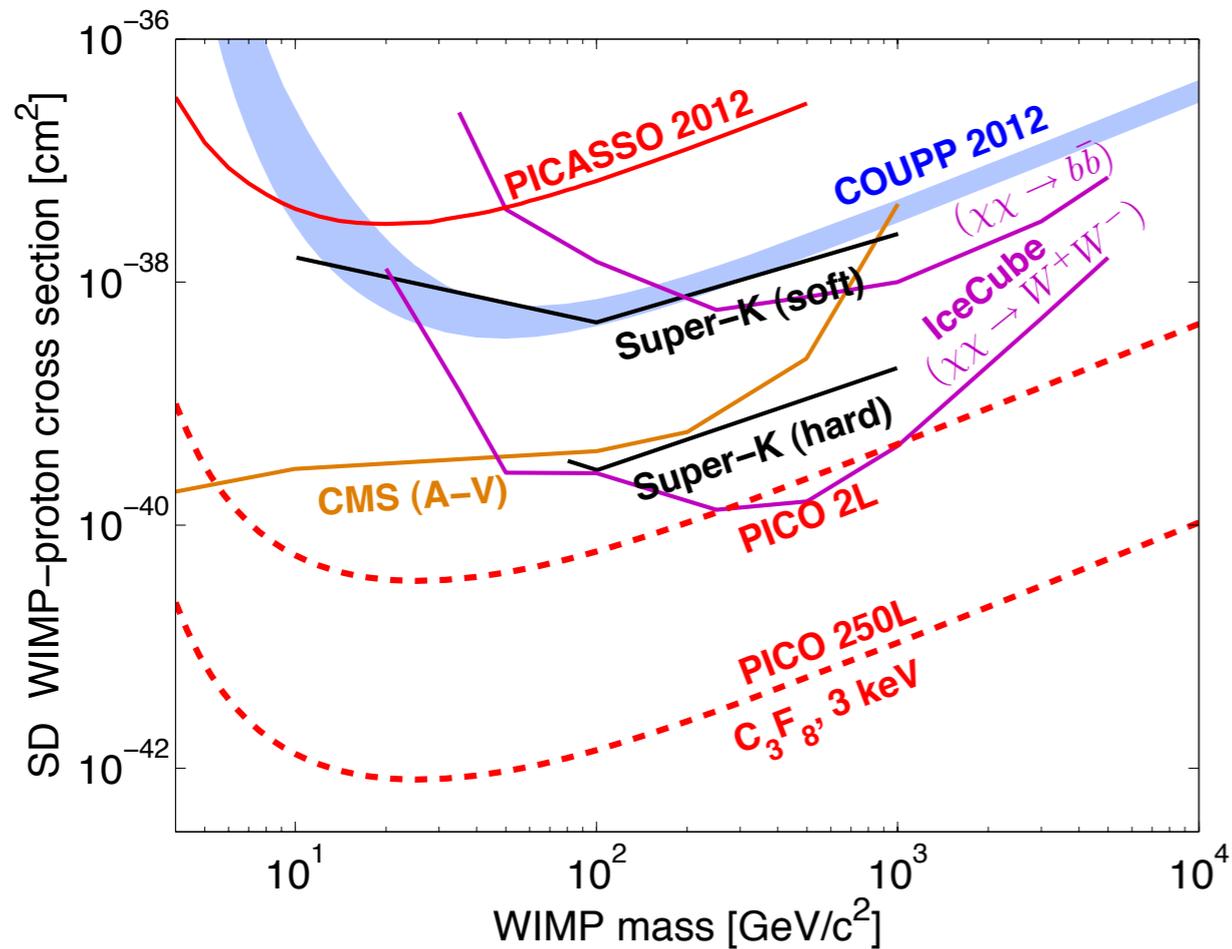
PICO-lite

- Joint effort between COUPP & PICASSO
- C_3F_8 chamber (2L) in existing COUPP-4 infrastructure at SNOLAB
- 3 keV threshold
- Excellent low-mass WIMP and SD coupling sensitivity
- CDMS-Si result gives 1 event/day in COUPP-4lite
- Deploy September 2013

PICO-250

- 250L bubble chamber design effort
- Well developed Conceptual Design
- Straightforward scale-up from COUPP-4 and COUPP-60
- Begin construction in 2014-2015

COUPP/PICASSO/ PICO Projections and Limits



- COUPP Limits: **Phys. Rev. D 86, 052001 (2012)**
 - 553 kg-days total exposure (4.0-kg CF₃I)
- PICASSO Limits: **Phys Lett B 711(2) (153-161)**
 - 114 kg-d exposure (10 modules, 0.72 kg of ¹⁹F).

Other Experiments

CUORE

Biassoni - Mon. DM II

Gaseous Detectors

Gerbier - Mon. DM II

NEXT

Renner - Wed. DM V

DAMIC

Chavarria - DM VI

TEXANO

Lin - Thurs. DM VIII

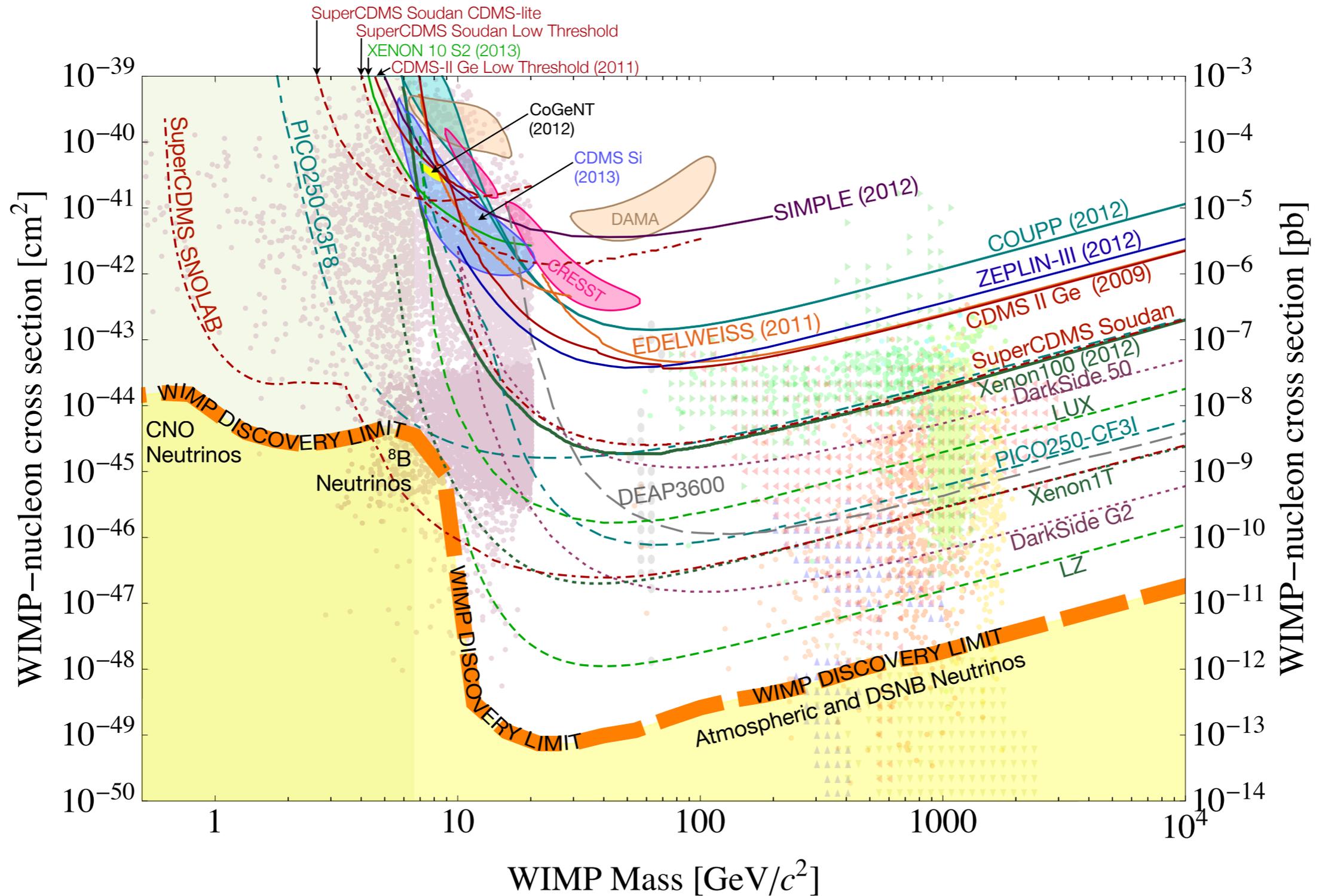
CDEX

Yue - Thurs. DM VIII

Kamland-PICO

Fushimi - Thurs. DM VIII

Where Are We Going?



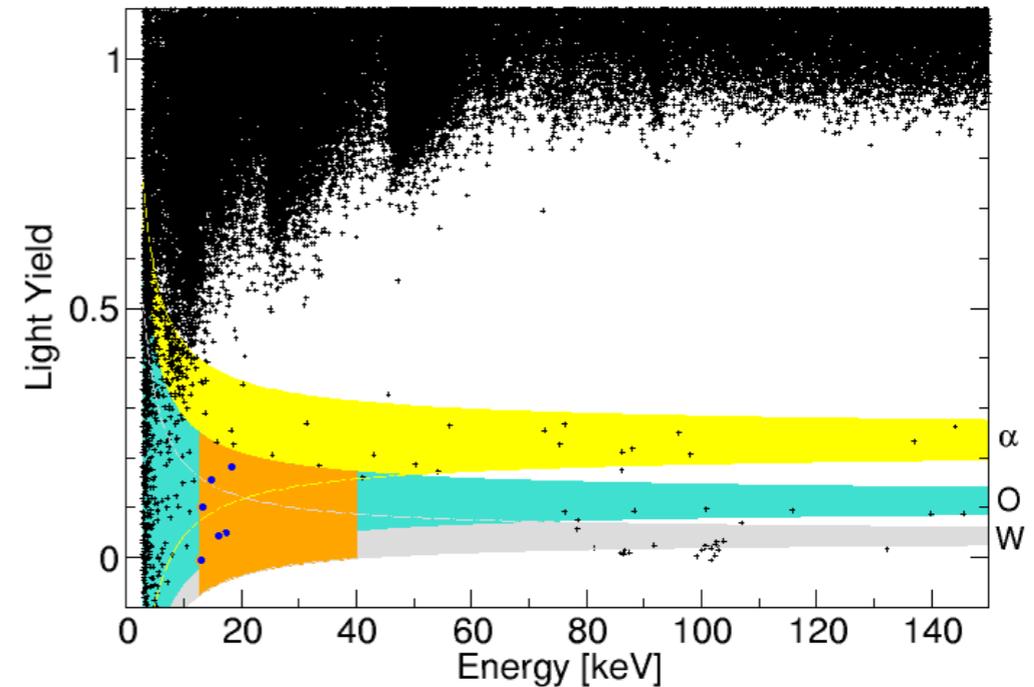
Summary/Outlook

- Dark matter experimentalists have come up with clever techniques to suppress backgrounds in an attempt to extract a dark matter signal.
- Four experiments have observed excess events. If these events are interpreted as dark matter CDMS Si and CoGeNT are compatible. However, it is difficult to reconcile their results from CRESST and DAMA.
- It is necessary to have several technologies in different locations.
- There are many experiments using different techniques currently running world wide. The techniques employed include solid-state devices, two-phase and single-phase noble liquid detectors, superheated detectors.
- There are many planned upgrades and extensions to existing experiments to achieve greater sensitivity.
- It is an interesting time to be working in this field!

Backup Slides

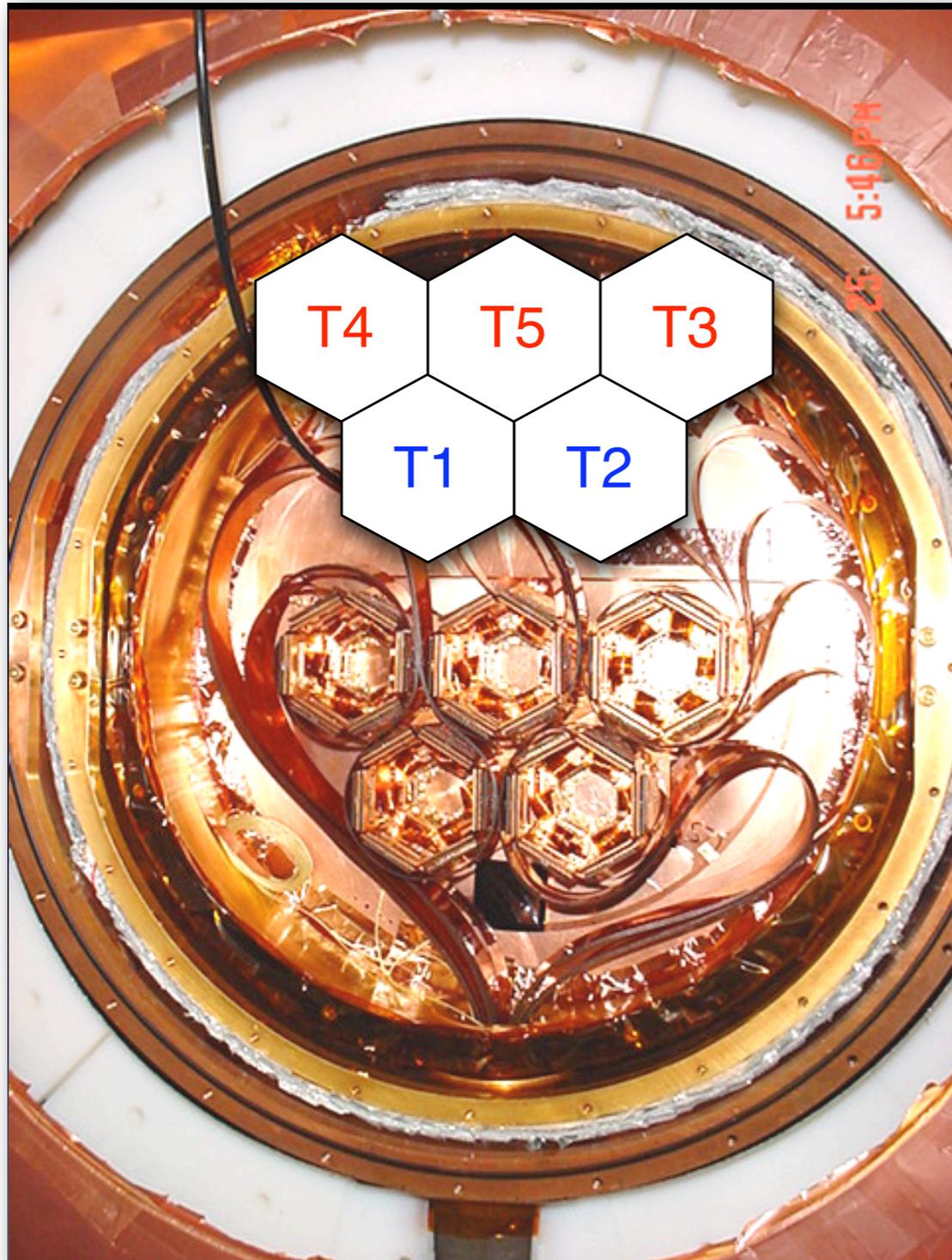
CRESST-II Data

- Net exposure: 730 kg-day (July 2009 - March 2011) from 8 detector modules.
- Observed 67 events in acceptance region (orange).
- Analysis used a maximum likelihood in which 2 regions favored a WIMP signal in addition to predict background.
 - M1 is global best fit (4.7σ)
 - M2 slightly disfavored (4.2σ)
- Excess events can not be explained by known backgrounds
- Large background contribution



	M1	M2
e/γ events	8.00 ± 0.05	8.00 ± 0.05
α events	$11.5^{+2.6}_{-2.3}$	$11.2^{+2.5}_{-2.3}$
neutron events	$7.5^{+6.3}_{-5.5}$	$9.7^{+6.1}_{-5.1}$
Pb recoils	$15.0^{+5.2}_{-5.1}$	$18.7^{+4.9}_{-4.7}$
signal events	$29.4^{+8.6}_{-7.7}$	$24.2^{+8.1}_{-7.2}$
m_χ [GeV]	25.3	11.6
σ_{WN} [pb]	$1.6 \cdot 10^{-6}$	$3.7 \cdot 10^{-5}$

CDMS II - Si Analysis



- 30 detectors installed and operated in Soudan from June 2006 - March 2009.

- ~4.75 kg Ge, ~1.1 kg Si

- Seven Total Data Runs

- R123- R124 (Oct. 2006 - July 2007)

- 55.9 kg-days in 6 Si detectors

- R125 - R128 (July 2007 - Sep. 2008)

- 140.23 kg-days in 8 Si detectors

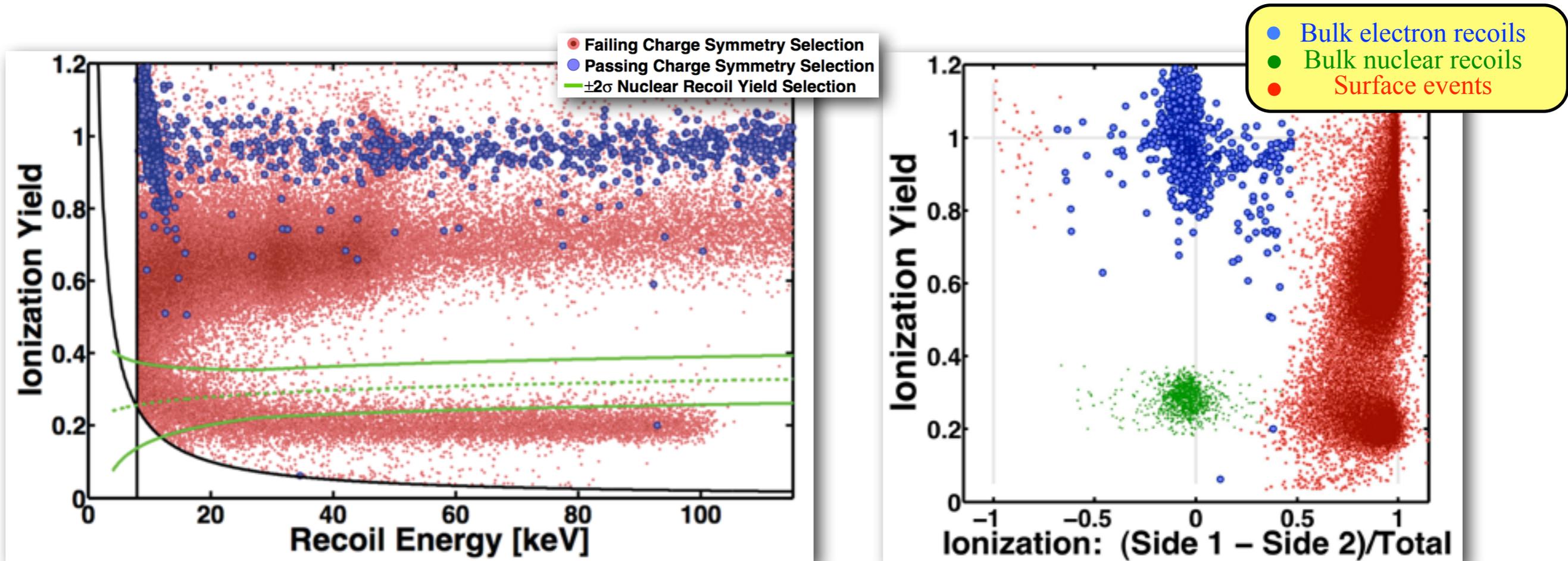
- R129 (Nov. 2008 - Mar. 2009)

	T1	T2	T3	T4	T5
Z1	G6	S14	S17	S12	G7
Z2	G11	S28	G25	G37	G36
Z3	G8	G13	S30	S10	S29
Z4	S3	S25	G33	G35	G26
Z5	G9	G31	G32	G34	G39
Z6	S1	S26	G29	G38	G24

Side View

SCDMs: ^{210}Pb Test

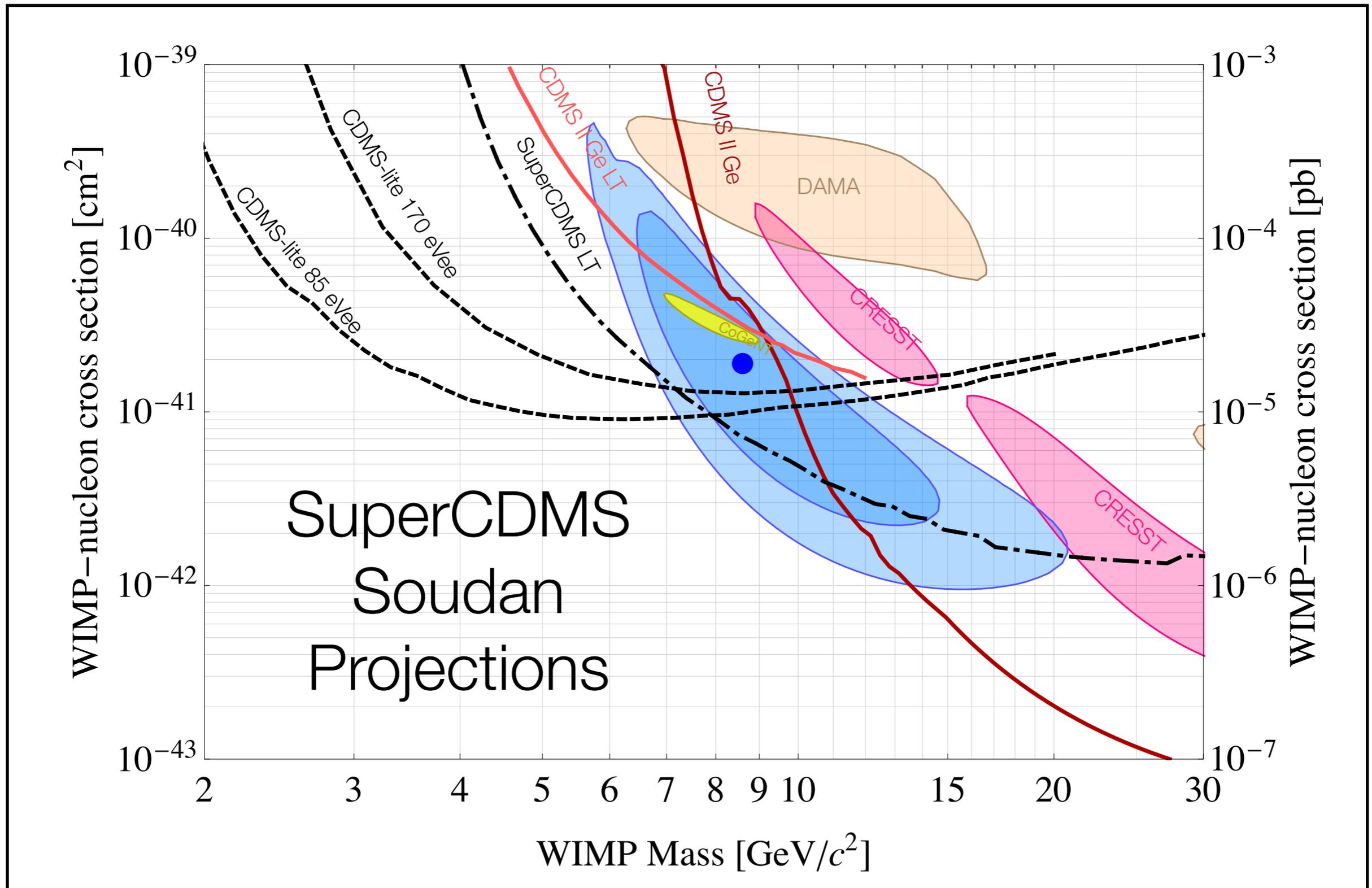
Two ^{210}Pb sources were deployed with the detectors to test surface rejection capabilities of the new iZIP detectors.



- 71,525 (38,178) electrons and 16,258 (7,007) ^{206}Pb recoil surface event collected from ^{210}Pb source in 905.5 (683.8) live hours
- In ~800 live hours 0 events leaking into the signal region ($< 1.7 \times 10^{-5}$ @90% C.L. misID)

- ~50% fiducial volume (8-115 keVr)
- < 0.6 events in 0.3 ton-years
- Good enough for a 200 kg experiment run for 4 years at SNOLAB!

SuperCDMS @ Soudan: Low Mass Projections

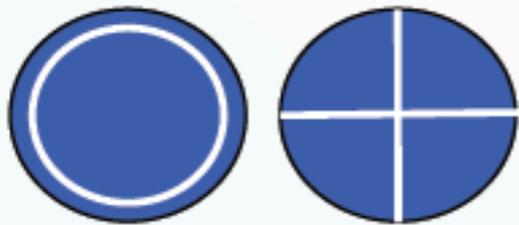


SNOLAB Detectors

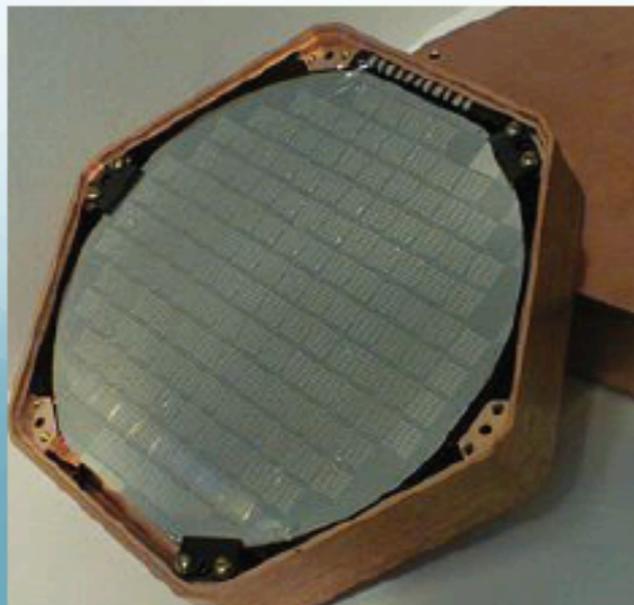
CDMS II

Single-sided
1 cm thick
3" diameter
250 g Ge

2 charge + 4 phonon



5 towers of 6 det each



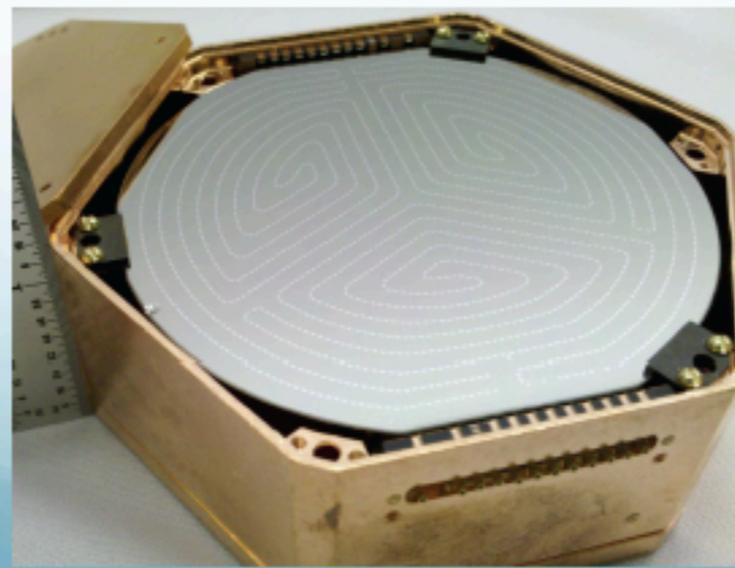
SuperCDMS Soudan

Double-sided
2.5 cm thick
3" diameter
620 g Ge

2 charge + 2 charge
4 phonon + 4 phonon



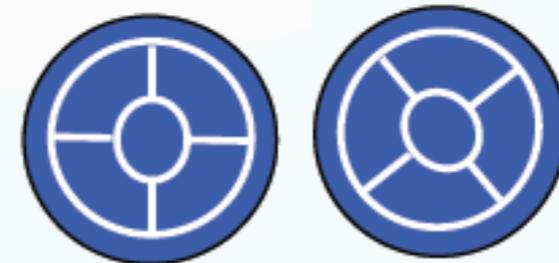
5 towers of 3 det each



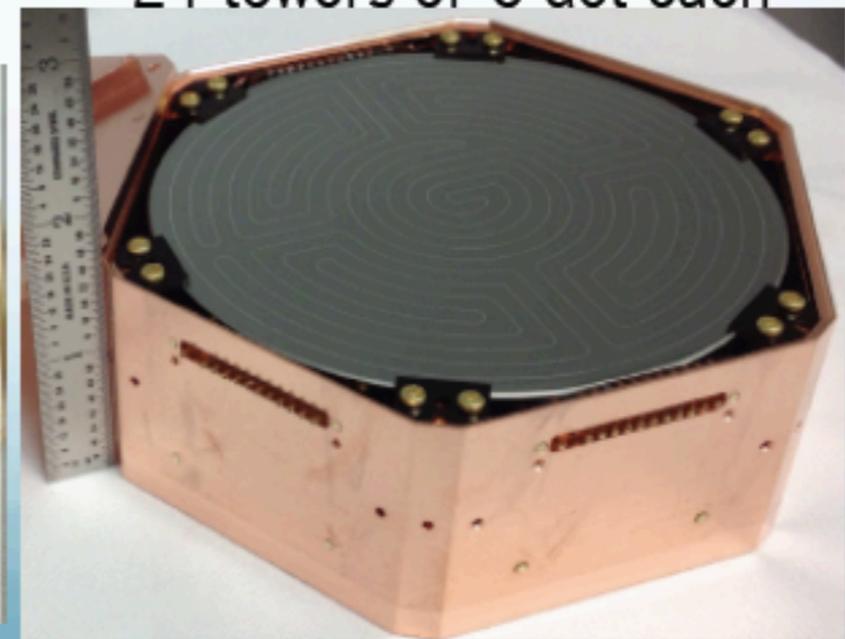
SuperCDMS SNOLAB

Double-sided
3.3 cm thick
4" diameter
1.38 kg Ge

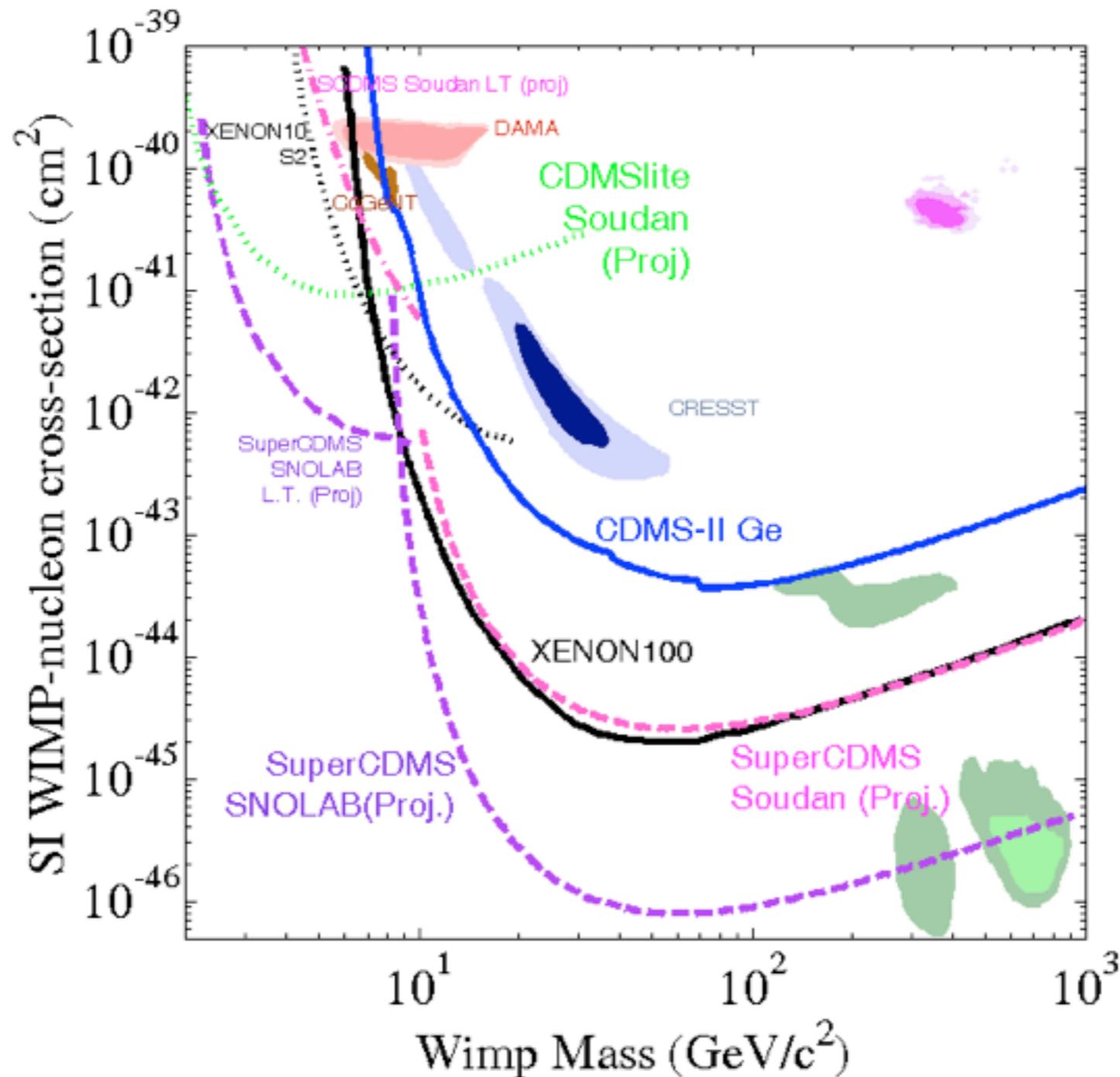
2 charge + 2 charge
6 phonon + 6 phonon



24 towers of 6 det each



SNOLAB Projections

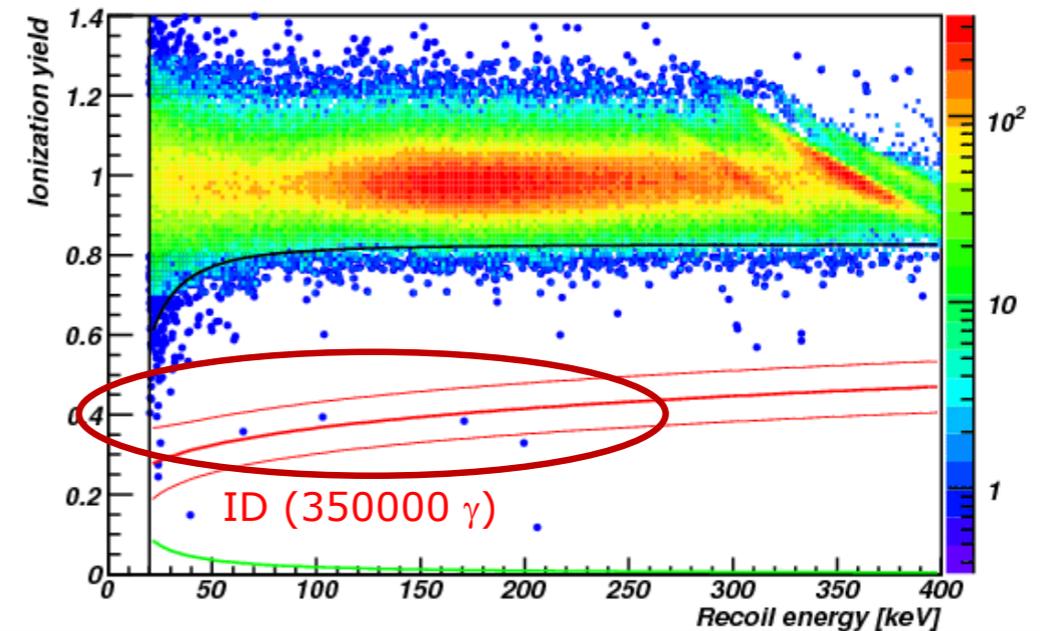
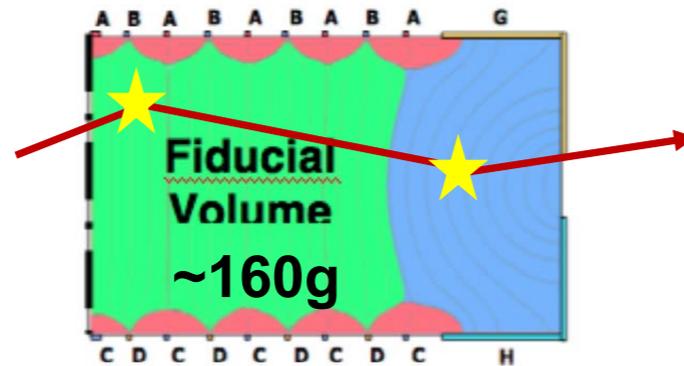
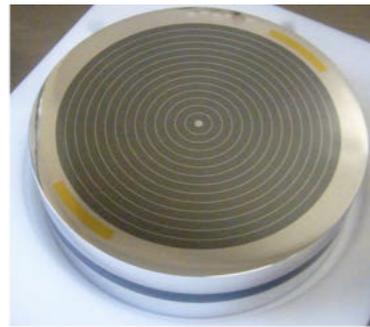


EDELWEISS

2. Improvement of γ discrimination

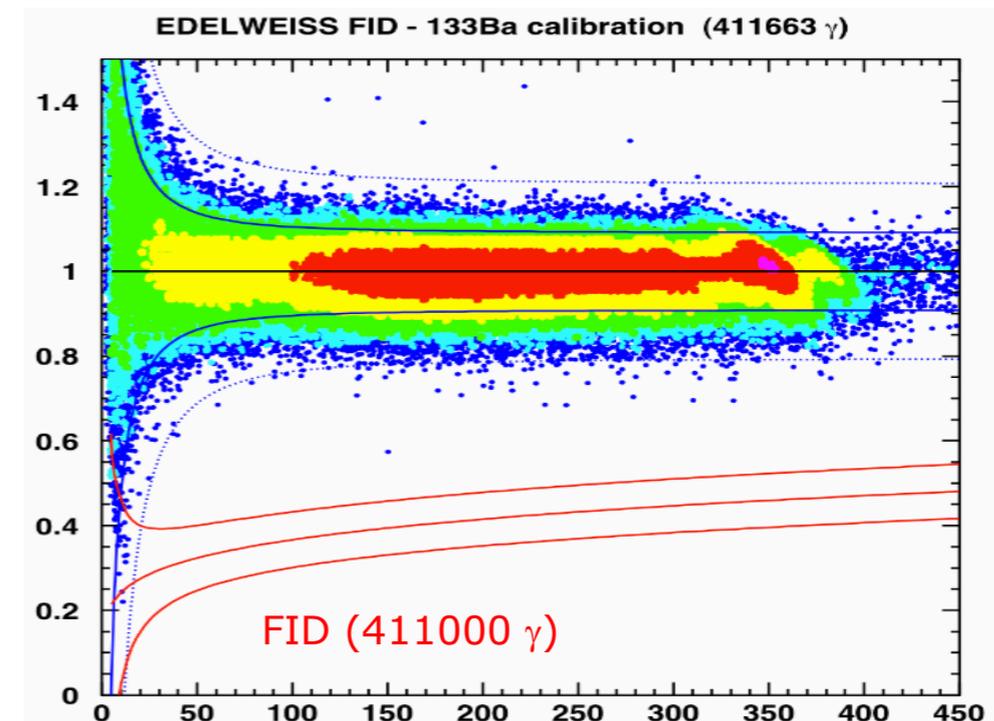
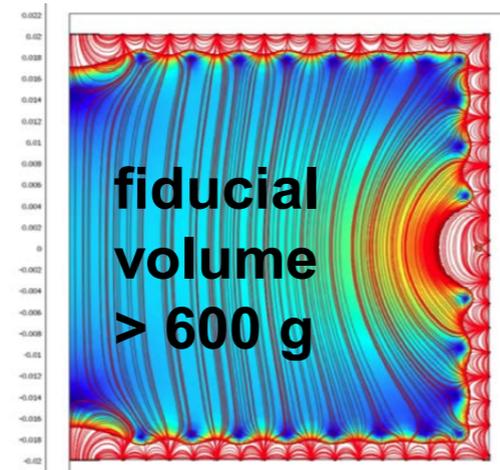
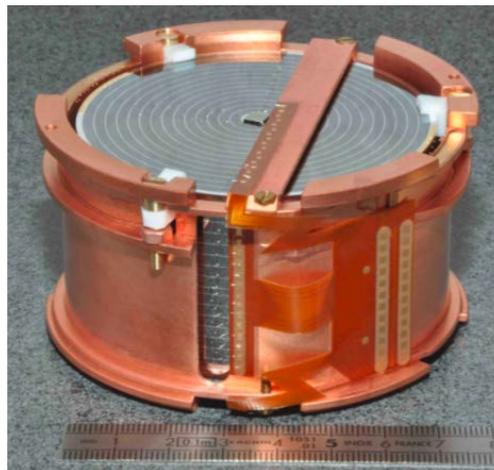
EDELWEISS-II

ID 400g with 10x 160g fiducial mass



EDELWEISS-III

FID 800g with 40x ~600g fiducial mass



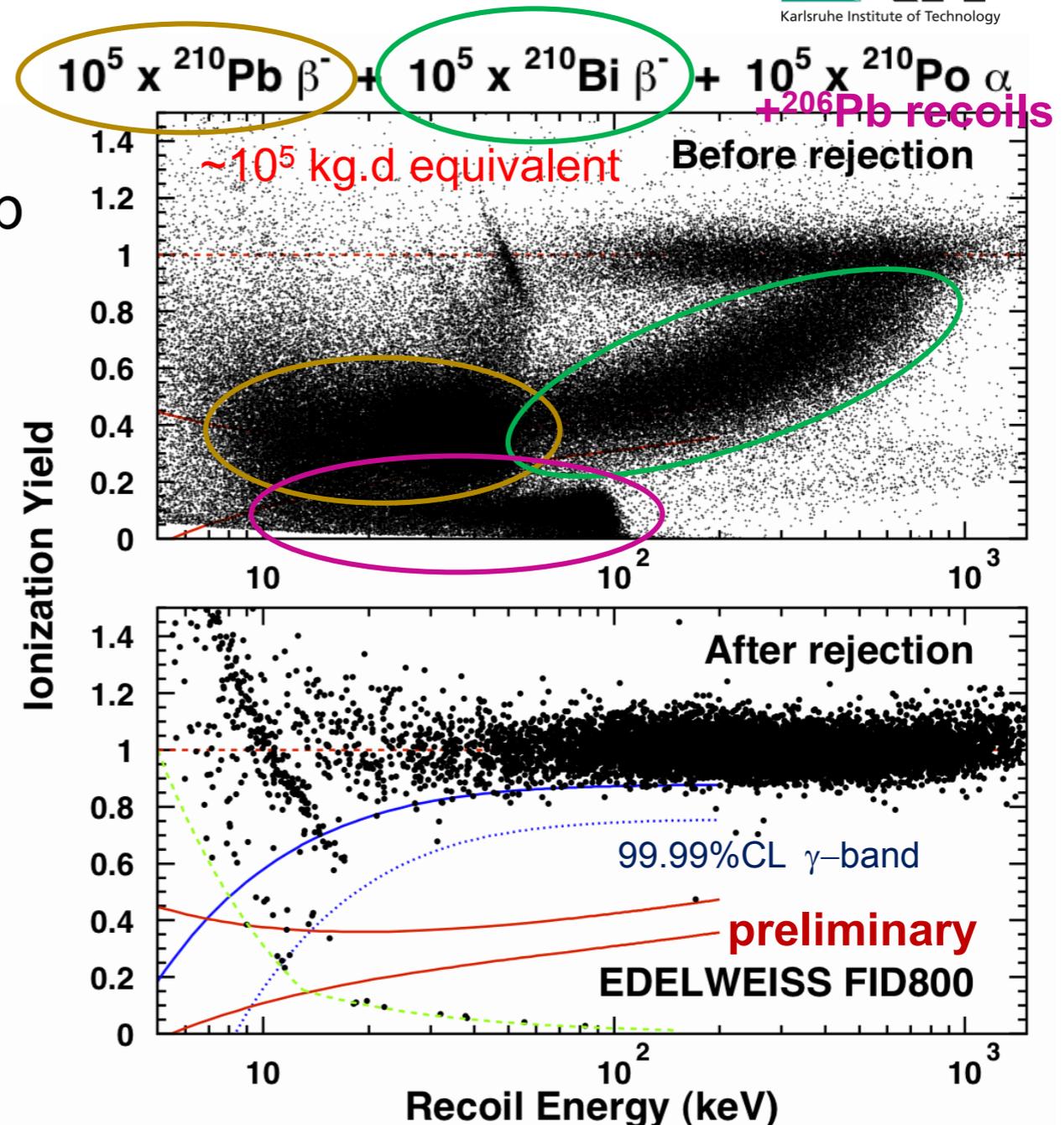
EDELWEISS

3. Surface rejection measurements – improved discrimination



- measurement with ^{210}Pb β^- -source
- surface rejection:
 - < 4×10^{-5} misidentified events per kg.d ($E_{\text{rec}} > 15$ keV)

better than previous EDELWEISS detectors
 (< 6×10^{-5} misidentified events per kg.d, $E_{\text{rec}} > 20$ keV)



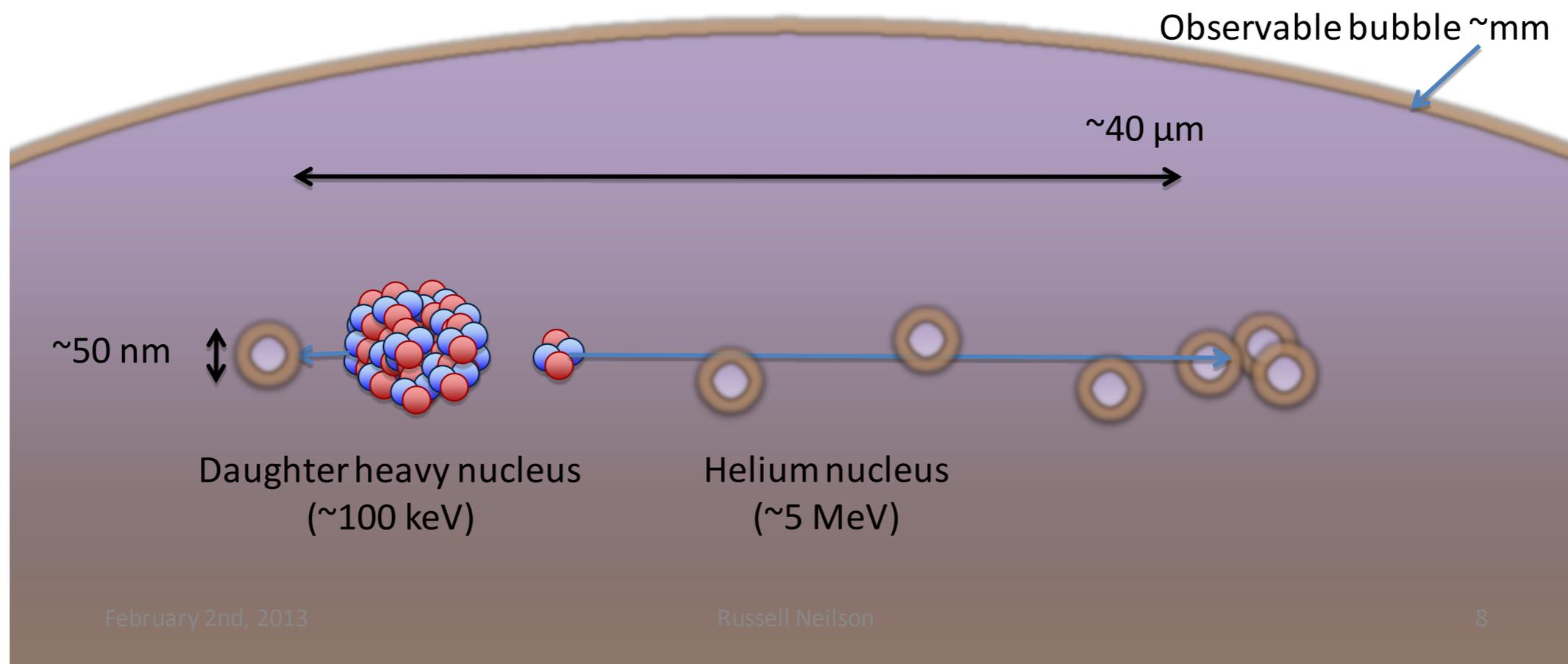
CUORE

- Detection mechanism: coherent scattering of WIMP on detector nuclei
- Spin-dependent cross section highly suppressed, almost only spin-independent interactions
- Both heavy (Te) and light (O) targets in the same detector
- Scattered nucleus recoils in the crystal lattice, energy converted into heat
 - PROS:
 - ✓ large mass, long term stability ==> seasonal modulation of events number
 - ✓ high energy resolution ==> seasonal modulation of spectral shape
 - ✓ quenching factor = 1, all energy detected for all recoils
 - ✓ detection efficiency = 1, basically all recoils are fully contained
 - CONS:
 - ✓ recoil energy is typically small (depending on nucleus and WIMP mass) ==> LOW THRESHOLD REQUIRED (<25keV)
 - ✓ no particle identification (nuclear recoils, gammas, betas, alphas all the same)

Great effort in lowering the threshold and reducing the background

COUPP

- Alphas deposit their energy over 10s of microns.
- Nuclear recoils deposit their energy over 10s of millimeters
- Alpha particles are louder than nuclear recoils. This can be measured by piezoelectric sensors.



DMTPC

10L



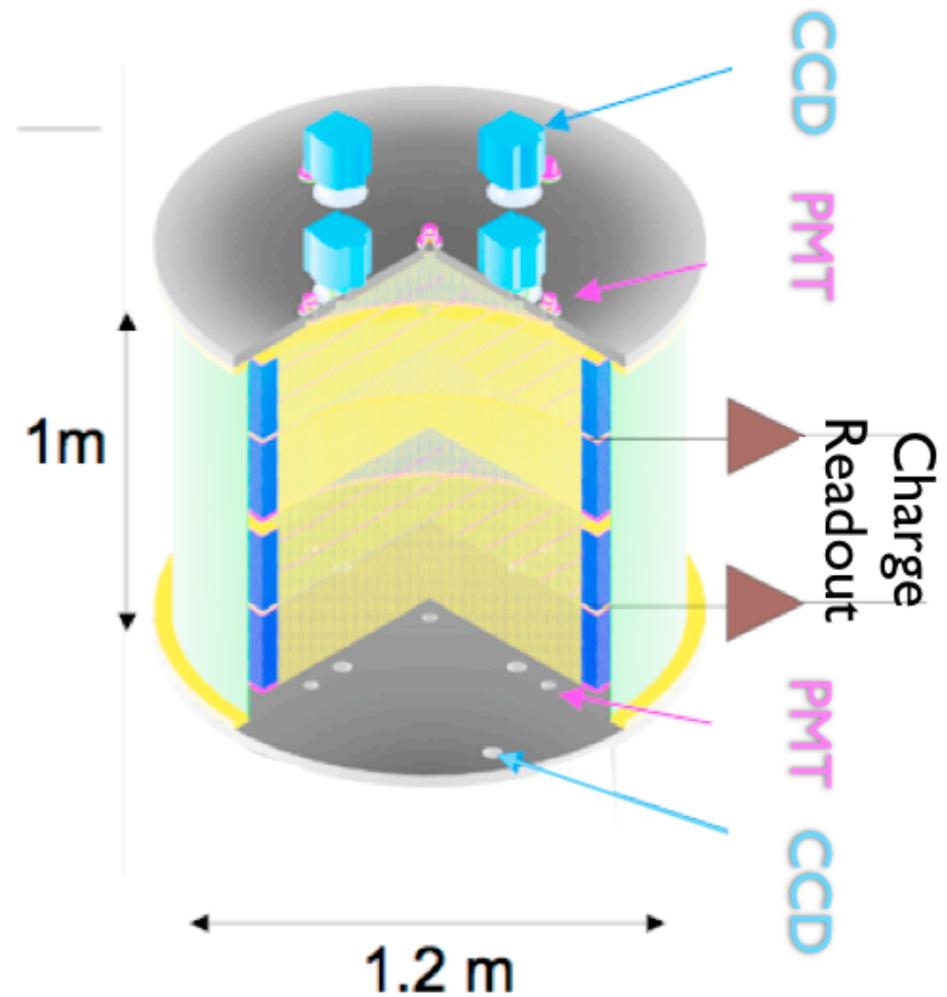
Underground
at WIPP

4Shooter (20L)



At MIT

DMTPCino (1 m³)

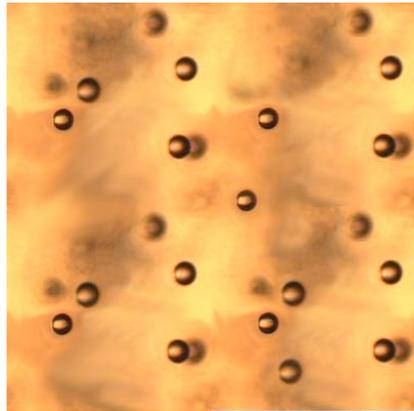


Funded by
NSF+DoE

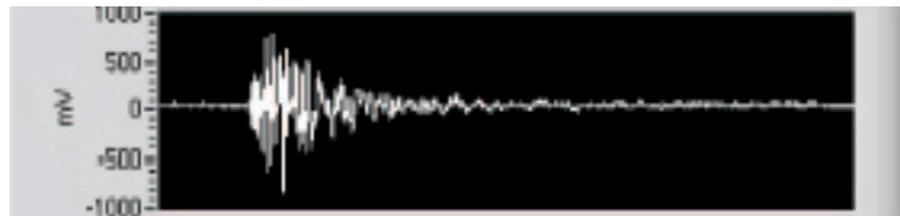
James Battat Bryn Mawr College

PICASSO

- Modular detector (32 modules).
- Uses C_4F_{10} droplets (~200 μm diameter)...
- ...suspended in polymerised aqueous gel matrix...



- ...in 4.5L acrylic cylindrical container.
- 9 piezoelectric transducers record sound.



- 40-50 hr data taking runs.
- 2-5 hr calibration runs with neutron source.
- 11 hr recompression between runs.

