

Other WIMP Direct Detection Experiments

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Outline

- Principles common to experiments
- The Experiments
 - Part 1: The low mass region
 - Part 2: The long standing DAMA/LIBRA experiment
 - Part 3: The search continues...
- Concluding Remarks

World-Wide Experiments

Phonon/Charge/Light:

CDMS/SuperCDMS

EDELWEISS

CRESST

Charge Only:

CoGeNT/C4

TEXANO

CDEX

CDMSlite

Multi-purpose:

Majorana Demonstrator

COURE-0/COURE

Modulation:

DAMA/LIBRA

DM-ICE

KIMS

ANAIS

SABRE

KamLAND-PICO

Bubble Chambers/Superheated:

PICASSO

COUPP

PICO

Directional:

DRIFT

DM-TPC

Other:

DAMIC

NEXT

*Experiments in red are presenting results or status in parallel sessions.

World-Wide Experiments

Phonon/Charge/Light:

CDMS/SuperCDMS

EDELWEISS

CRESO

Charge

CoGeNT

TEXA

CDEX

CDMS

Modulation:

DAMA/LIBRA

DM ICE

Directional:

DRIFT

DM-TPC

Too Many Experiments, Too Little
Time - My Apology for not
Covering All

er:

AMIC

EXT

eated:

Multi-purpose:

Majorana Demonstrator

COURE-0/COURE

PICASSO

COUPP

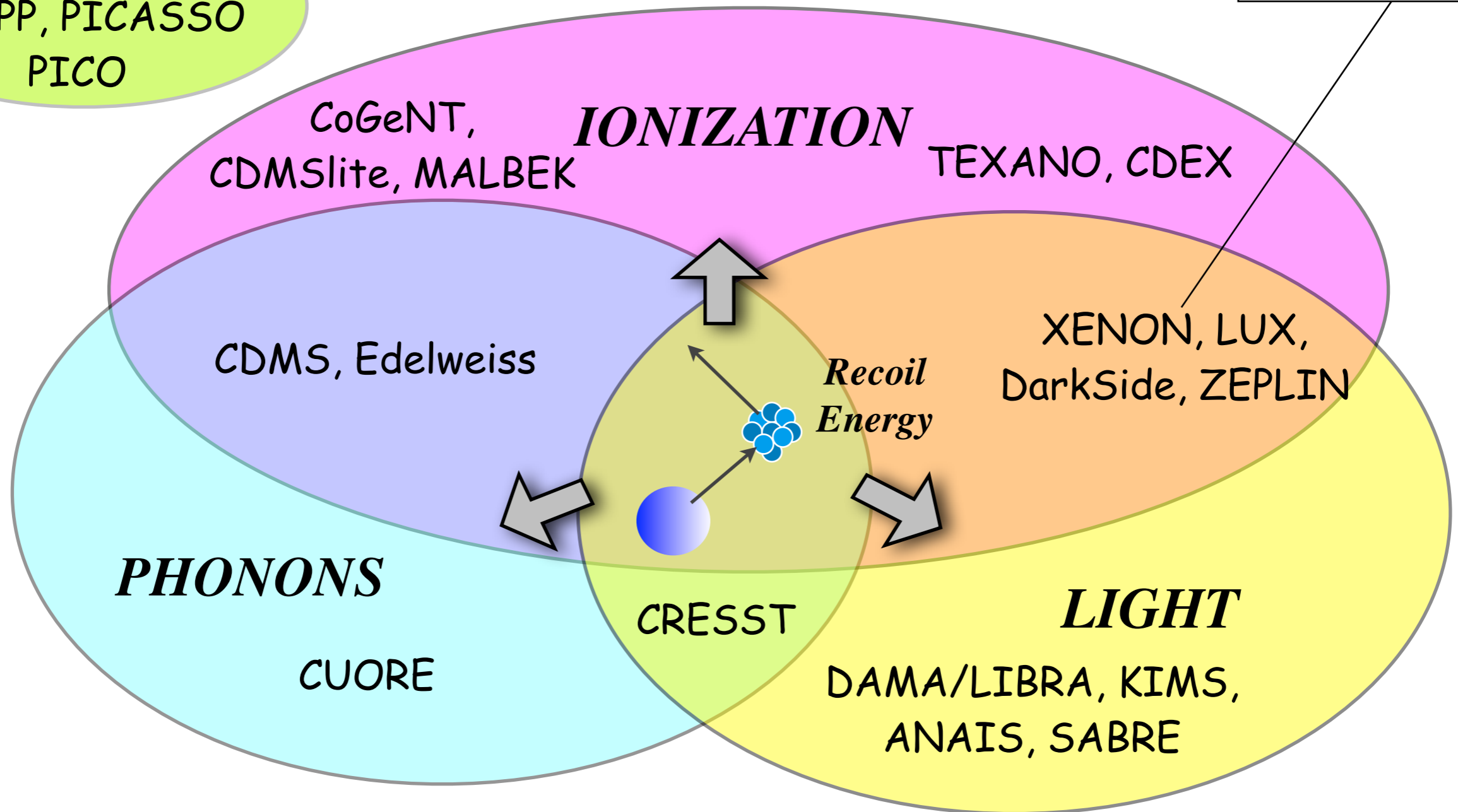
PICO-lite

*Experiments in red are presenting results or status in parallel sessions.

Direct Detection

SuperHeated
COUPP, PICASSO
PICO

Baudis - Tues.
Plenary



Minimize Backgrounds



Need at least 1000 m rock (~ 3000 mwe) overburden
to reduce muon rate by $\sim 10^5$

Nigel Smith - DM2012

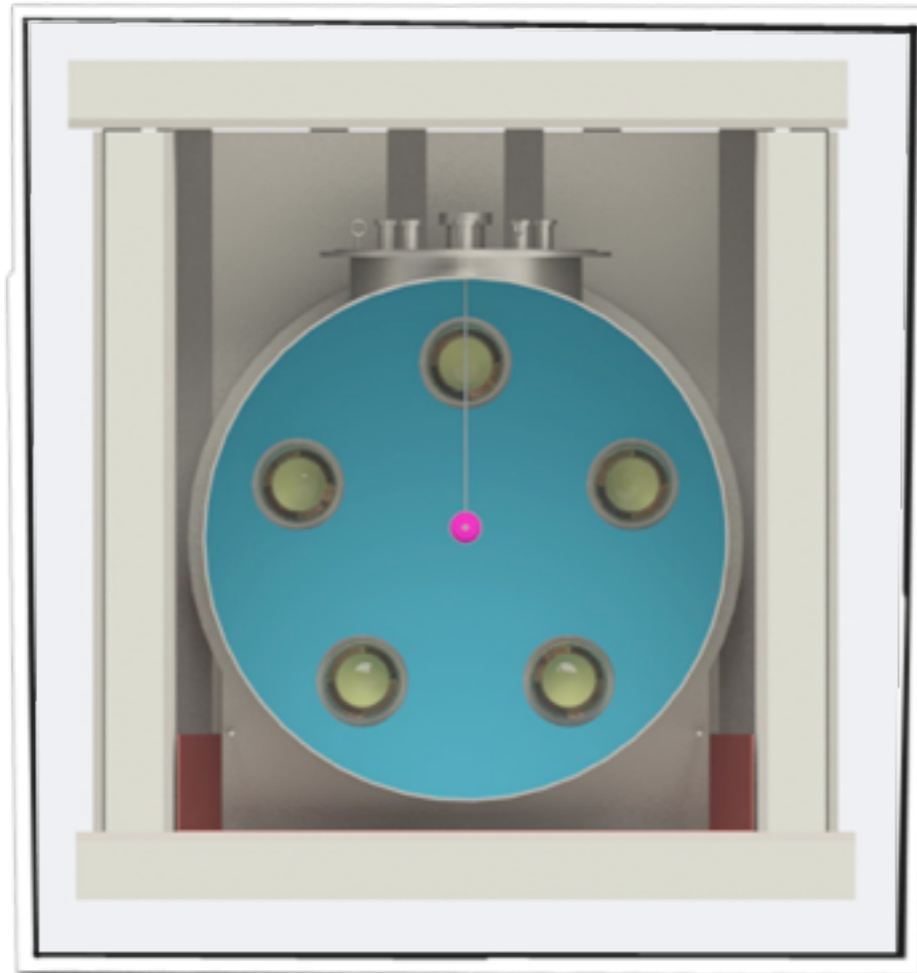
Site experiments underground.

Minimize Backgrounds

Active Muon Veto:

rejects events from cosmic rays

- Scintillating panels
- Water/Liquid Scintillator Shield



SABRE LAB shield design



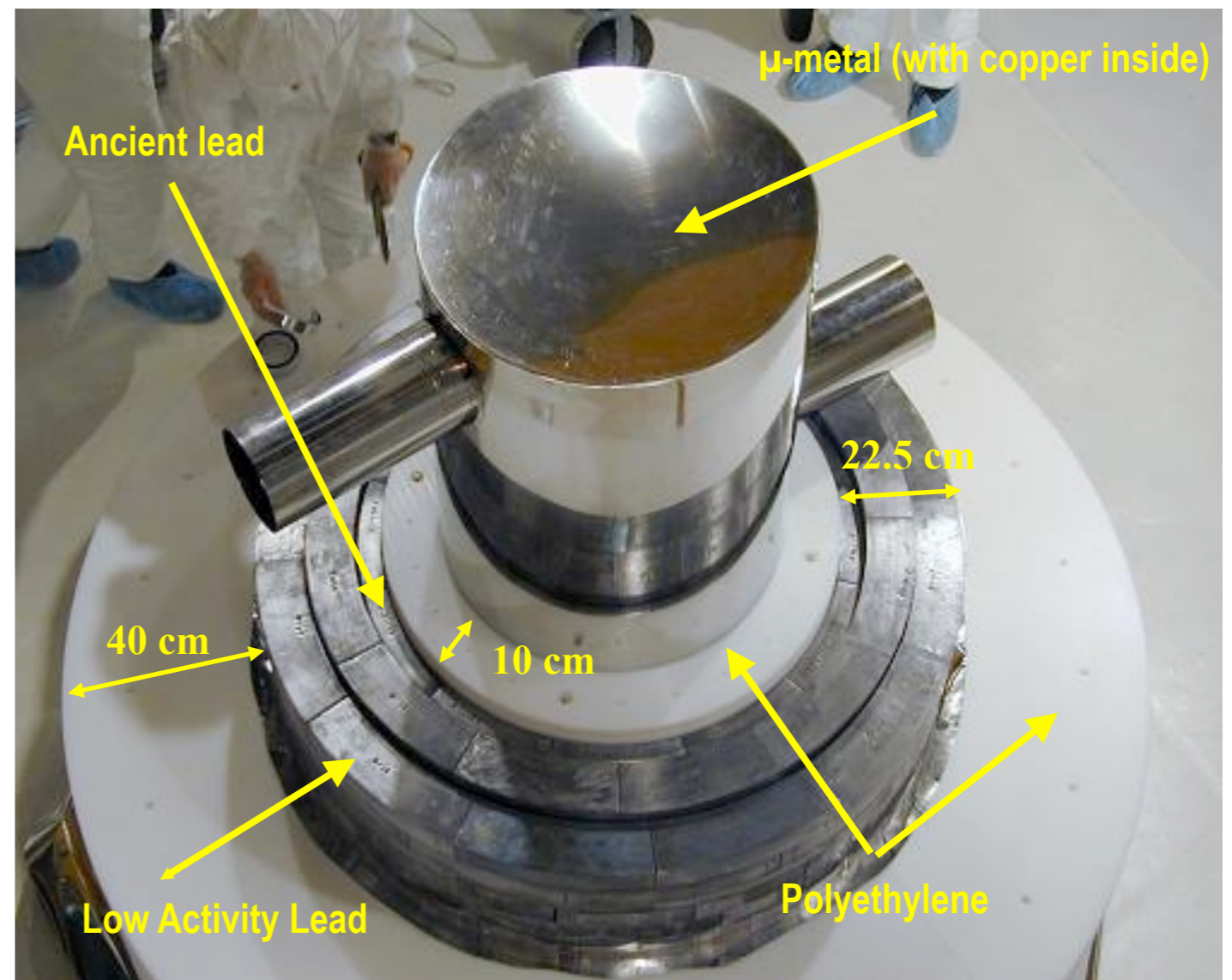
SCDMS active muon veto

Minimize Backgrounds

Use Passive Shielding

Pb: shielding from gammas resulting from radioactivity

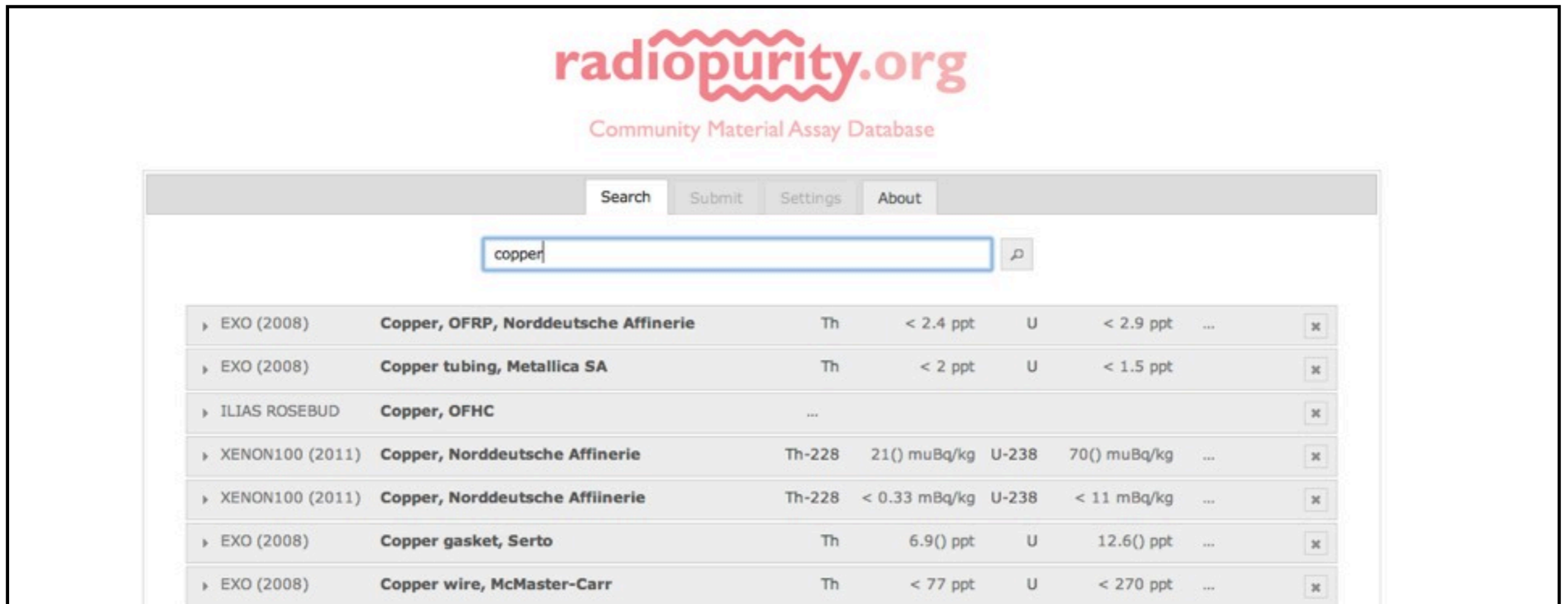
Polyethylene: moderate neutrons produced from fission decays and from (α, n) interactions resulting from U/Th decays



SCDMS - Layers of Polyethylene and Lead

Minimize Backgrounds

Use Clean Materials



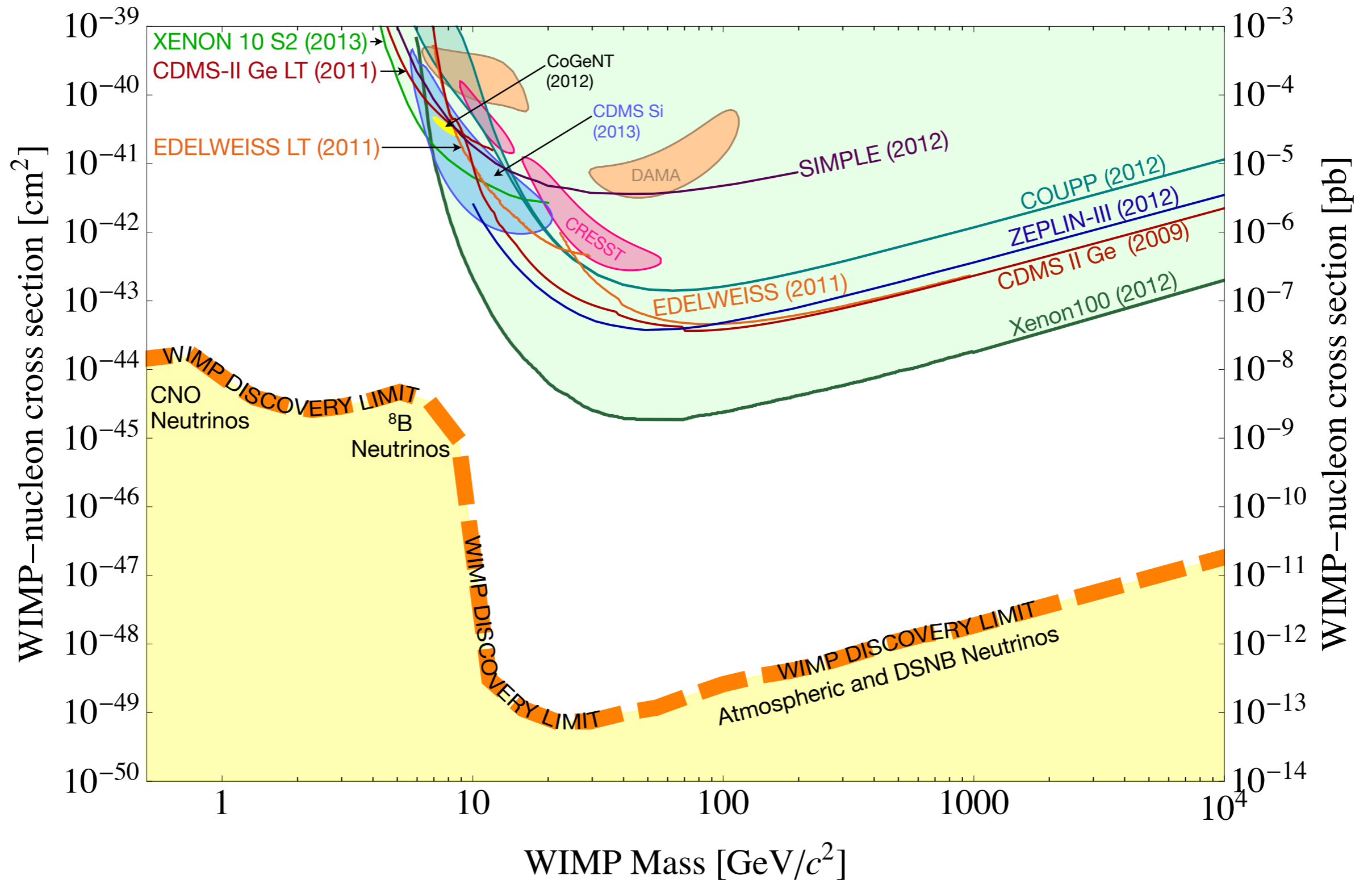
The screenshot shows the radiopurity.org website interface. At the top, the logo 'radiopurity.org' is displayed in red, with the tagline 'Community Material Assay Database' below it. A navigation bar contains 'Search', 'Submit', 'Settings', and 'About' buttons. A search input field contains the text 'copper'. Below the search bar, a table lists search results for various copper materials and their associated radionuclide levels.

Sample ID	Material	Radionuclide	Activity	Unit	Activity	Unit	...	Action
EXO (2008)	Copper, OFRP, Norddeutsche Affinerie	Th	< 2.4 ppt	U	< 2.9 ppt		...	X
EXO (2008)	Copper tubing, Metallica SA	Th	< 2 ppt	U	< 1.5 ppt			X
ILIAS ROSEBUD	Copper, OFHC							X
XENON100 (2011)	Copper, Norddeutsche Affinerie	Th-228	21() muBq/kg	U-238	70() muBq/kg		...	X
XENON100 (2011)	Copper, Norddeutsche Affinerie	Th-228	< 0.33 mBq/kg	U-238	< 11 mBq/kg		...	X
EXO (2008)	Copper gasket, Serto	Th	6.9() ppt	U	12.6() ppt		...	X
EXO (2008)	Copper wire, McMaster-Carr	Th	< 77 ppt	U	< 270 ppt		...	X

<http://radiopurity.org>

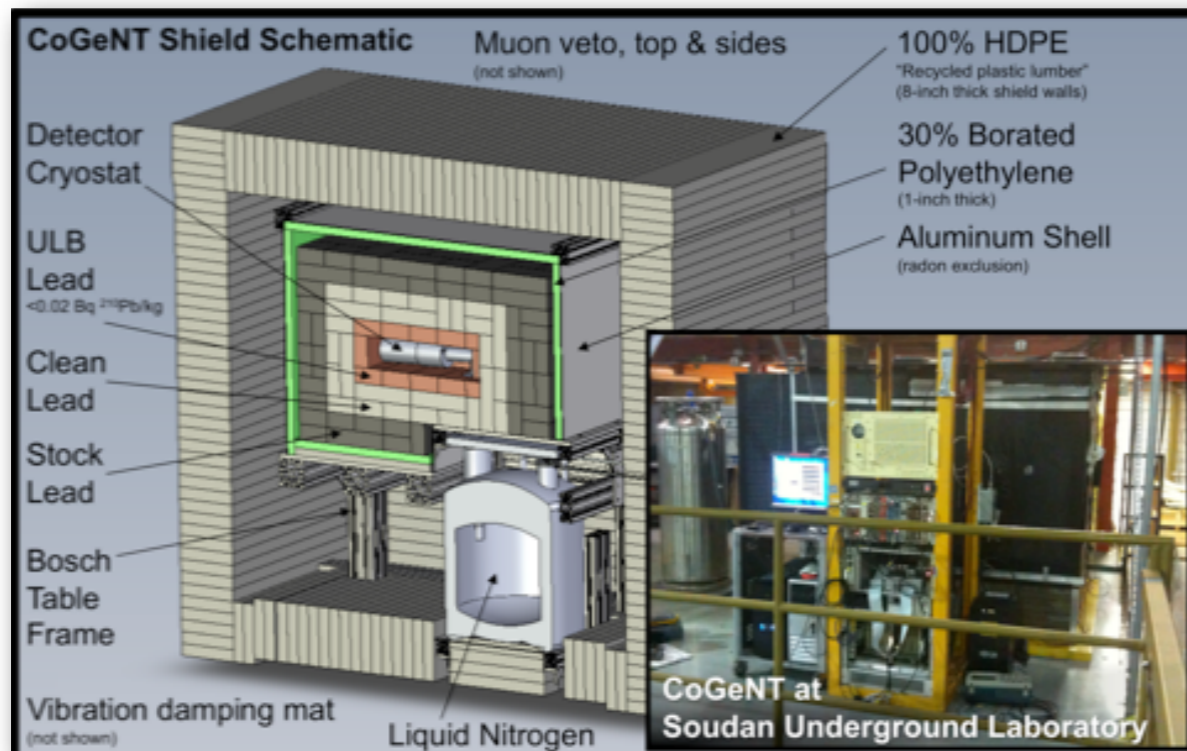
Supported by AARM, LBNL, MAJORANA, SMU, SJTU & others

Where Are We Now?



The Experiments Part 1: The Low Mass Region
Excesses Reported by DAMA/LIBRA,
CoGeNT, CRESST and CDMS

CoGeNT

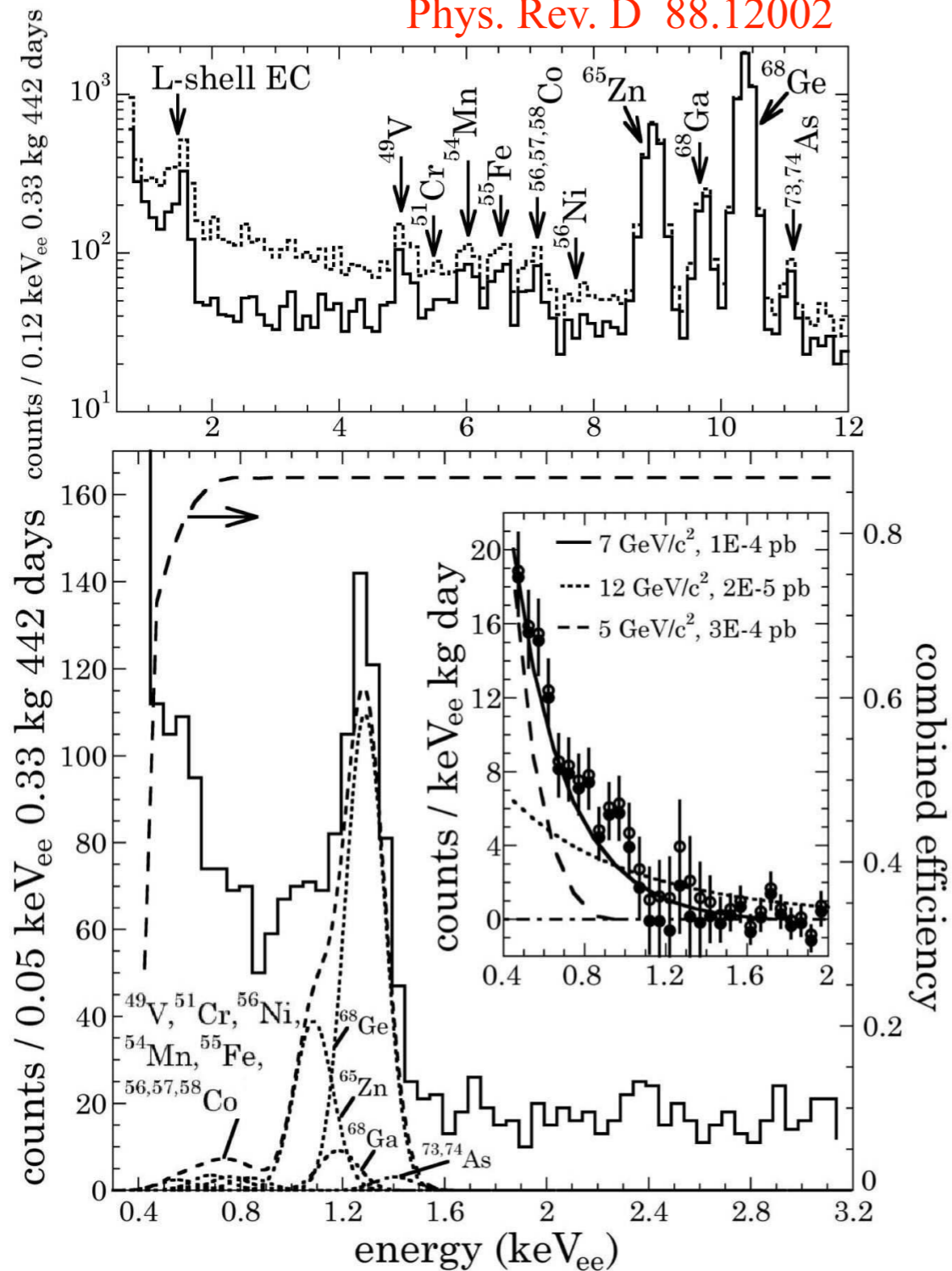


- Location: Soudan Underground Laboratory, Minnesota, USA
- 440 g HPGe ionization spectrometer
- Data collection from Dec. 4, 2009 - Mar. 6, 2011 (442 live days)
- Data collection interrupted due to fire.
- Data collection resumed July 2011.



CoGeNT

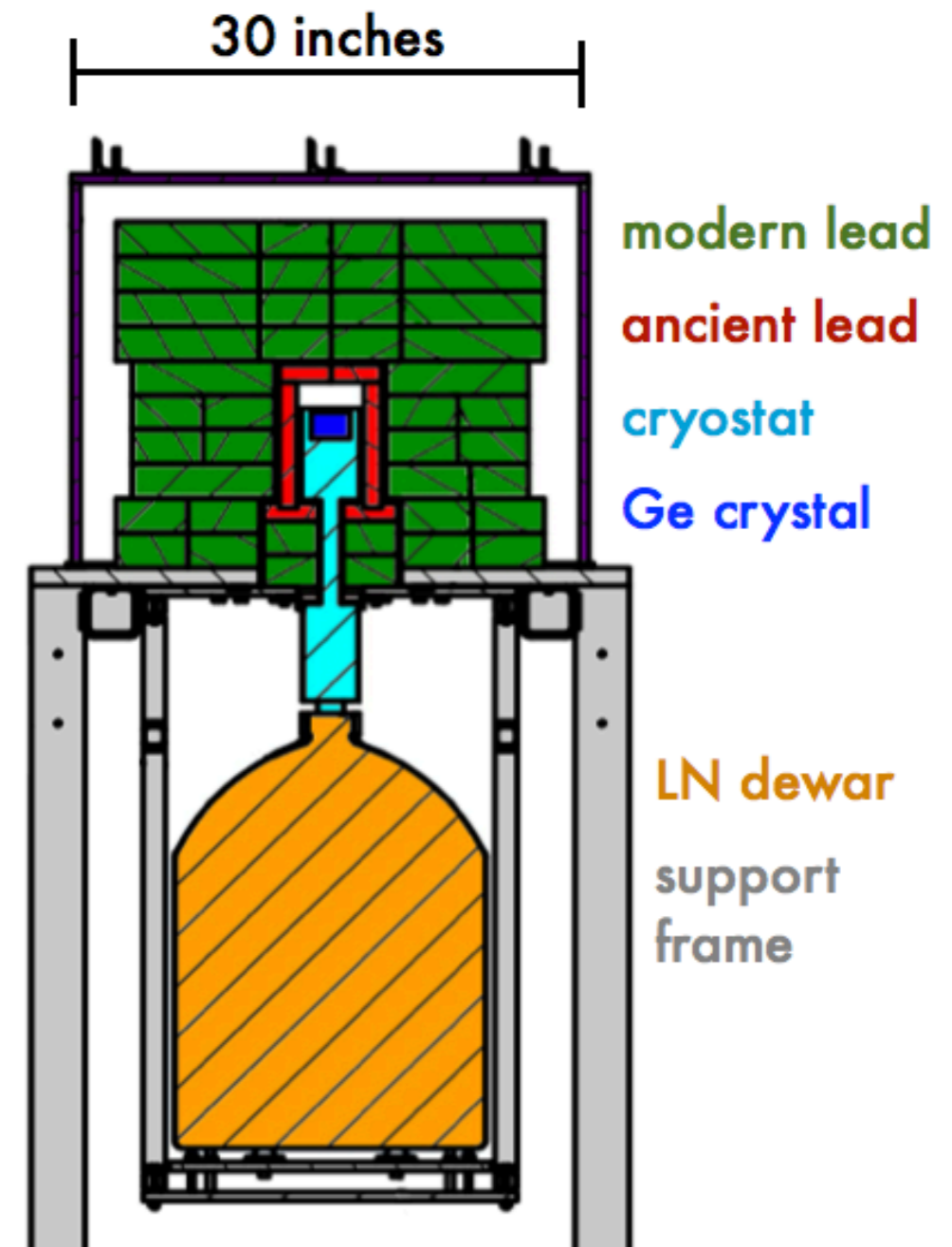
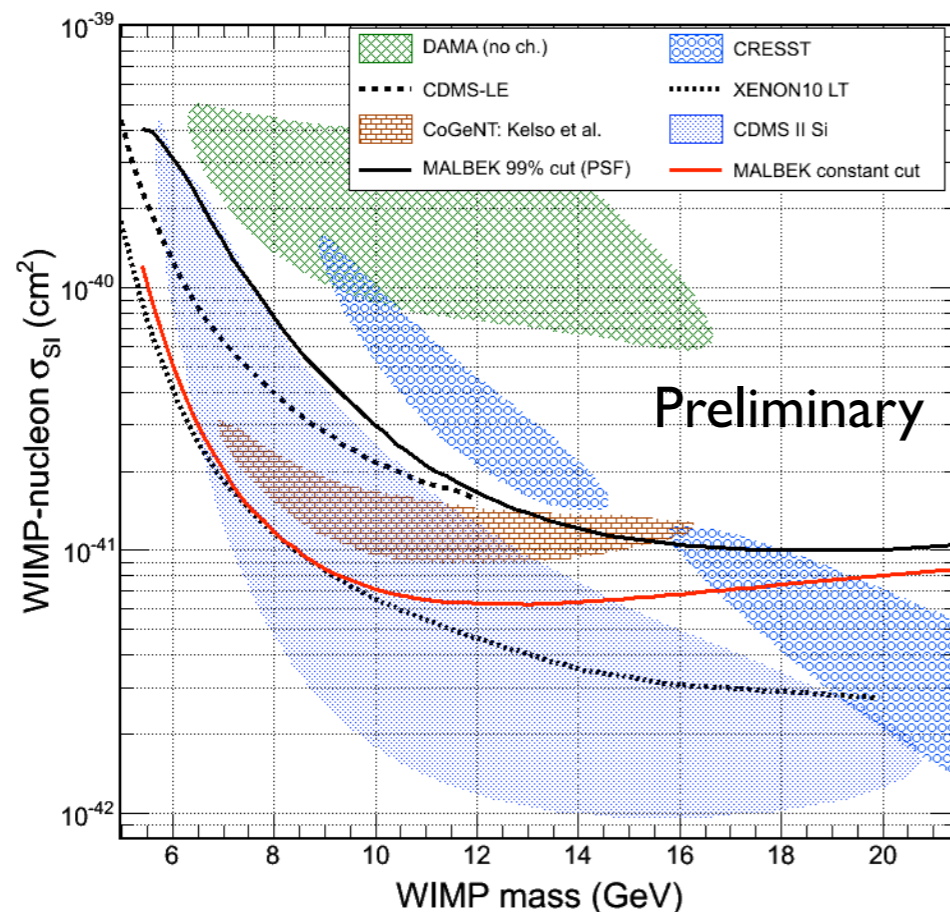
Phys. Rev. D 88.12002



- First claim of excess in 2010.
- Reject surface events using risetime cut (2011).
- Peaks due to cosmogenic activation of Ge
- After subtraction of known background, an exponential excess of events remains
- Fits to a variety of light-WIMP masses and couplings shown in inset of lower figure.
- Publication of new data coming soon.

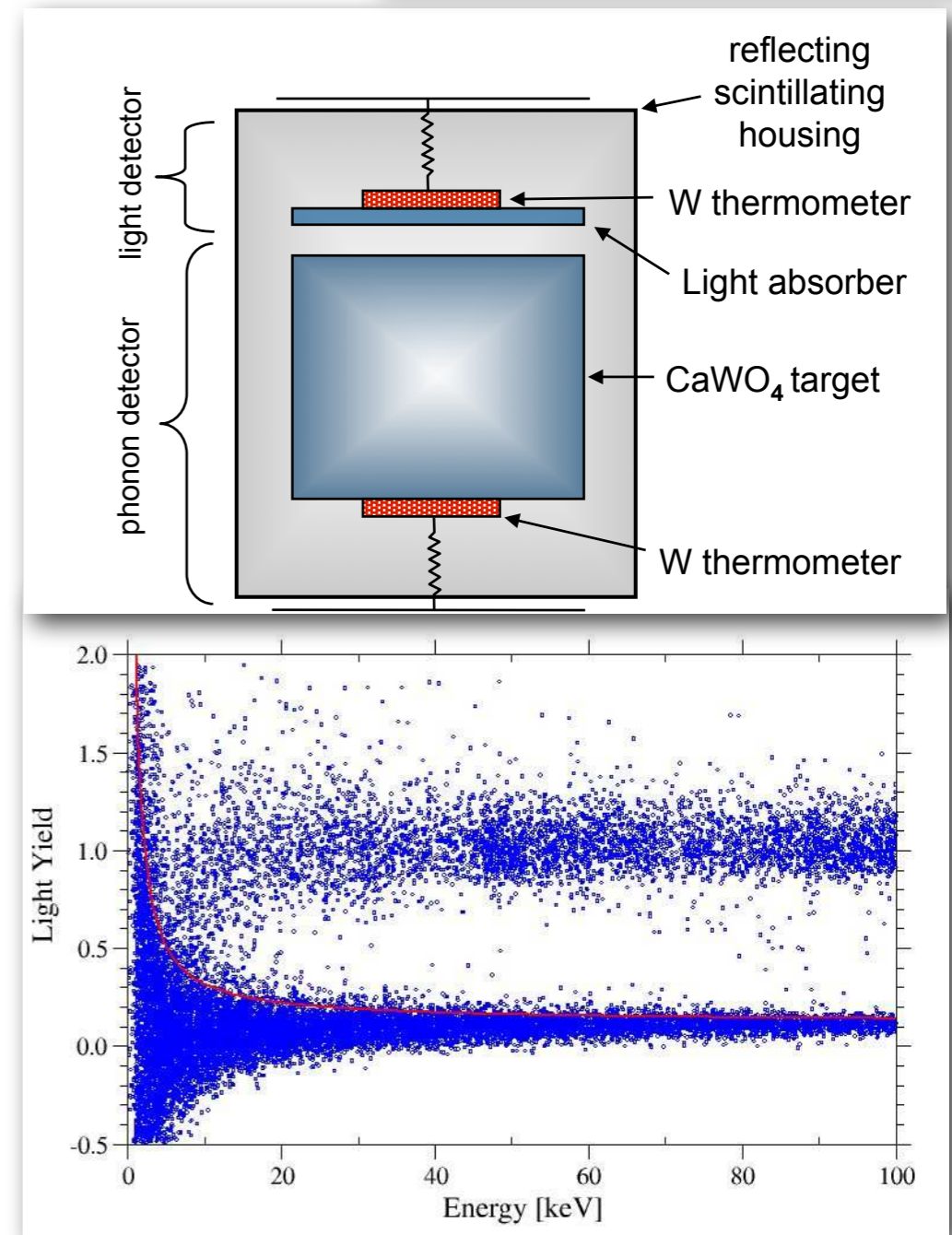
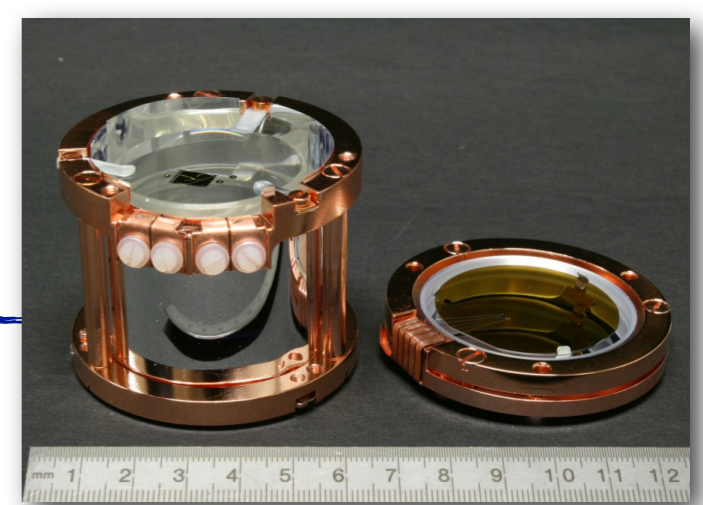
MALBEK

- **MAJ**ORANA **L**ow-background **BE**Ge detector at **K**URF.
- 450g Canberra Broad Energy Ge (BEGe) detector with ultra-low background components provided by J.I. Collar.
- Location: Kimballton Underground Research Facility (KURF), VA at 1450 mwe.
- 90% exclusion limits from 221 day data run.



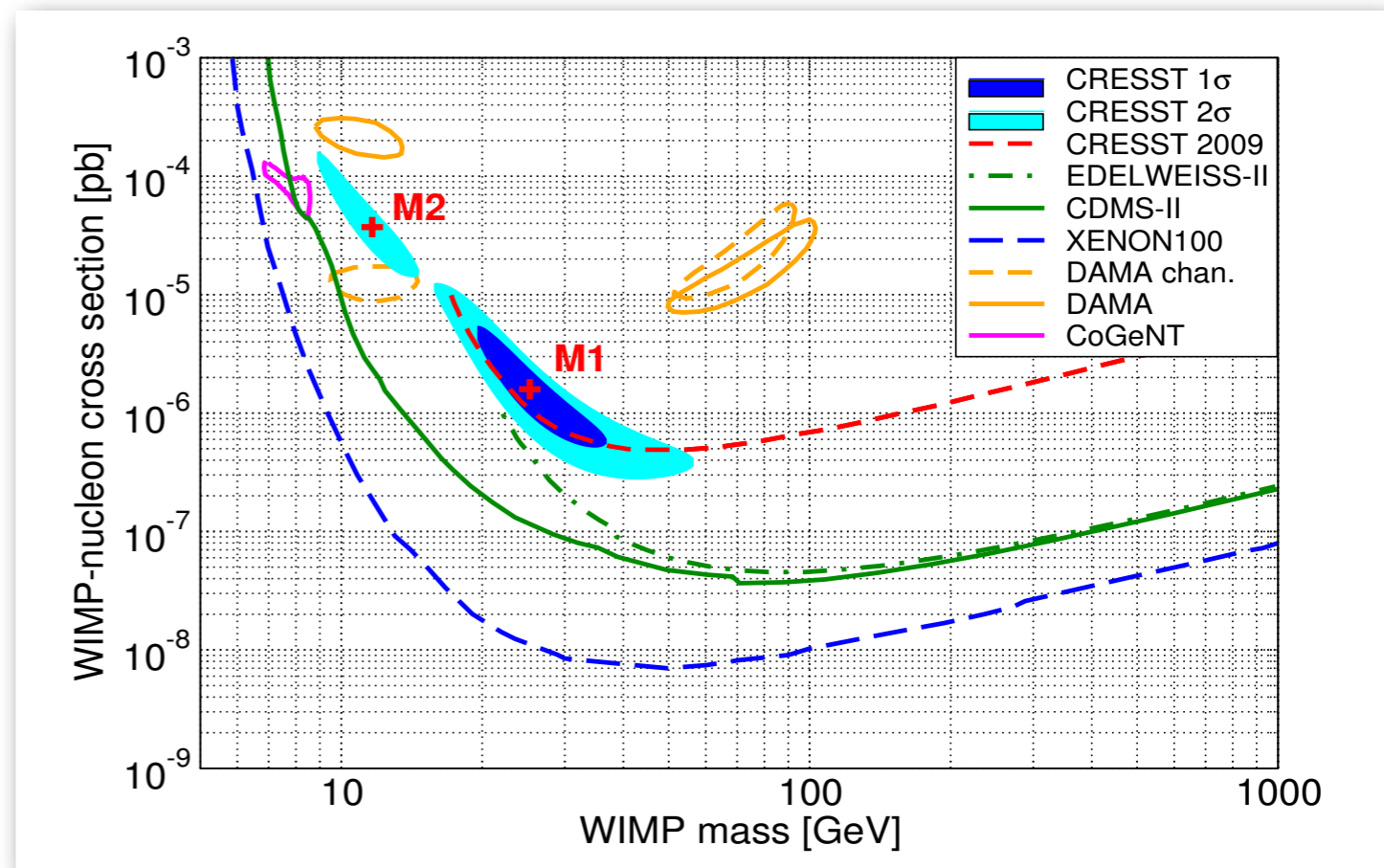
CRESST

- Cryogenic CaWO_4 crystals (~ 300 g each) are instrumented to readout phonon energy and scintillation.
- Location: Laboratori Nazionali del Gran Sasso, Italy
- Discrimination between ER and NR events via light yield (light/phonon energy)
- Net exposure: 730 kg-day (July 2009 - March 2011) from 8 detector modules.
- Observed 67 events in acceptance region (orange). [arXiv:1109.0702](https://arxiv.org/abs/1109.0702)
 - Analysis used a maximum likelihood in which 2 regions favored a WIMP signal in addition to predict background.
 - Excess events can not be explained by known backgrounds
 - Large background contribution



CRESST Plans

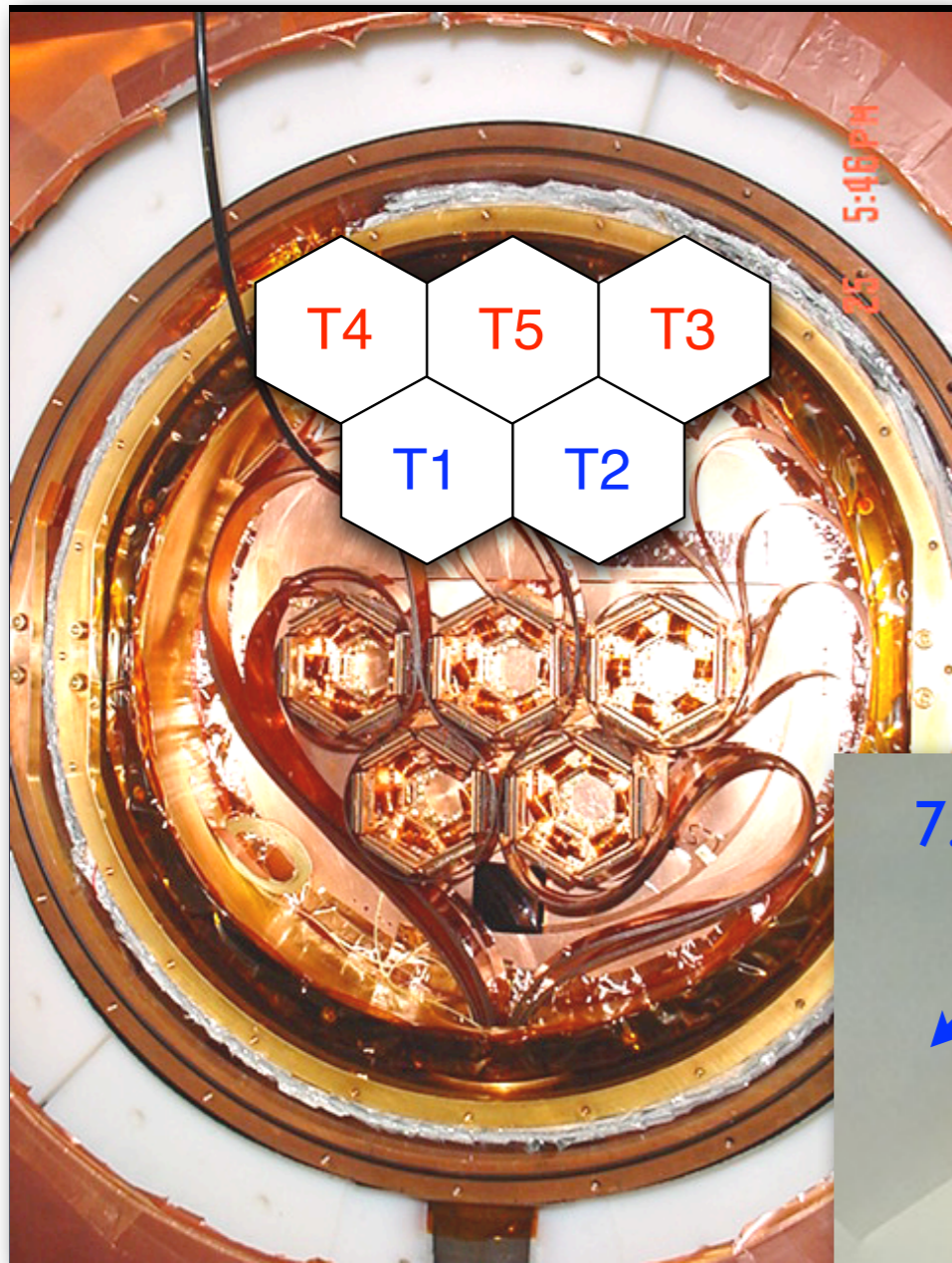
- Current data run aims to reduce background, increase detector mass.
 - Alphas - new clamping design and material
 - Detector assembly in a radon free environment
 - New detector design to discriminate ^{206}Po recoils
 - Add additional shielding to reduce neutron background
- June & July calibration runs with ^{57}Co source were successful.
- **July 30th, 2013 Science Runs Begin!**



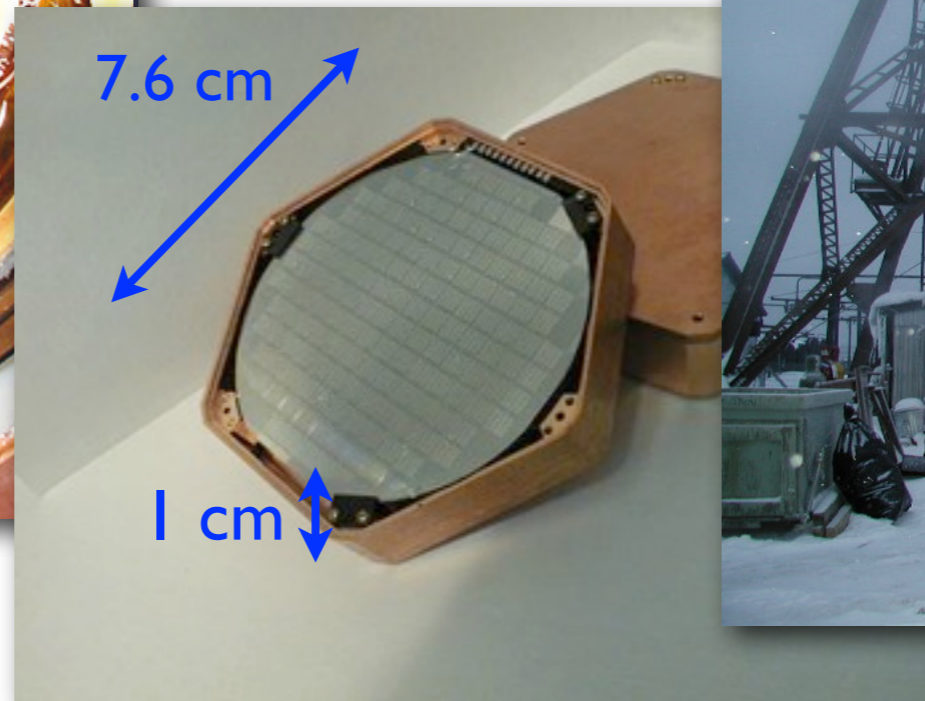
[arXiv:1109.0702](https://arxiv.org/abs/1109.0702)

CDMS II

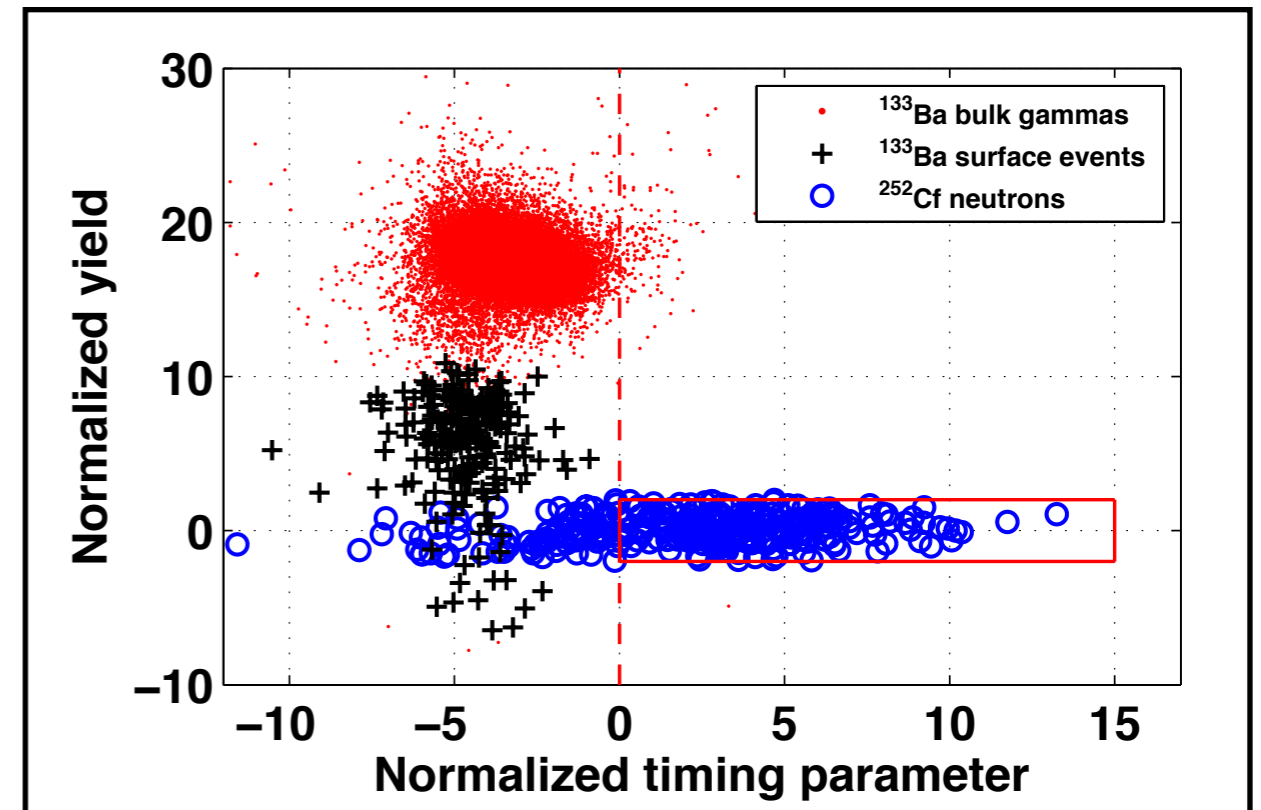
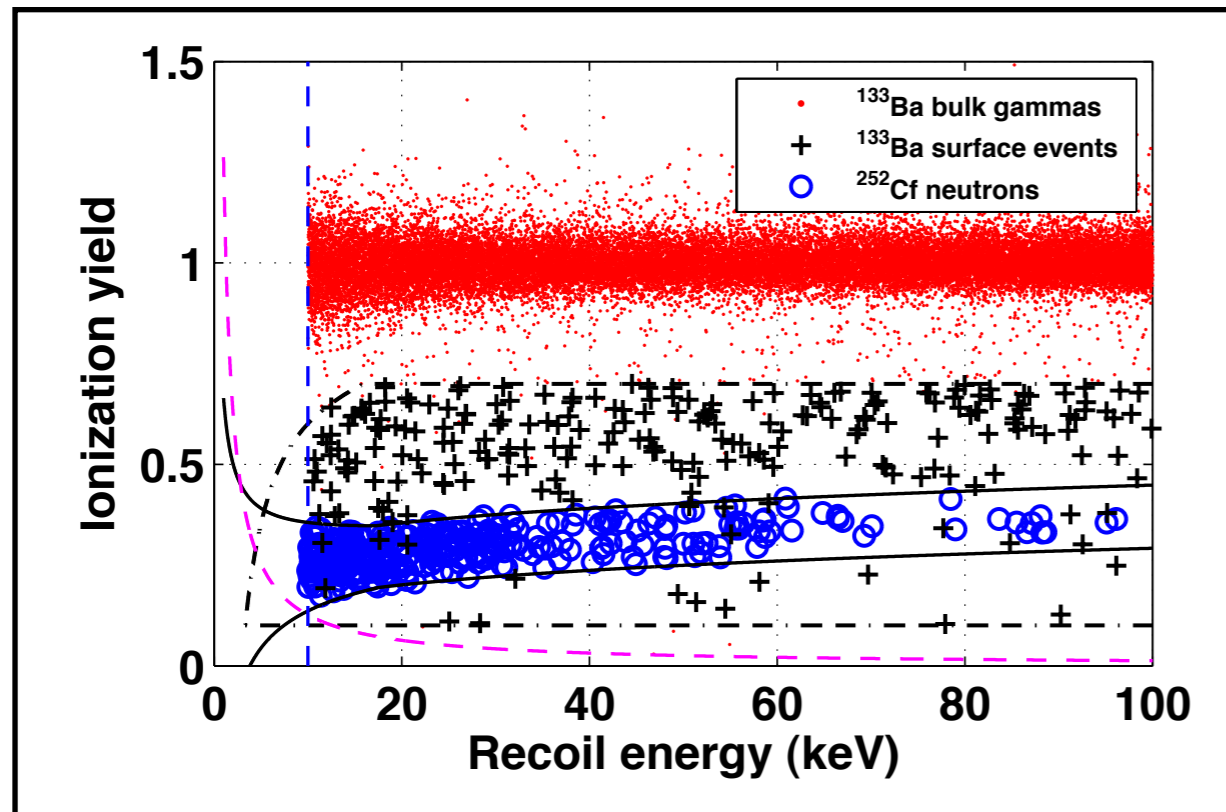
Billard - Mon. DM I
Nelson - Mon. DM II
Speller - Poster



- CDMS II: 30 detectors (19 Ge, 11 Si) installed and operated in the Soudan Underground Laboratory, MN, USA from Jun. 06 - Mar. 09.
- Measures ionization and phonons (read out by TES)
- Science Results: CDMS-Ge, CDMS-Si, Ge-Low Threshold, Annual Modulation

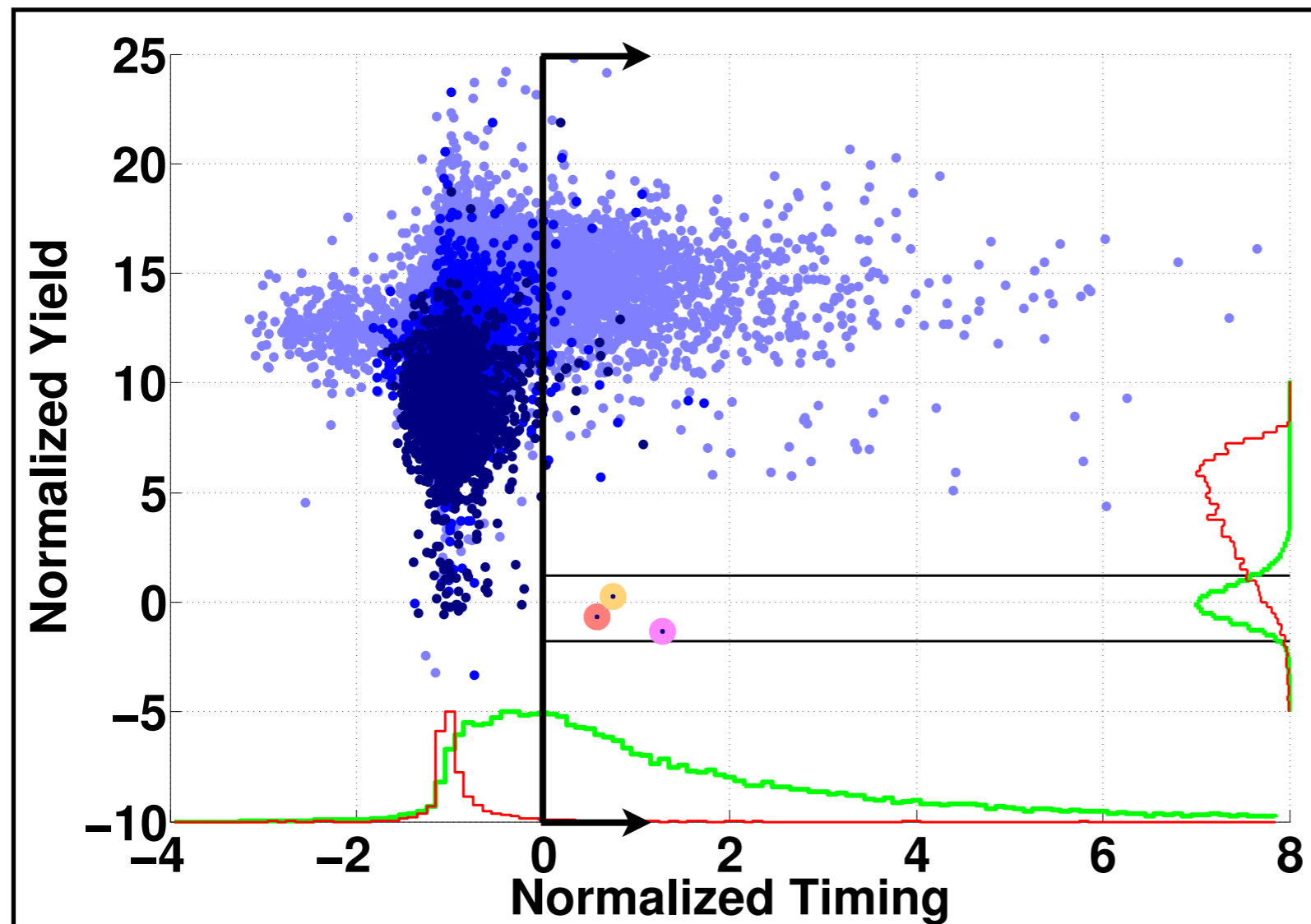


Background Rejection



- Most backgrounds produce electron recoils and have yield (ionization/phonon energy) ~ 1 .
- WIMPs and neutrons produce nuclear recoils and have yield ~ 0.3 .
- Surface events can be identified using timing properties of the phonon and charge pulses.

Recent Results: CDMS II-Si Detectors



Observed 3 events.

- Shades of blue indicate three separate timing cut energy ranges.

- 7- 20 keV

- 20 - 30 keV

- 30 - 100 keV

- Background Estimate

- Surface Events

$$0.41^{+0.20}_{-0.08}(\text{stat.})^{+0.28}_{-0.24}(\text{syst.})$$

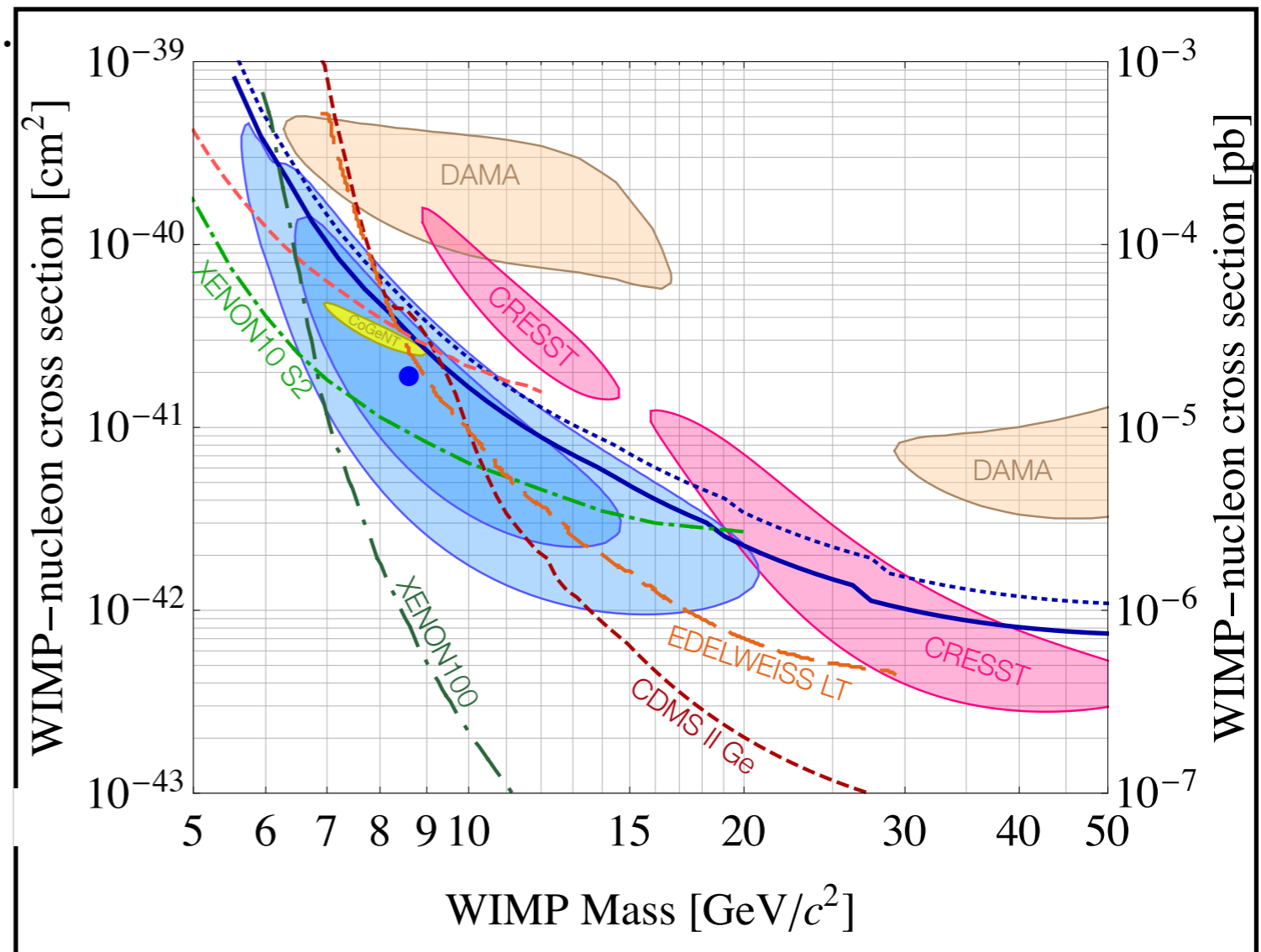
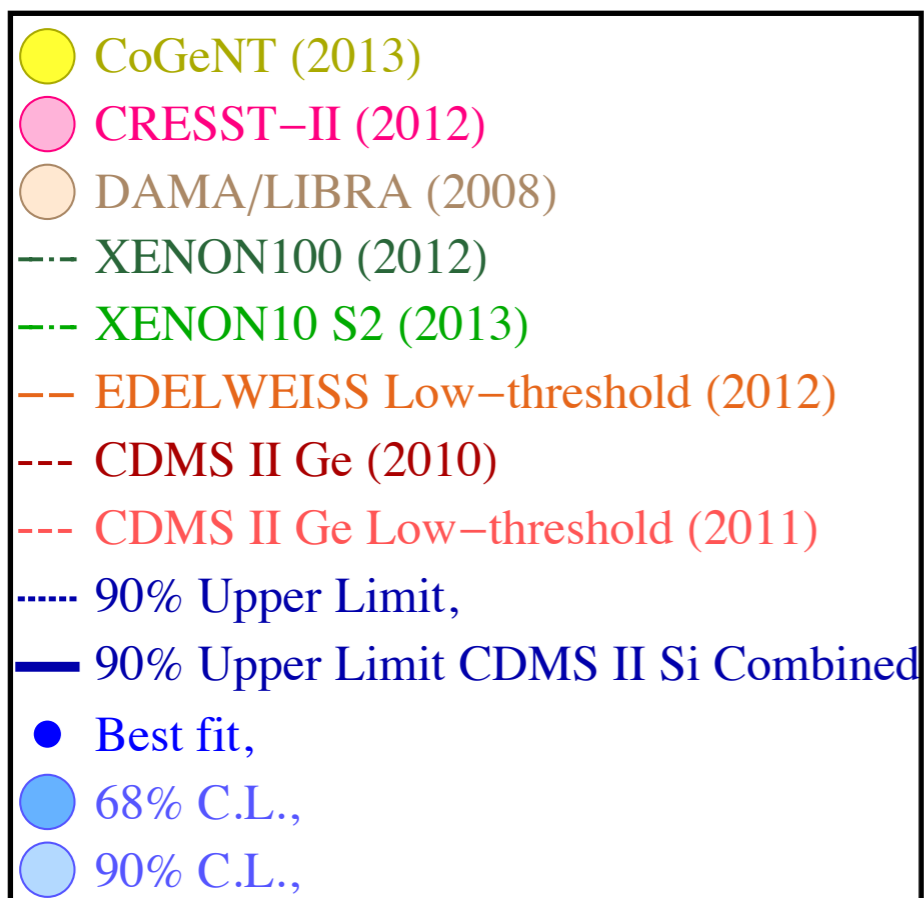
- < 0.13 neutrons from
Cosmogenics &
Radiogenics

- < 0.08 ^{206}Pb recoils from
 ^{210}Pb decays

CDMS II - Si Results

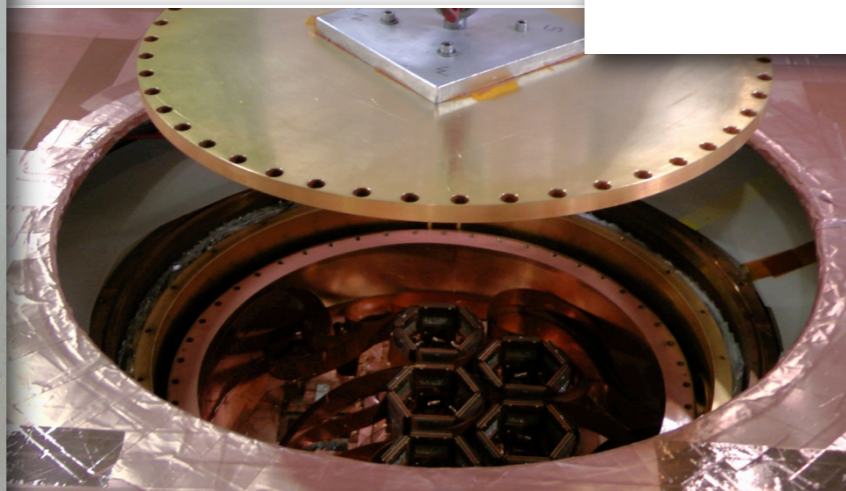
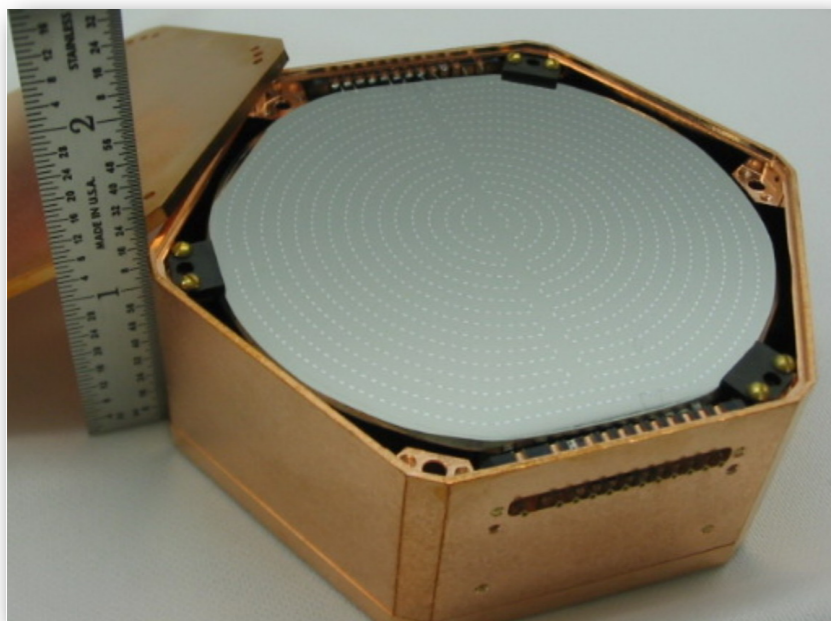
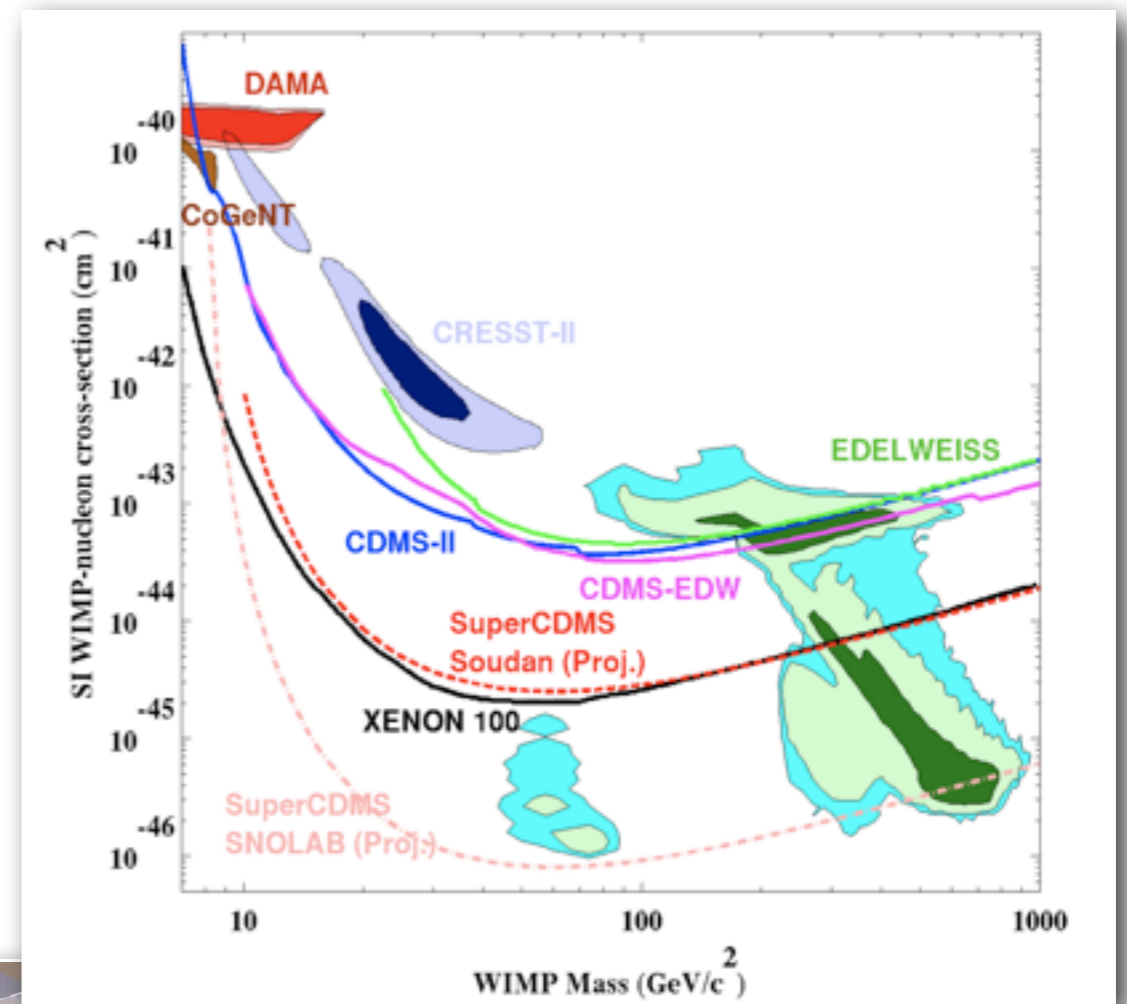
- Three events observed in the signal region.
- A profile likelihood analysis favors a WIMP+background hypothesis over the known background estimate as the source of our signal at the 99.81% C.L. ($\sim 3\sigma$, p-value: 0.19%)

- The maximum likelihood occurs at a WIMP mass of $8.6 \text{ GeV}/c^2$ and WIMP-nucleon cross section of 1.9×10^{-41} .
- Does not rise to level of discovery, but does call for further investigation.



SuperCDMS @ Soudan

- Currently operating 5 towers of advanced iZIP detectors (~9 kg Ge) in the existing cryostat at the Soudan Underground Laboratory.
- After 3 years of operation, expected to improve sensitivity to spin-independent WIMP-nucleon interactions by a factor of ~10 over existing CDMS II results.



Installation complete Nov. 8, 2011. Operating with final detector settings since Mar. 2012.

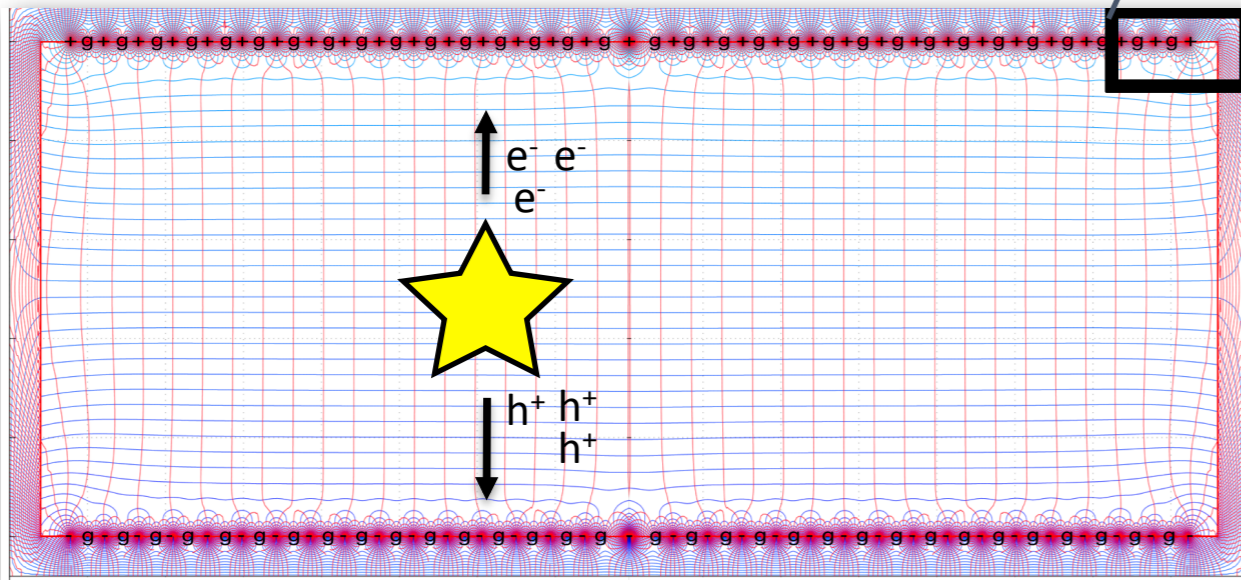
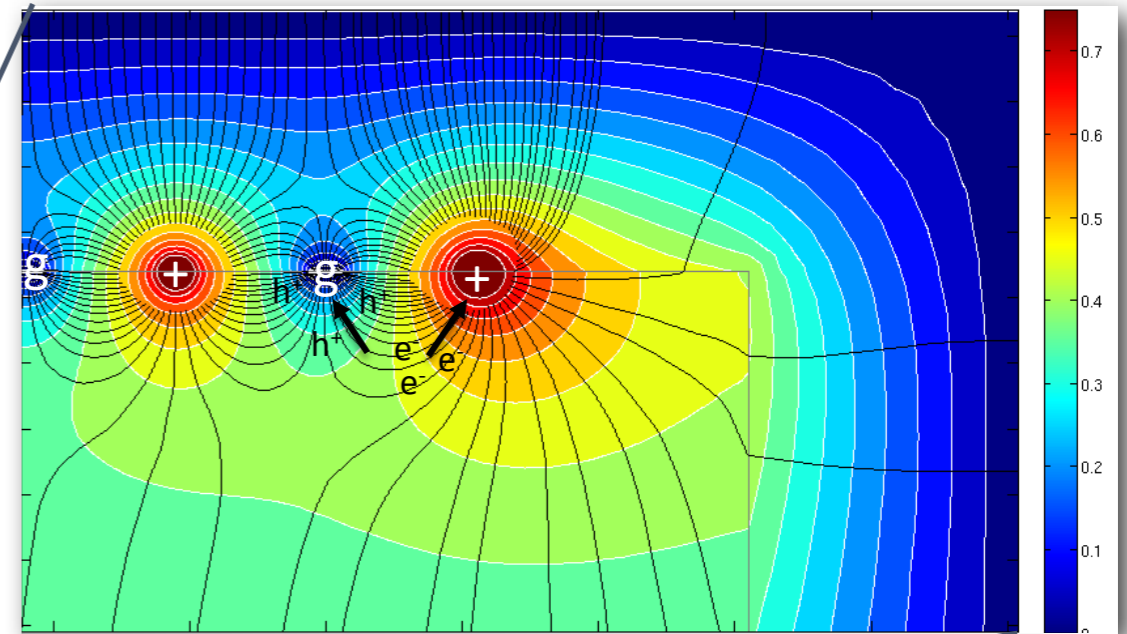
SCDMS iZIPs: Charge Signal

Bulk Events:

Equal but opposite ionization signal appears on both faces of detector (symmetric)

Surface Events:

Ionization signal appears on one detector face (asymmetric)



arXiv:1305.2405

SCDMS iZIPs: Charge Signal

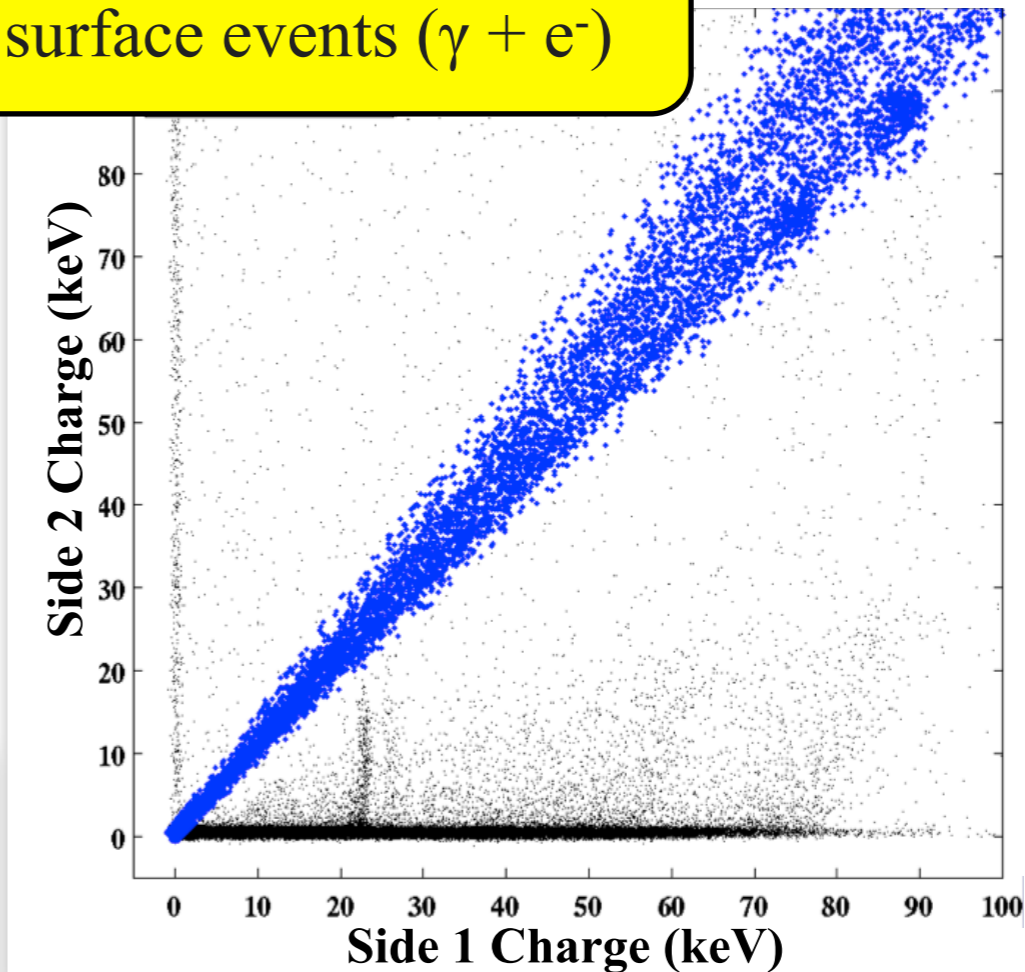
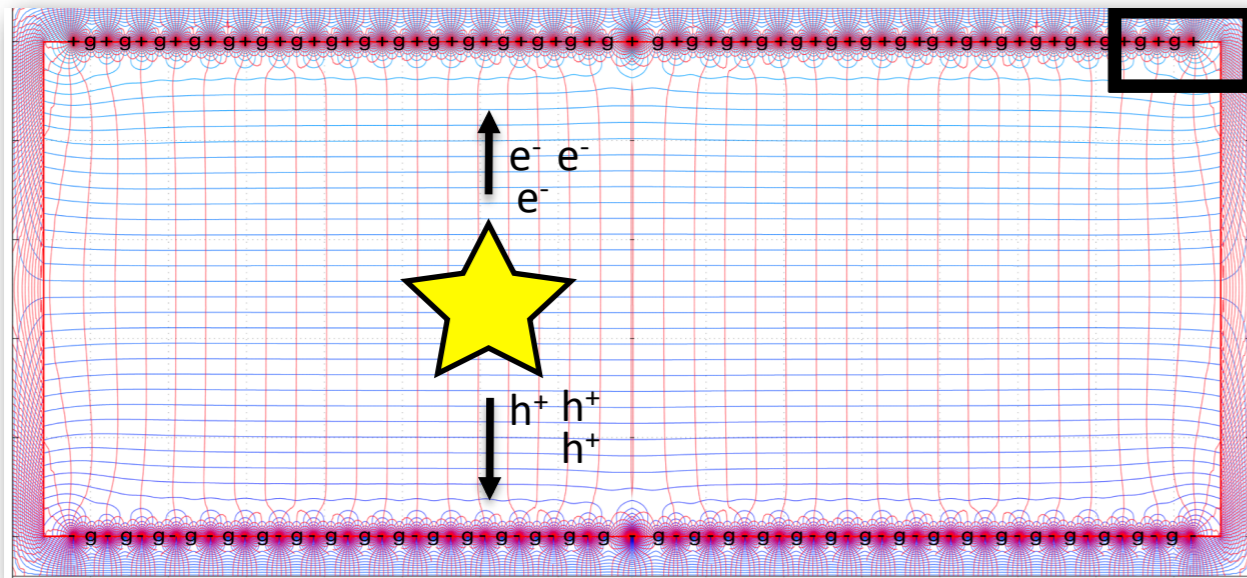
- bulk events (γ)
- surface events ($\gamma + e^-$)

Bulk Events:

Equal but opposite ionization signal appears on both faces of detector (symmetric)

Surface Events:

Ionization signal appears on one detector face (asymmetric)



- ~50% fiducial volume (8-115 keVr)
- < 0.6 events in 0.3 ton-years
- **Good enough for a 200 kg experiment run for 4 years at SNOLAB!**

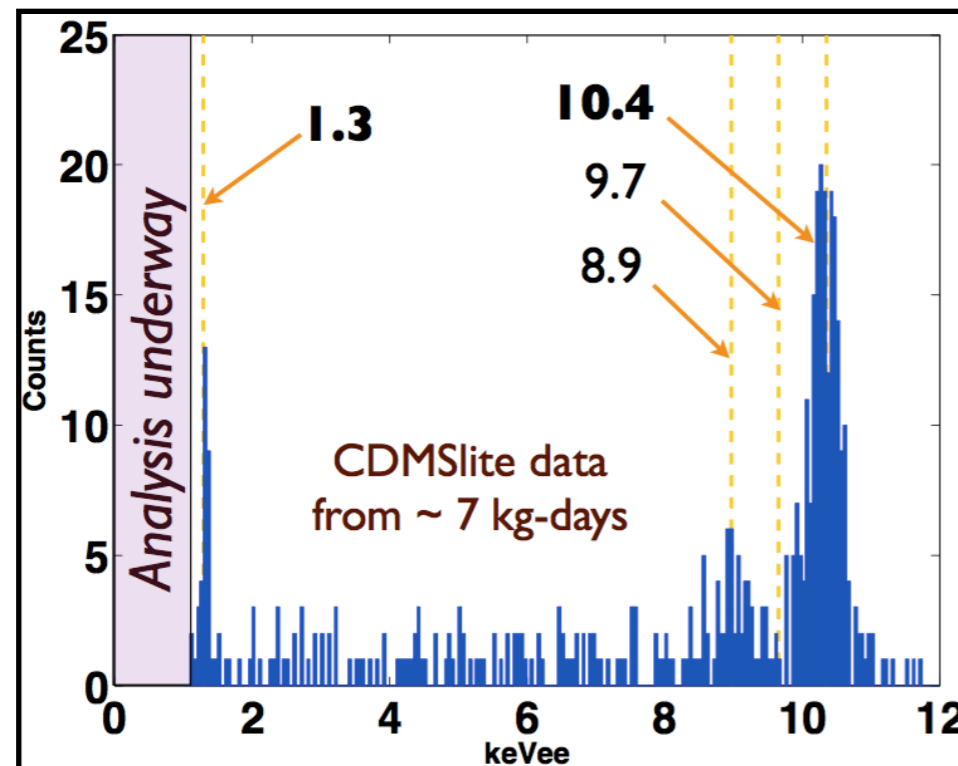
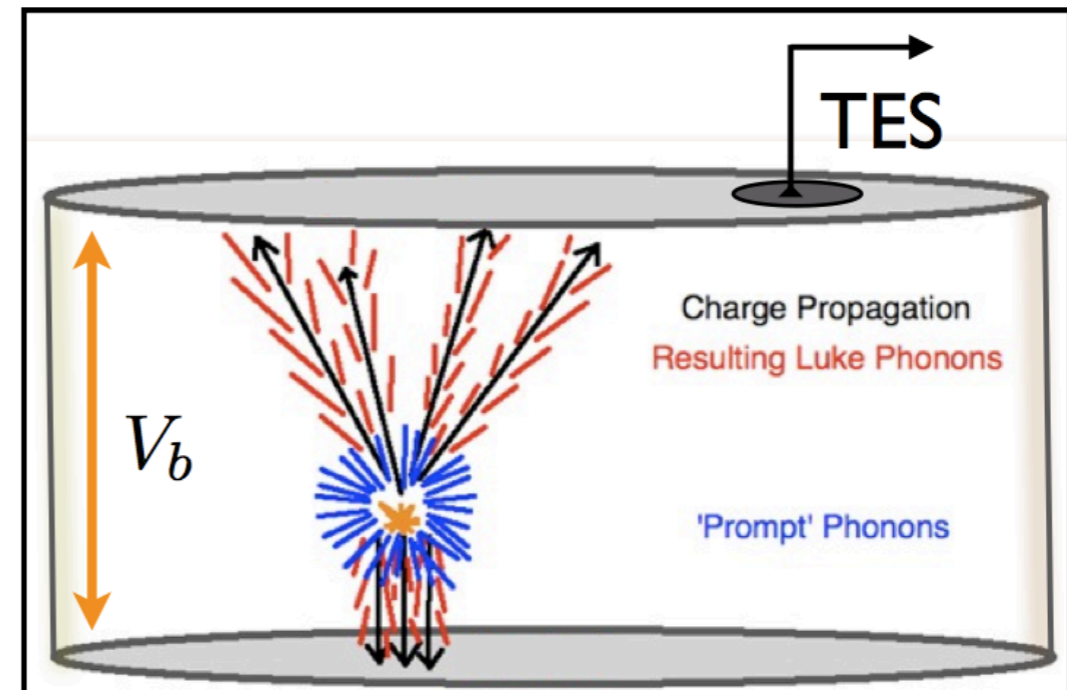
arXiv:1305.2405

CDMSlite

- Alternate running mode to explore low mass WIMPs utilizing Luke phonons

$$E_{\text{luke}} = N_{e/h} \times eV_b$$

- Luke energy scales as bias voltage and noise remains constant until breakdown



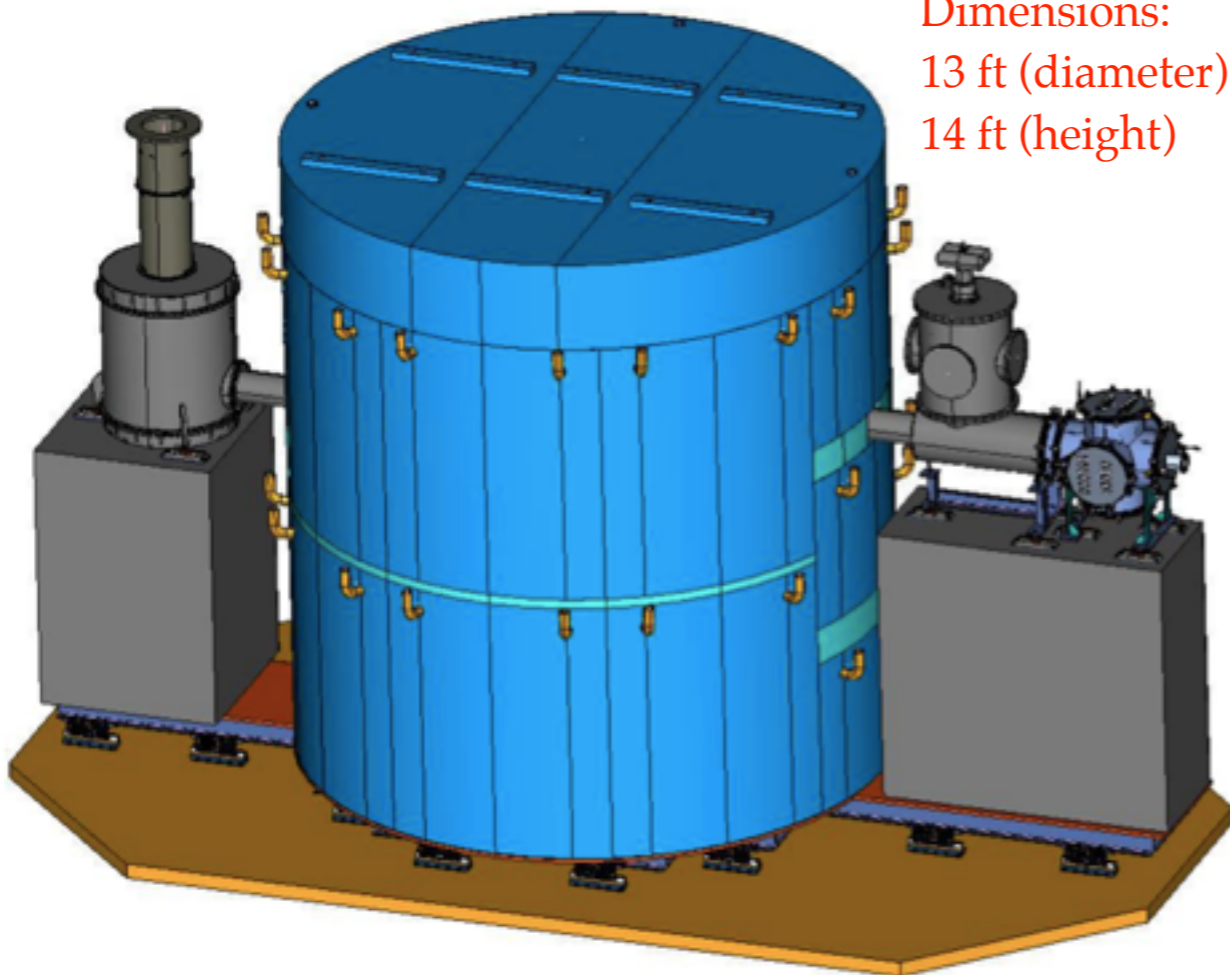
- Resulting Luke amplification has excellent energy resolution potentially down to 1.3 eV_{ee} .
- Resolution of various Ge activation lines.

**New Results to be
Announced Wednesday!**

Future: SuperCDMS @ SNOLAB

Planned Setup

Dimensions:
13 ft (diameter)
14 ft (height)

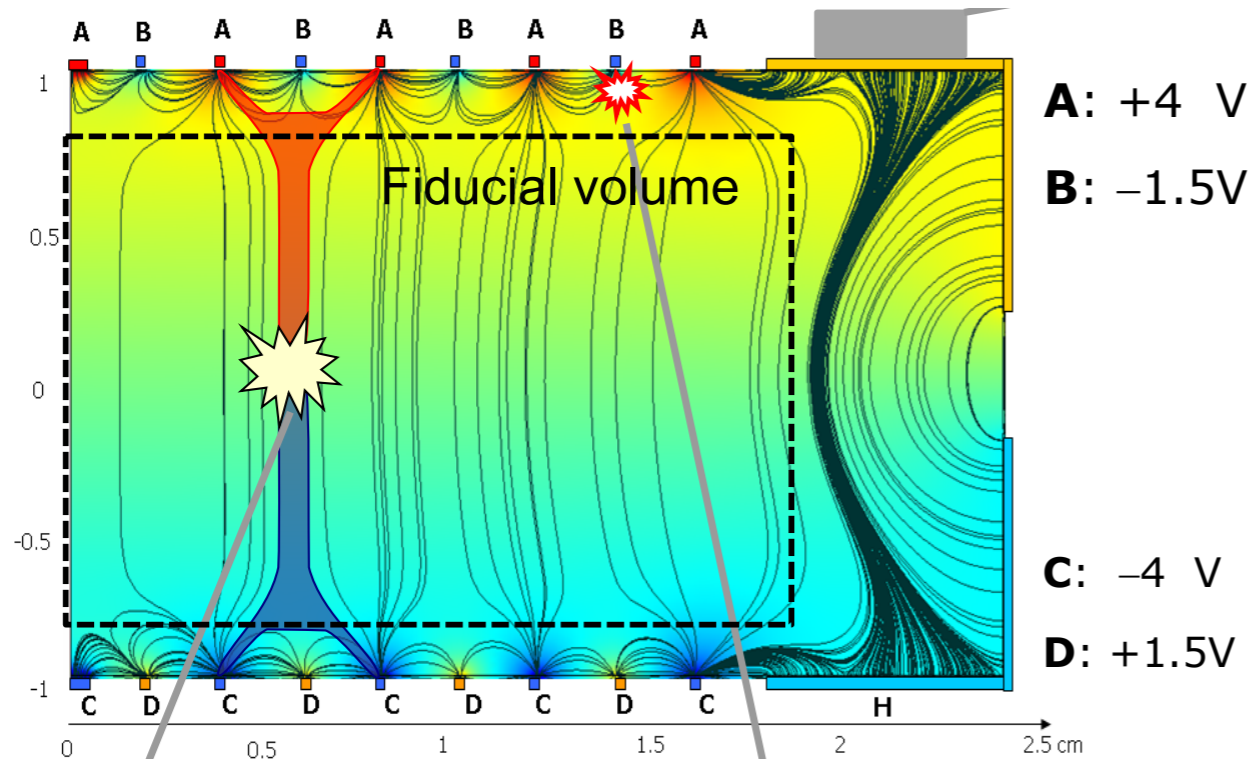


- cryostat volume of up to 400 kg target
- 200 kg experiment with sensitivity of $8 \times 10^{-47} \text{ cm}^2$ at $60 \text{ GeV}/c^2$
- Pb/Cu shielding for external radiation
- increased PE shielding (neutrons)
- possible neutron veto

- Calibration runs at Soudan indicate that the new iZIPs have good enough surface rejection capabilities for a 200 kg experiment at SNOLAB to run 4 years! ([arXiv:1305.2405](https://arxiv.org/abs/1305.2405))

EDELWEISS III

NTD thermal sensor



Bulk/Fiducial event
Charge collected on
electrodes A&C

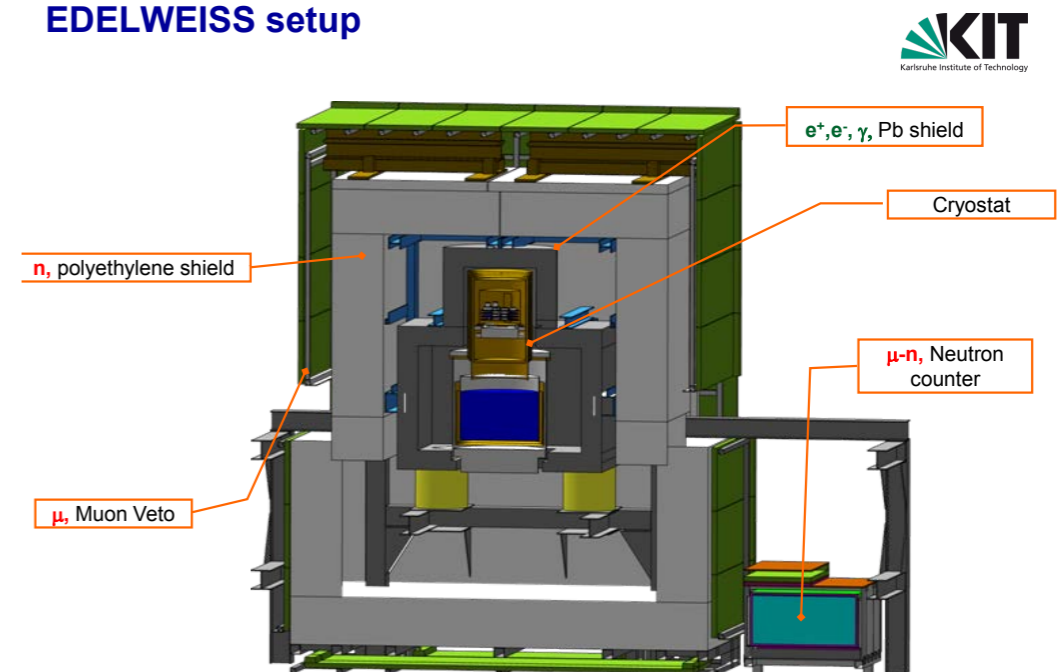
Surface event
Charge collected on
electrodes A&B

- Discrimination from ionization yield and charge collection symmetry.

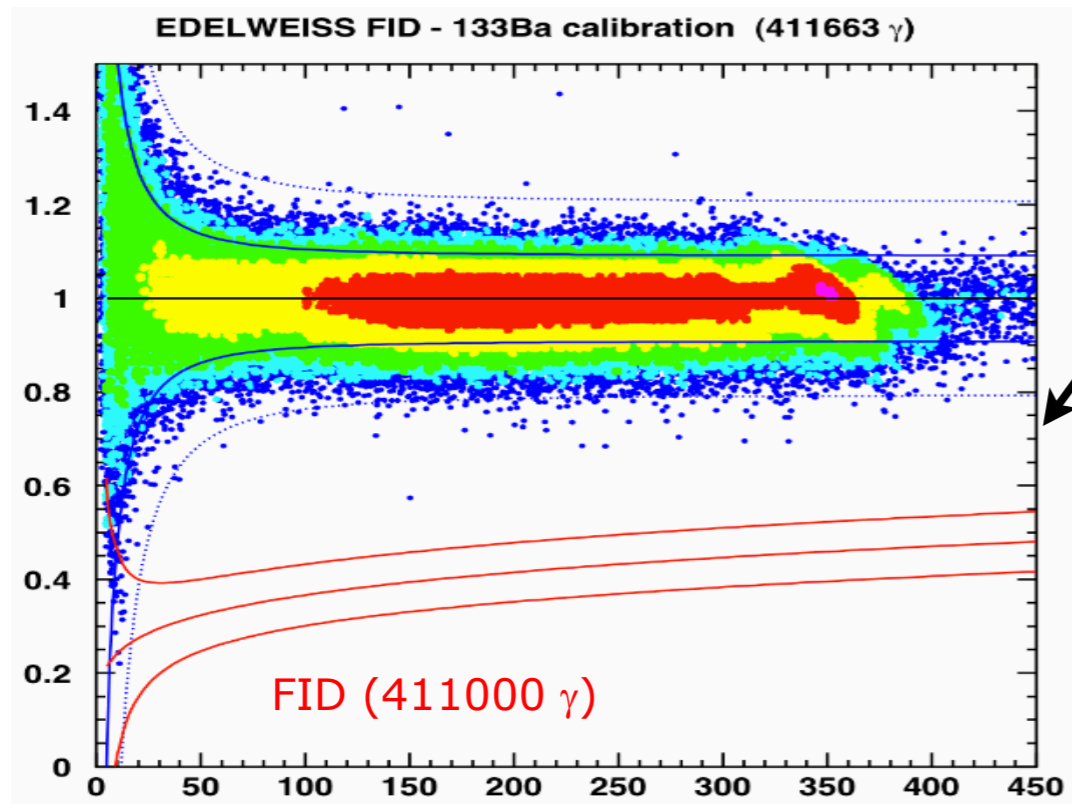
- Located in the Laboratoire Souterrain de Modane (LSM) between Italy and France.
- Detectors instrumented with electrodes to measure charge and NTD thermal sensors to measure phonon signal.

or

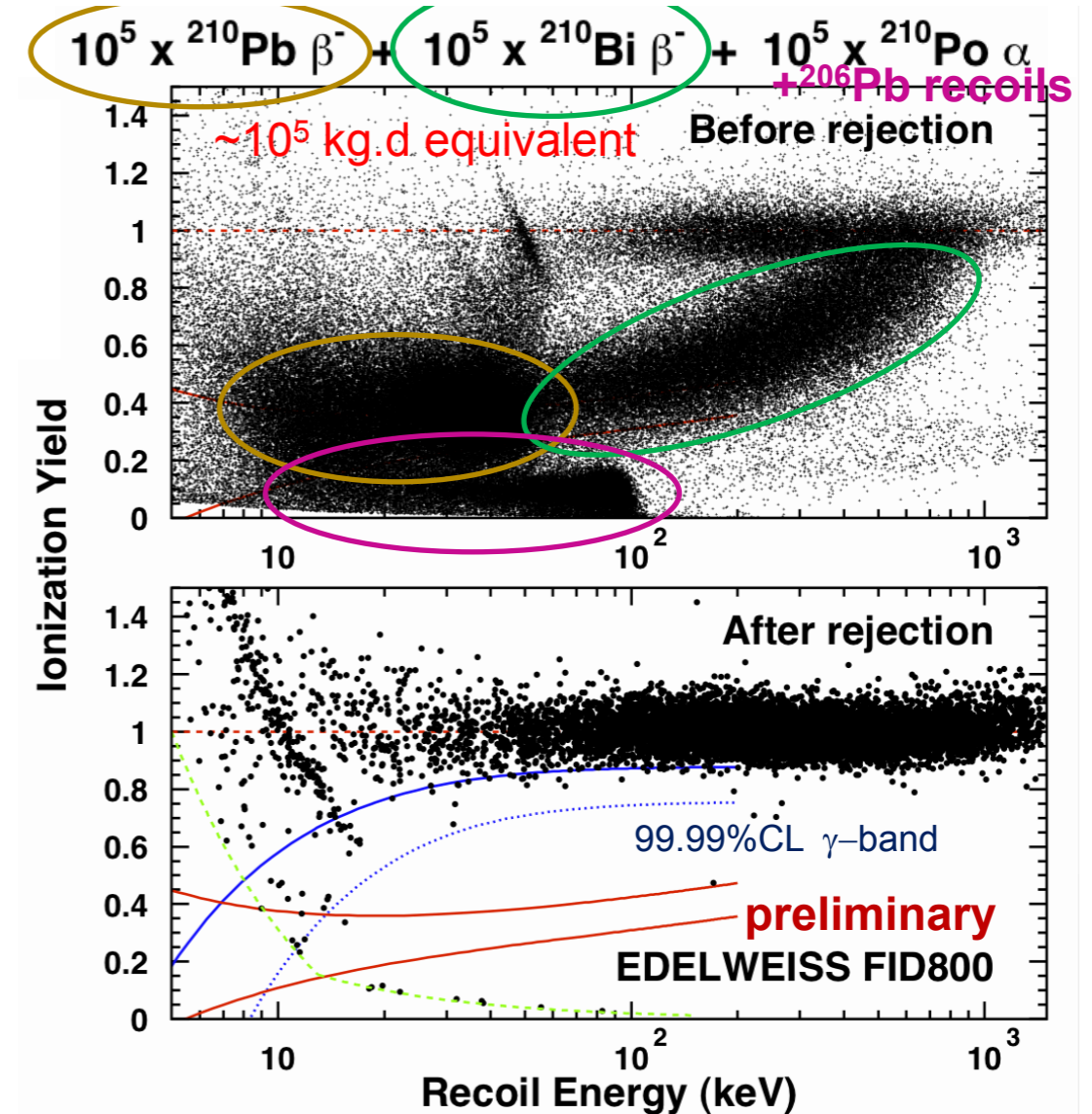
EDELWEISS setup



EDELWEISS III



-Improvements to γ -discrimination



-Improvements to surface event discrimination

- $< 4 \times 10^{-5}$ misID events per kg-d ($E_r > 15$ keV)

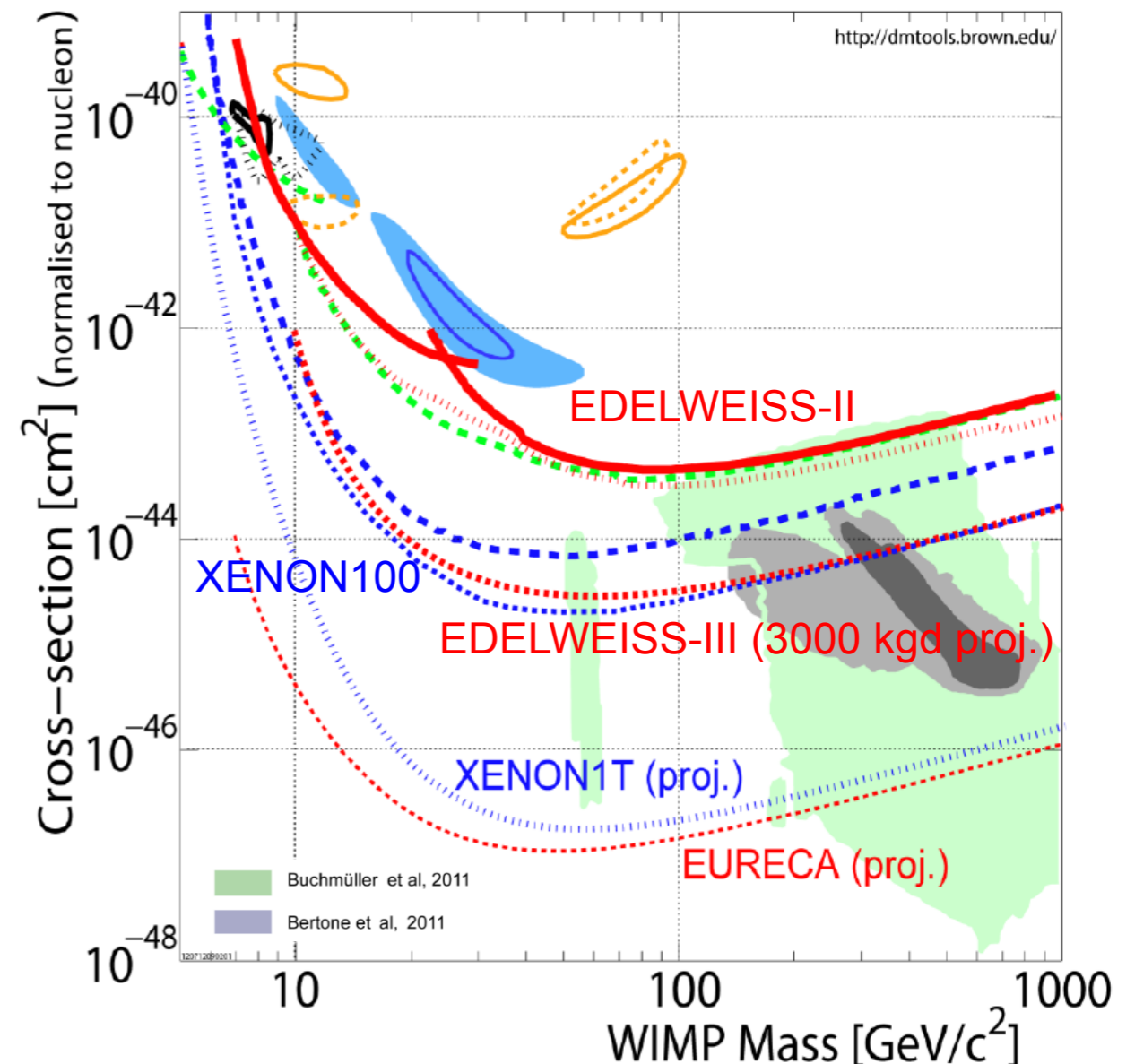
EDELWEISS III - Projections

Sept. 2013

- EDELWEISS III
Commissioning runs underway
- 15 FID detectors of mass 800g each
- upgraded cryostat, readout electronics and kapton cables
- New PE shield and copper screens

End of 2013

- Fully equipped cryostat with 40 FID detectors of 800g mass each.

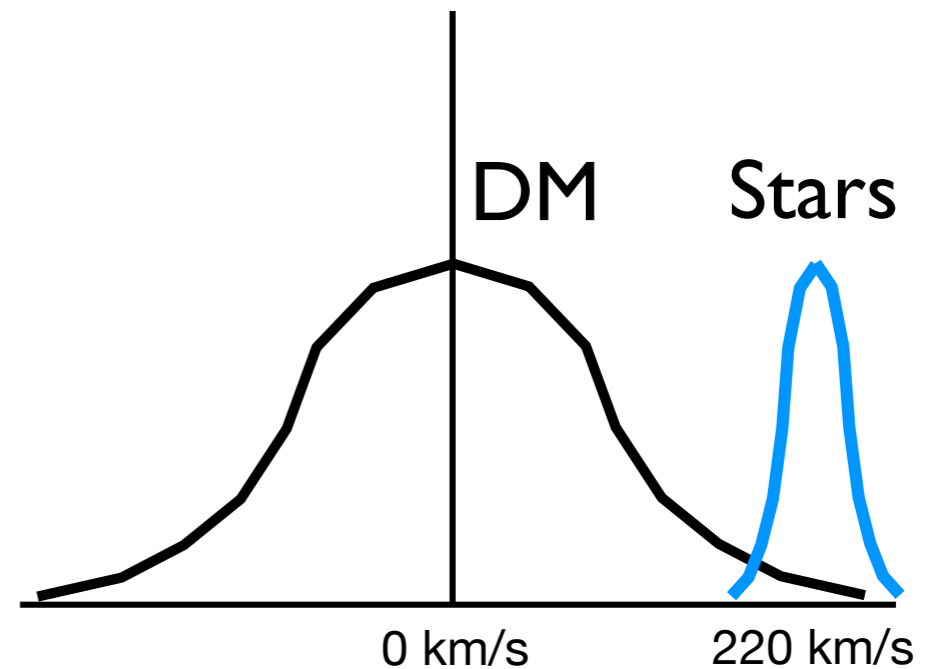


The Experiments Part 2: Addressing a Long Standing Issue

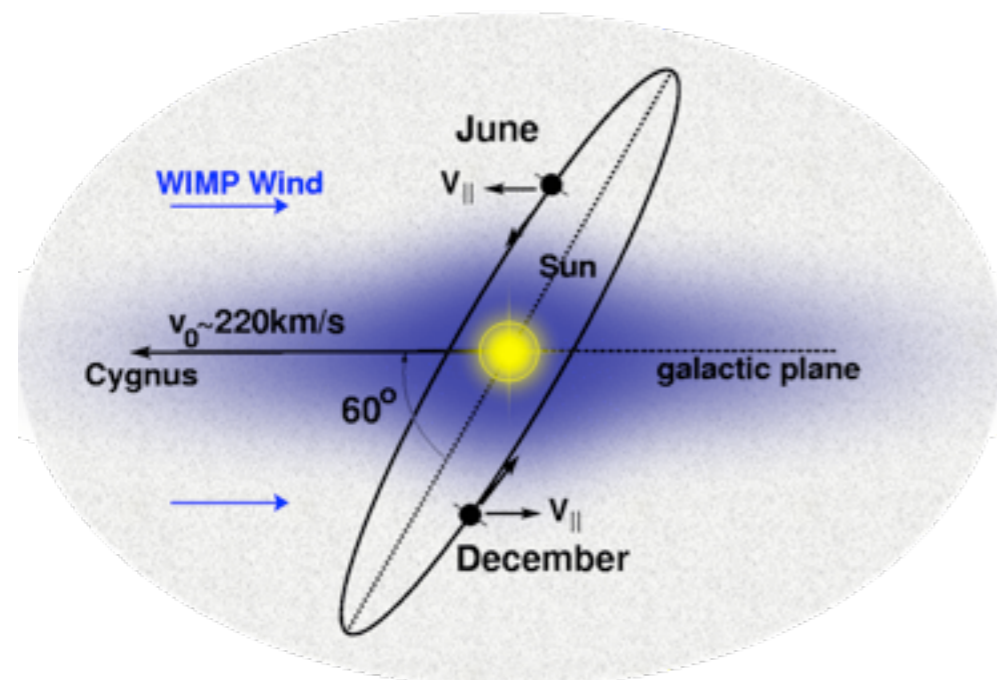
DAMA/LIBRA Modulation Signal

Signal Modulation

- Baryons travel together in roughly circular orbits with small velocity dispersion
- Dark matter particles travel individually with no circular dependence and large velocity dispersion



V_θ (at out galactic radius)



- As a result, the flux of WIMPs passing through Earth modulate over the course of a year as Earth rotates around the sun.