

# Other WIMP Direct Detection Experiments

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# Outline

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- Principles common to experiments
- The Experiments
  - Part 1: The low mass region
  - Part 2: The long standing DAMA/LIBRA experiment
  - Part 3: The search continues...
- Concluding Remarks

# World-Wide Experiments

## Phonon/Charge/Light:

CDMS/SuperCDMS

EDELWEISS

CRESST

## Charge Only:

CoGeNT/C4

TEXANO

CDEX

CDMSlite

## Multi-purpose:

Majorana Demonstrator

COURE-0/COURE

## Modulation:

DAMA/LIBRA

DM-ICE

KIMS

ANAIS

SABRE

KamLAND-PICO

## Bubble Chambers/Superheated:

PICASSO

COUPP

PICO

## Directional:

DRIFT

DM-TPC

## Other:

DAMIC

NEXT

\*Experiments in red are presenting results or status in parallel sessions.

# World-Wide Experiments

## Phonon/Charge/Light:

CDMS/SuperCDMS

EDELWEISS

CRESO

## Charge

CoGeNT

TEXA

CDEX

CDMS

## Modulation:

DAMA/LIBRA

DM ICE

## Directional:

DRIFT

DM-TPC

Too Many Experiments, Too Little  
Time - My Apology for not  
Covering All

## er:

AMIC

EXT

## eated:

## Multi-purpose:

Majorana Demonstrator

COURE-0/COURE

PICASSO

COUPP

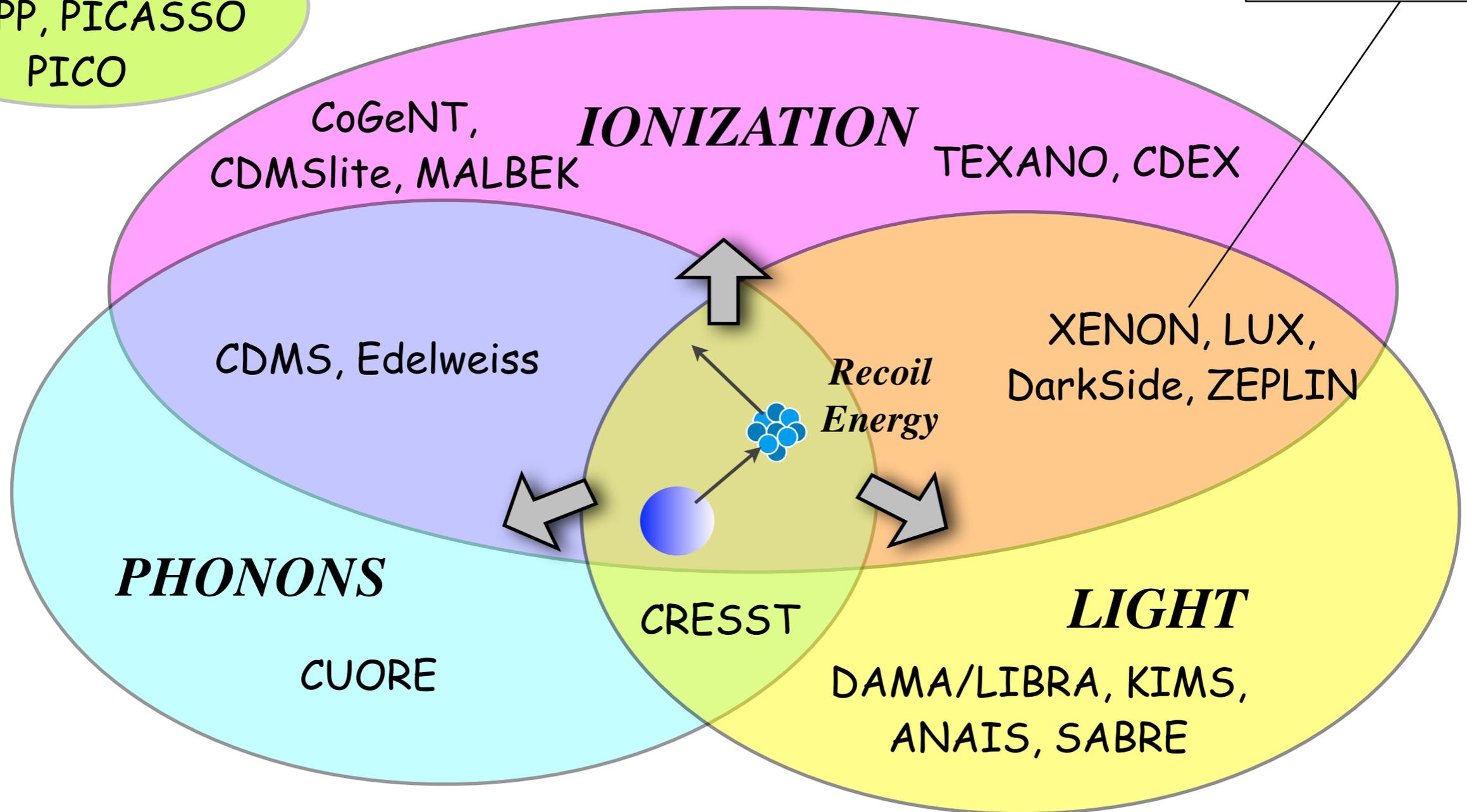
PICO-lite

\*Experiments in red are presenting results or status in parallel sessions.

# Direct Detection

*SuperHeated*  
COUPP, PICASSO  
PICO

Baudis - Tues.  
Plenary



# Minimize Backgrounds



Need at least 1000 m rock (~3000 mwe) overburden  
to reduce muon rate by  $\sim 10^5$

Nigel Smith - DM2012

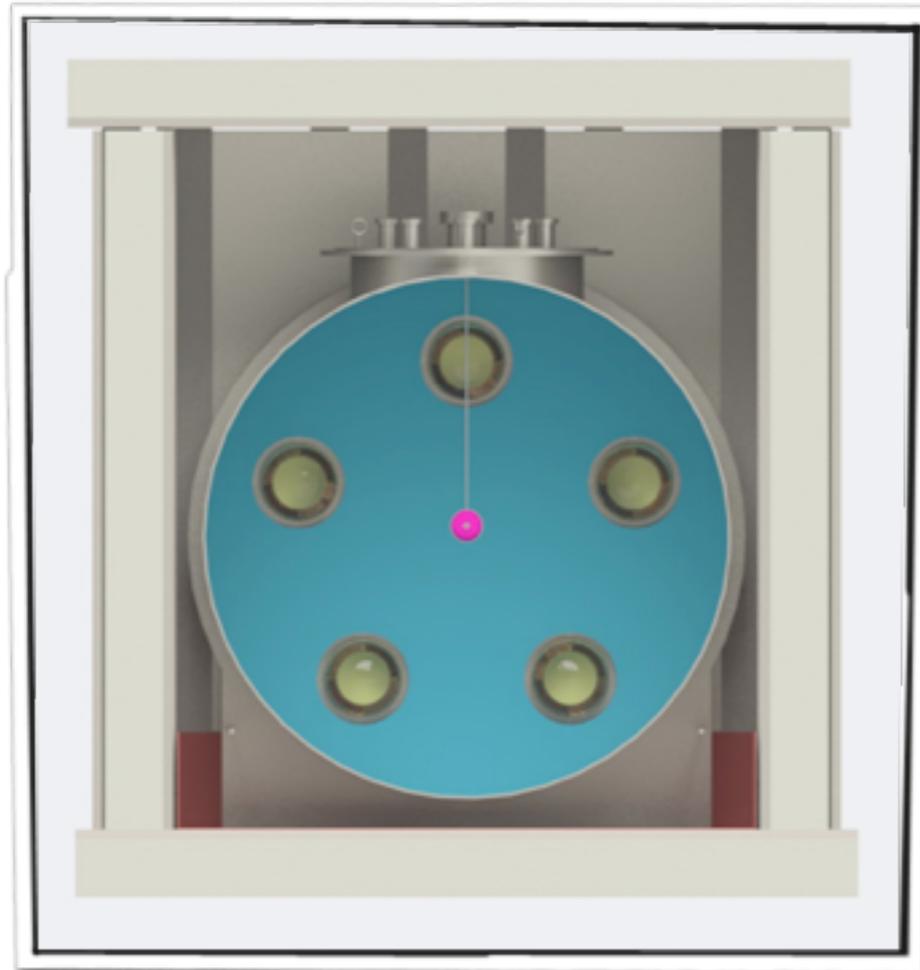
**Site experiments underground.**

# Minimize Backgrounds

## Active Muon Veto:

rejects events from cosmic rays

- Scintillating panels
- Water/Liquid Scintillator Shield



SABRE LAB shield design



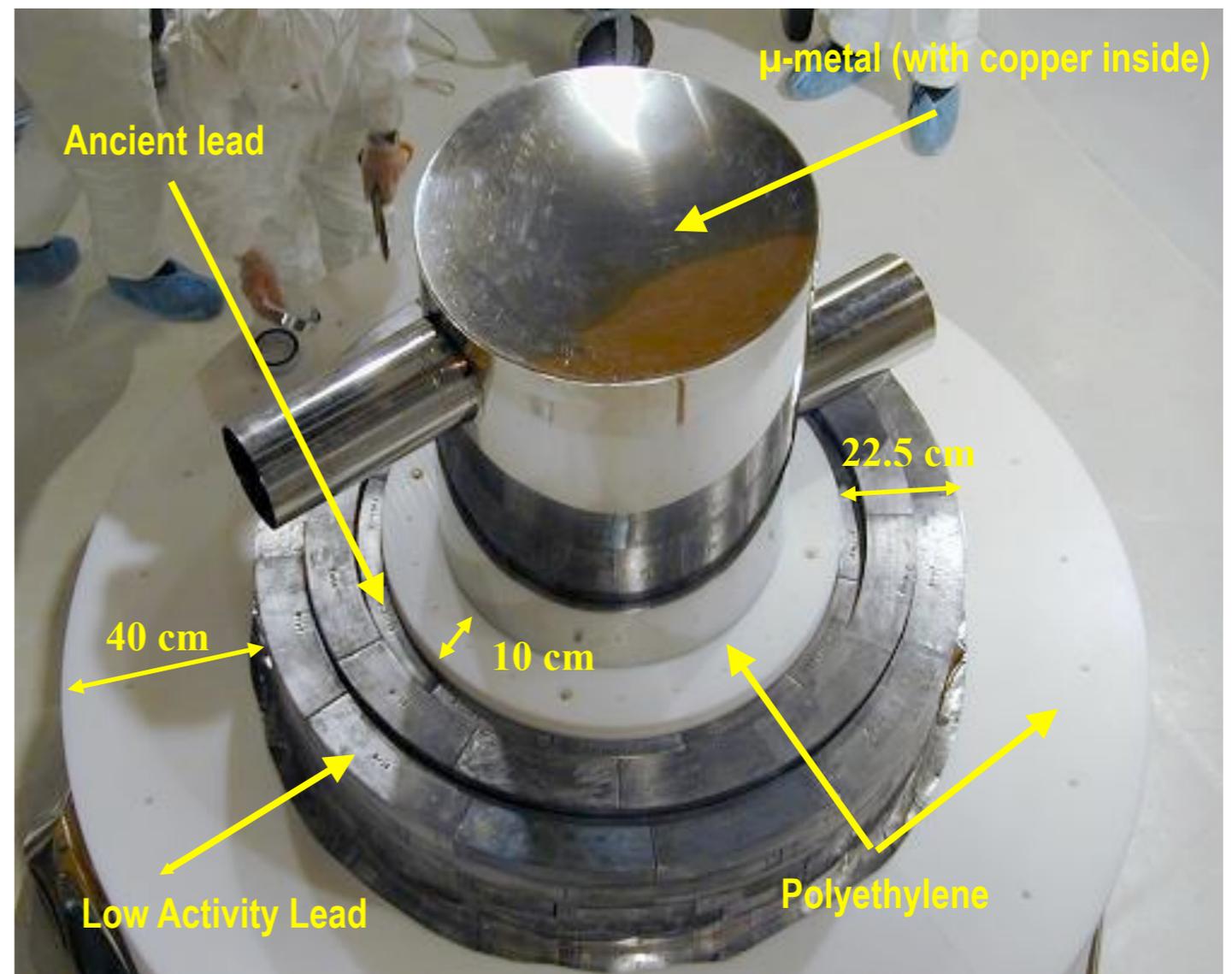
SCDMS active muon veto

# Minimize Backgrounds

## Use Passive Shielding

**Pb:** shielding from gammas resulting from radioactivity

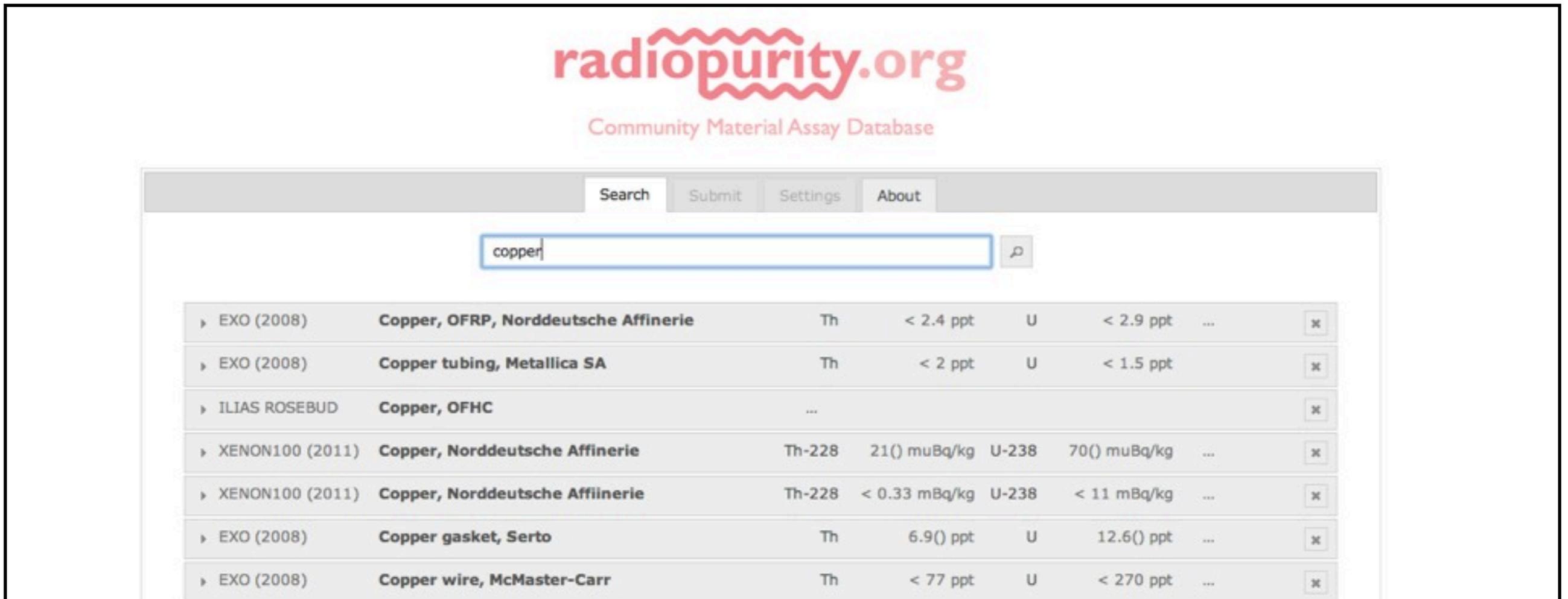
**Polyethylene:** moderate neutrons produced from fission decays and from  $(\alpha, n)$  interactions resulting from U/Th decays



SCDMS - Layers of Polyethylene and Lead

# Minimize Backgrounds

## Use Clean Materials



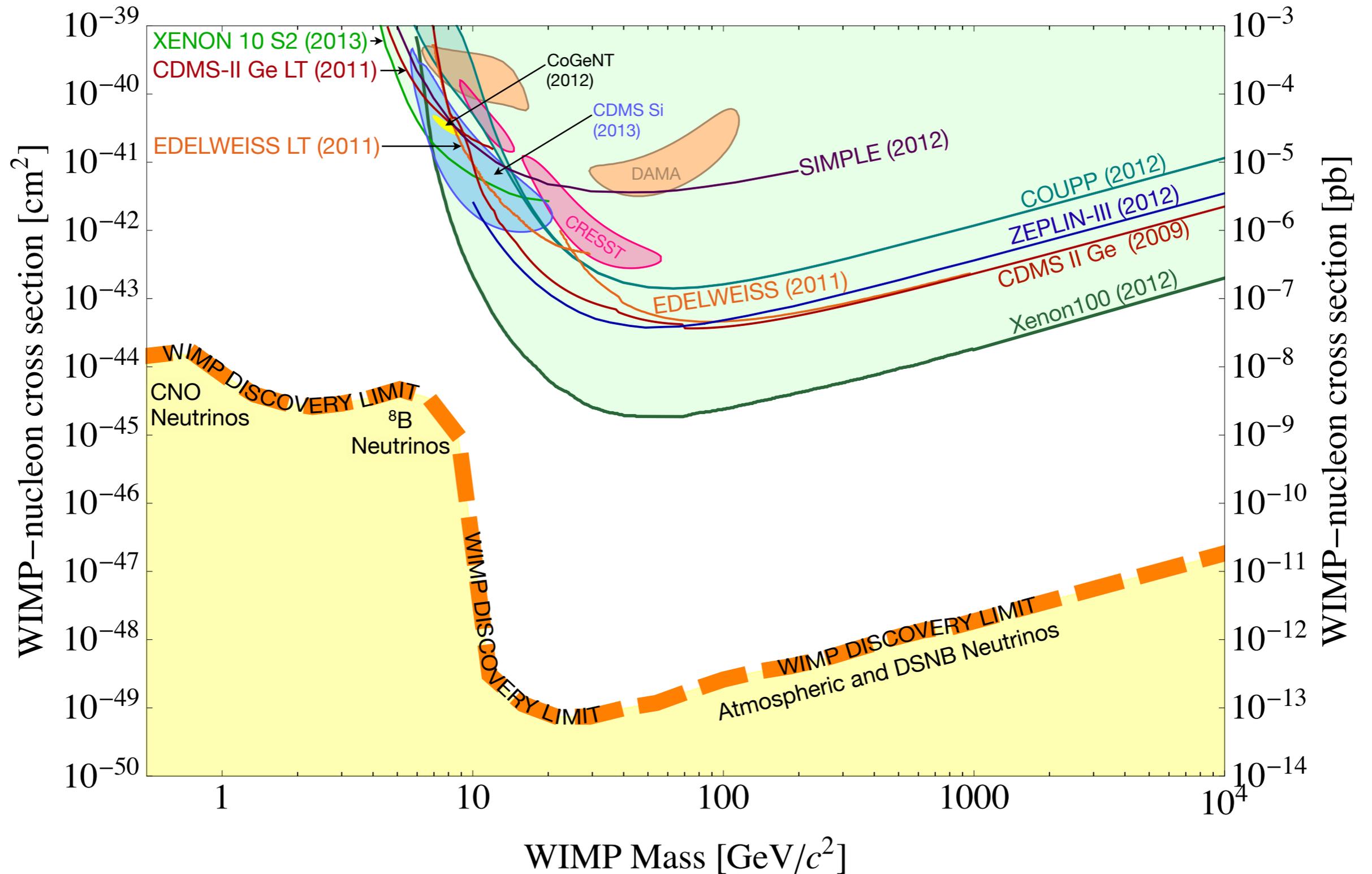
The screenshot shows the radiopurity.org website interface. At the top, the logo 'radiopurity.org' is displayed in red, with the tagline 'Community Material Assay Database' below it. A navigation bar contains 'Search', 'Submit', 'Settings', and 'About' buttons. A search input field contains the text 'copper'. Below the search bar, a table lists search results for various copper materials and their associated radionuclide levels.

Sample ID	Material	Radionuclide	Activity	Unit	Activity	Unit	...	Action
EXO (2008)	Copper, OFRP, Norddeutsche Affinerie	Th	< 2.4 ppt	U	< 2.9 ppt		...	✕
EXO (2008)	Copper tubing, Metallica SA	Th	< 2 ppt	U	< 1.5 ppt			✕
ILIAS ROSEBUD	Copper, OFHC	...						✕
XENON100 (2011)	Copper, Norddeutsche Affinerie	Th-228	21() muBq/kg	U-238	70() muBq/kg		...	✕
XENON100 (2011)	Copper, Norddeutsche Affinerie	Th-228	< 0.33 mBq/kg	U-238	< 11 mBq/kg		...	✕
EXO (2008)	Copper gasket, Serto	Th	6.9() ppt	U	12.6() ppt		...	✕
EXO (2008)	Copper wire, McMaster-Carr	Th	< 77 ppt	U	< 270 ppt		...	✕

<http://radiopurity.org>

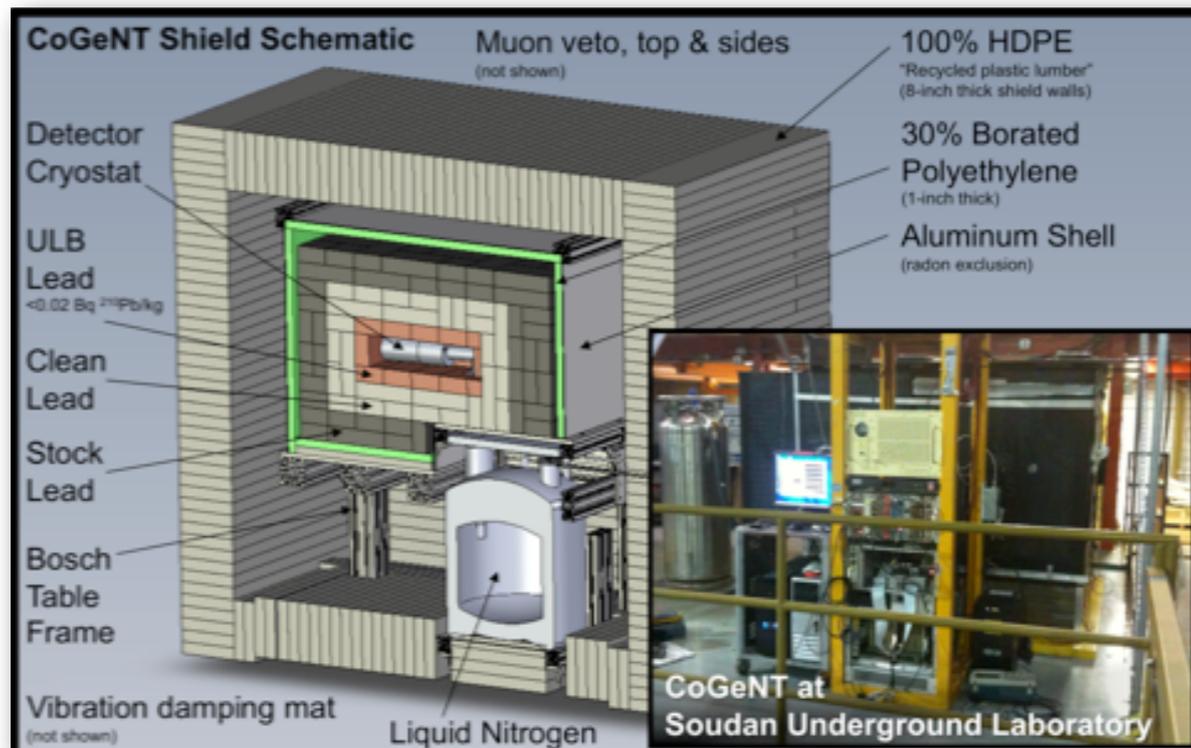
Supported by AARM, LBNL, MAJORANA, SMU, SJTU & others

# Where Are We Now?



The Experiments Part 1: The Low Mass Region  
Excesses Reported by DAMA/LIBRA,  
CoGeNT, CRESST and CDMS

# CoGeNT

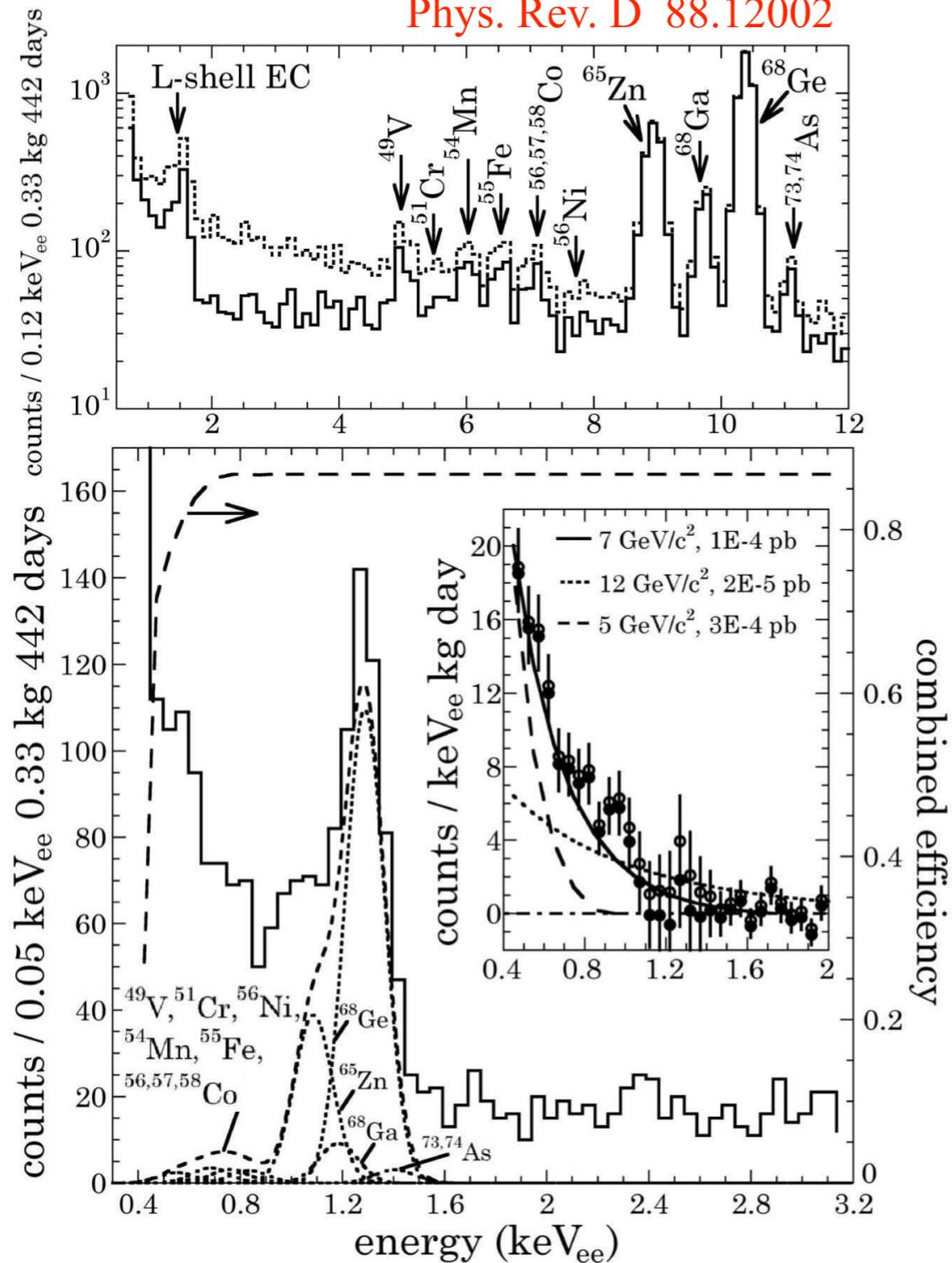


- Location: Soudan Underground Laboratory, Minnesota, USA
- 440 g HPGe ionization spectrometer
- Data collection from Dec. 4, 2009 - Mar. 6, 2011 (442 live days)
- Data collection interrupted due to fire.
- Data collection resumed July 2011.



# CoGeNT

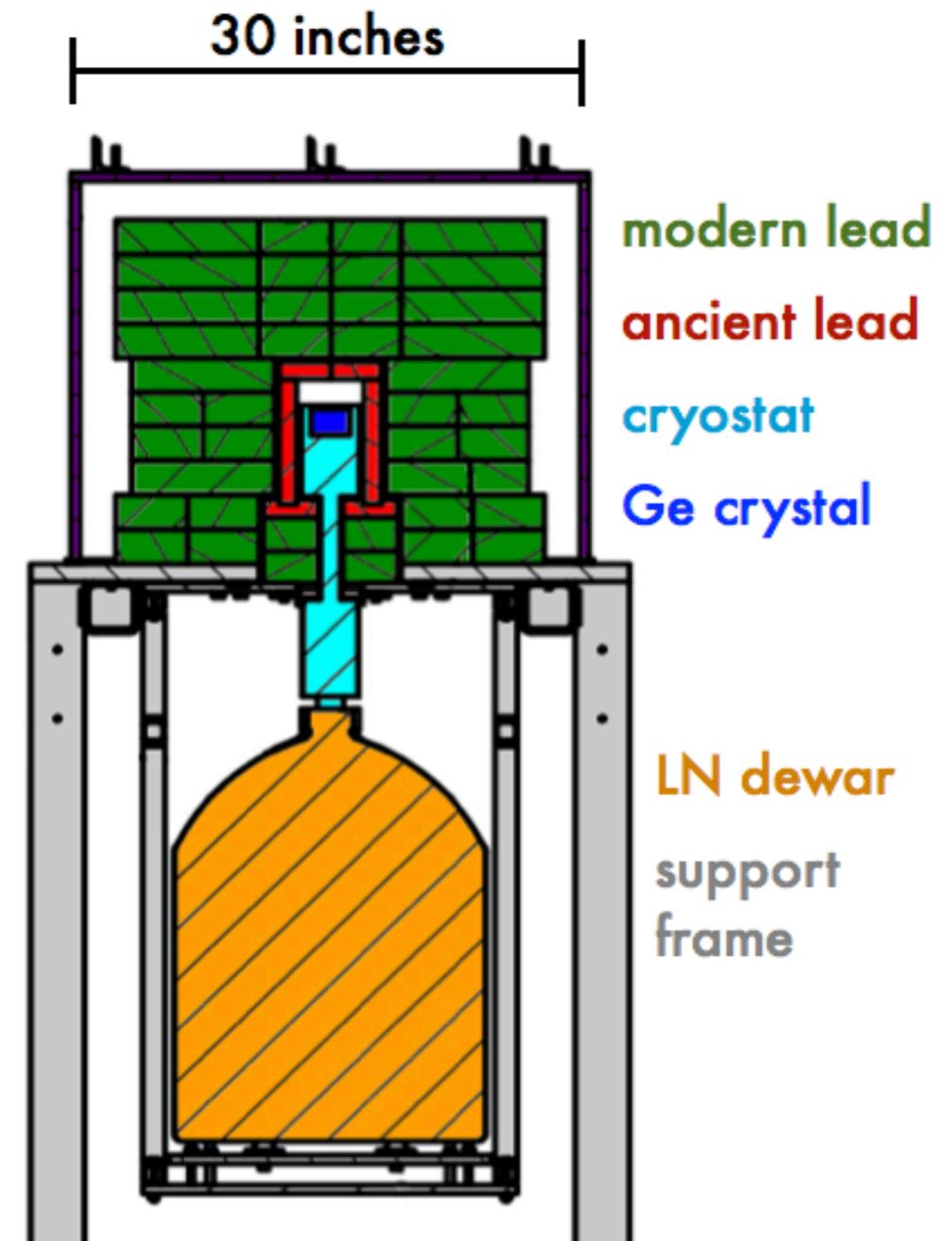
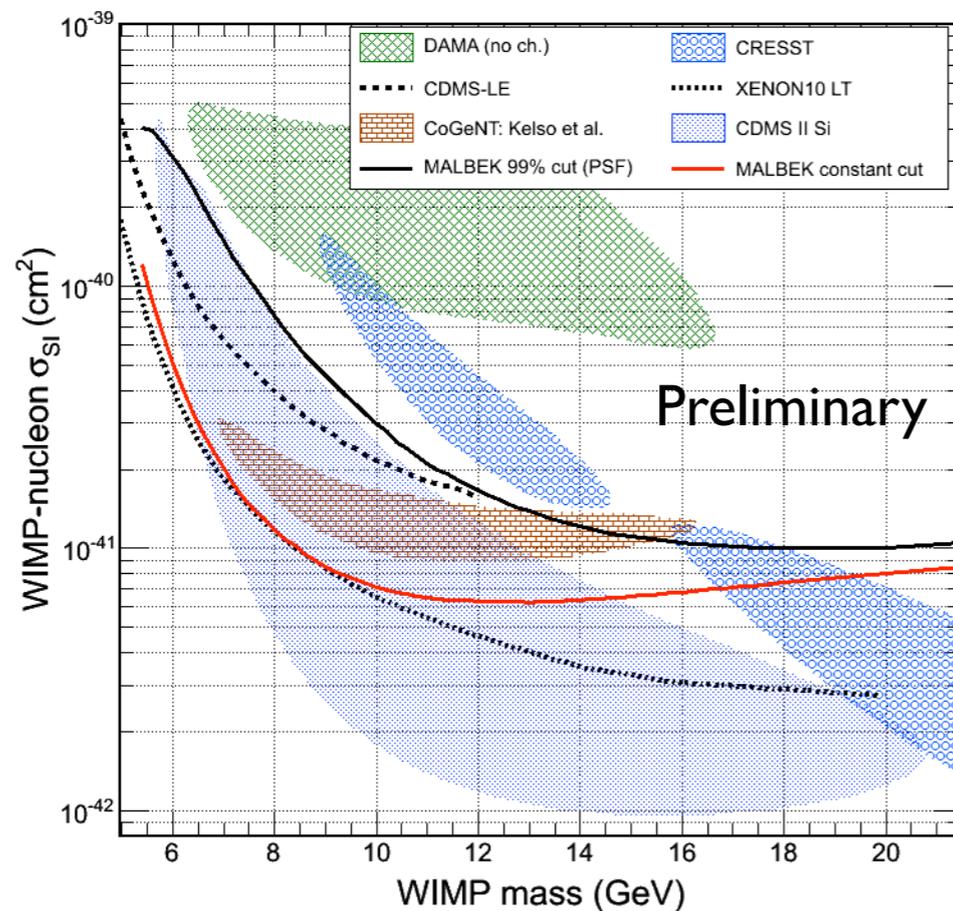
Phys. Rev. D 88.12002



- First claim of excess in 2010.
- Reject surface events using risetime cut (2011).
- Peaks due to cosmogenic activation of Ge
- After subtraction of known background, an exponential excess of events remains
- Fits to a variety of light-WIMP masses and couplings shown in inset of lower figure.
- Publication of new data coming soon.

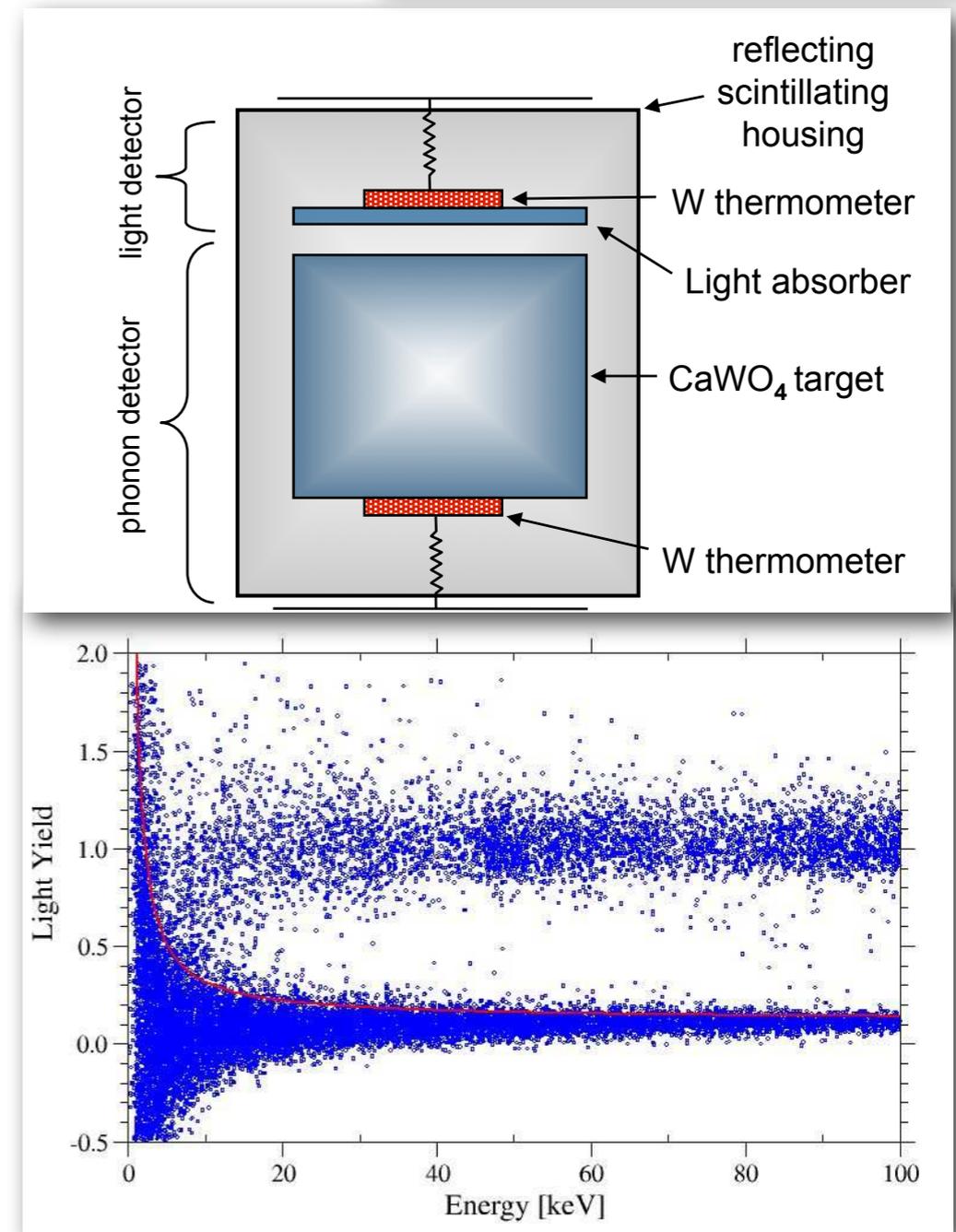
# MALBEK

- **MAJORANA** Low-background **BE**Ge detector at **KURF**.
- 450g Canberra Broad Energy Ge (BEGe) detector with ultra-low background components provided by J.I. Collar.
- Location: Kimballton Underground Research Facility (KURF), VA at 1450 mwe.
- 90% exclusion limits from 221 day data run.



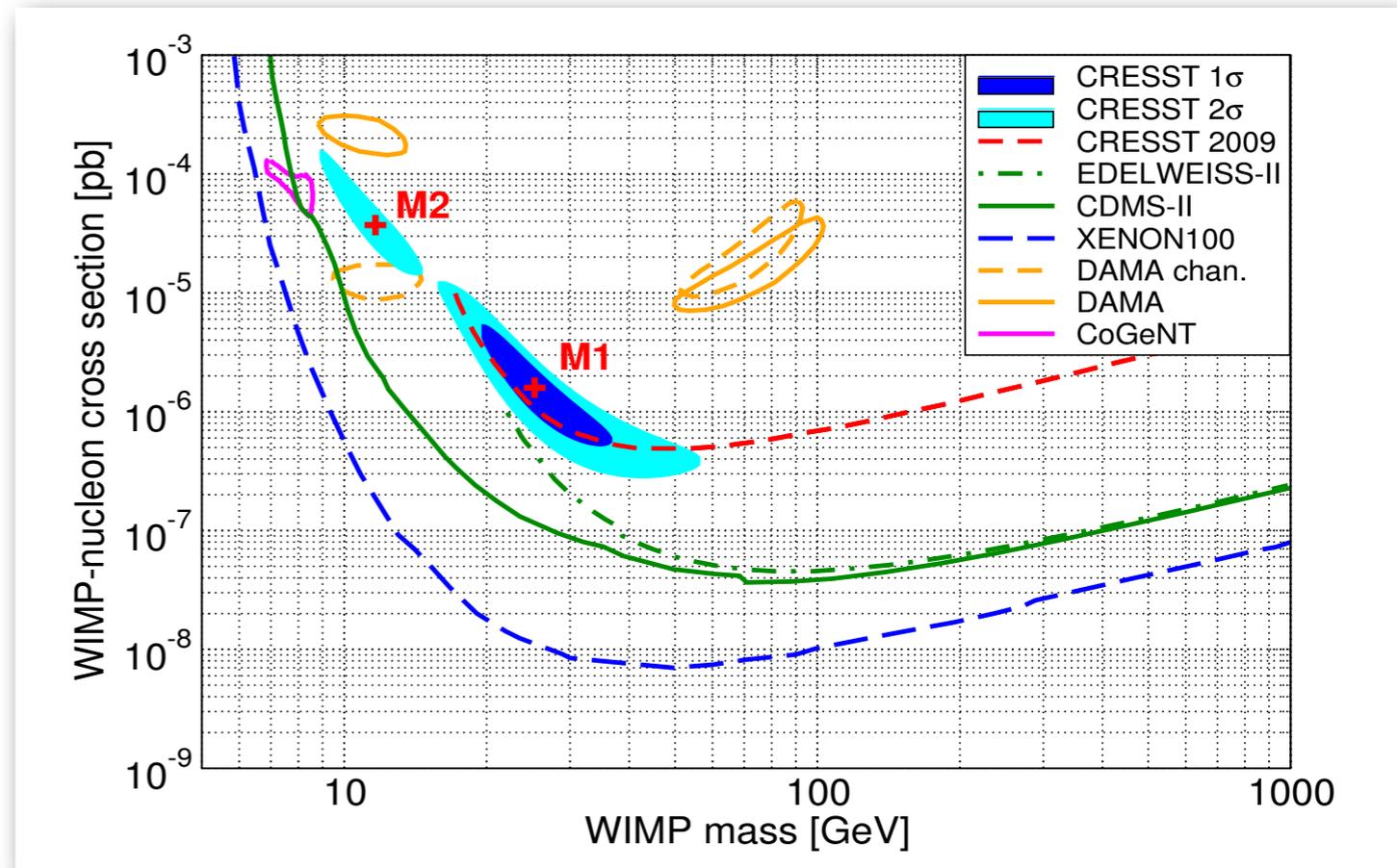
# CRESST

- Cryogenic  $\text{CaWO}_4$  crystals ( $\sim 300$  g each) are instrumented to readout phonon energy and scintillation.
- Location: Laboratori Nazionali del Gran Sasso, Italy
- Discrimination between ER and NR events via light yield (light/phonon energy)
- Net exposure: 730 kg-day (July 2009 - March 2011) from 8 detector modules.
- Observed 67 events in acceptance region (orange). [arXiv:1109.0702](https://arxiv.org/abs/1109.0702)
  - Analysis used a maximum likelihood in which 2 regions favored a WIMP signal in addition to predict background.
  - Excess events can not be explained by known backgrounds
  - Large background contribution



# CRESST Plans

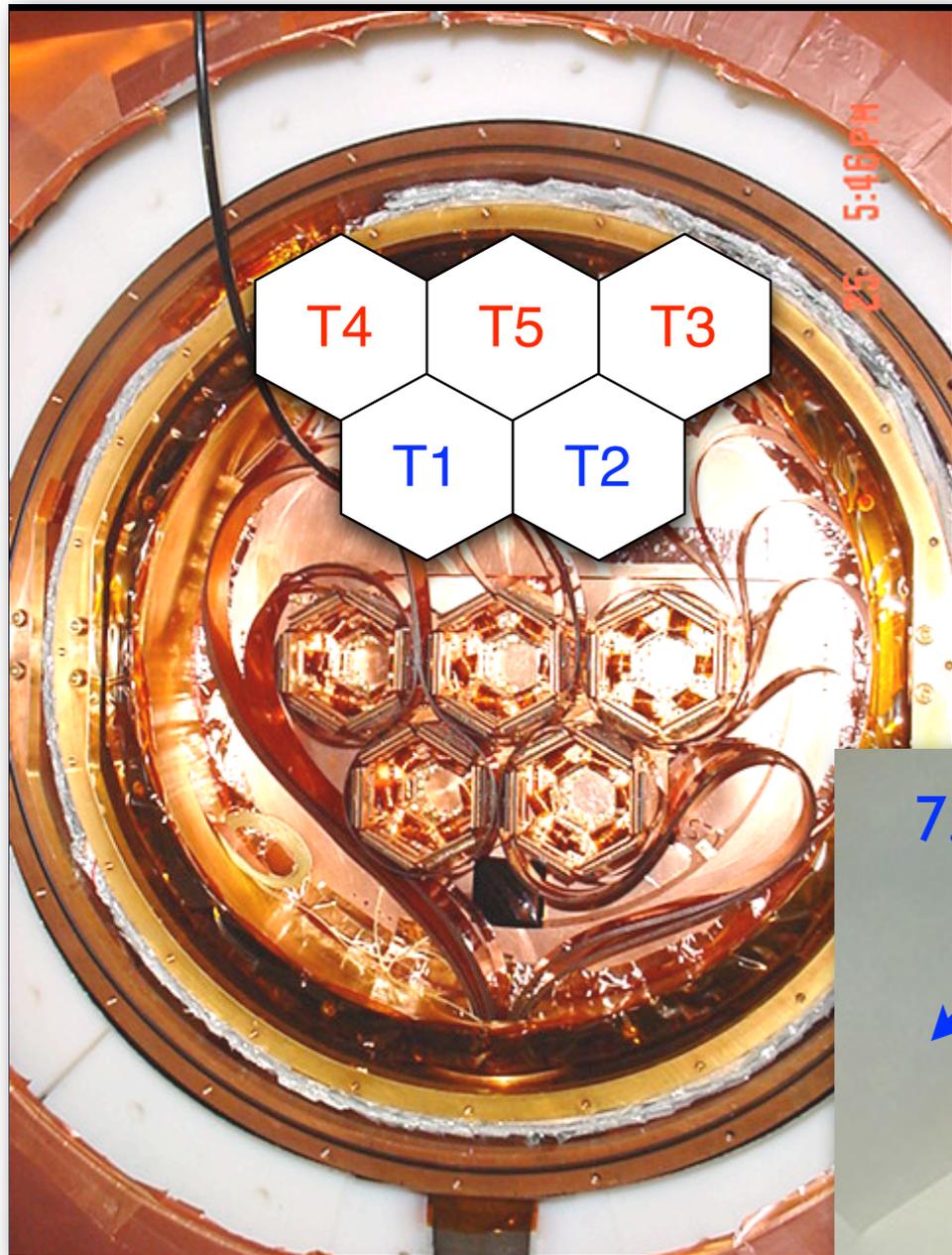
- Current data run aims to reduce background, increase detector mass.
  - Alphas - new clamping design and material
  - Detector assembly in a radon free environment
  - New detector design to discriminate  $^{206}\text{Po}$  recoils
  - Add additional shielding to reduce neutron background
- June & July calibration runs with  $^{57}\text{Co}$  source were successful.
- **July 30th, 2013 Science Runs Begin!**



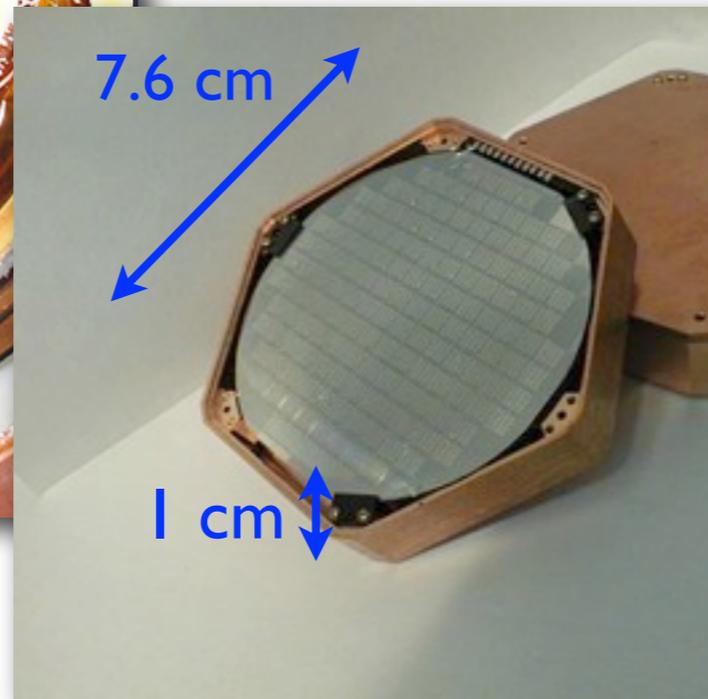
[arXiv:1109.0702](https://arxiv.org/abs/1109.0702)

# CDMS II

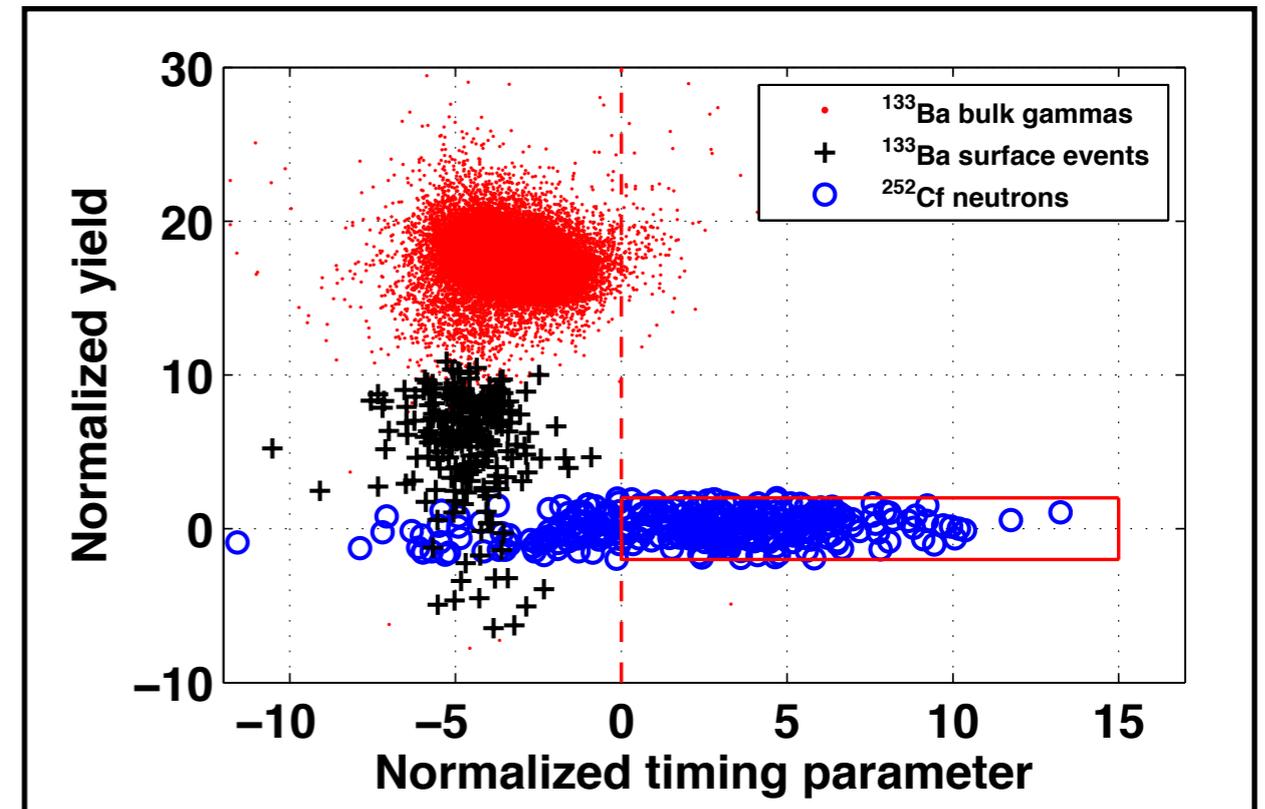
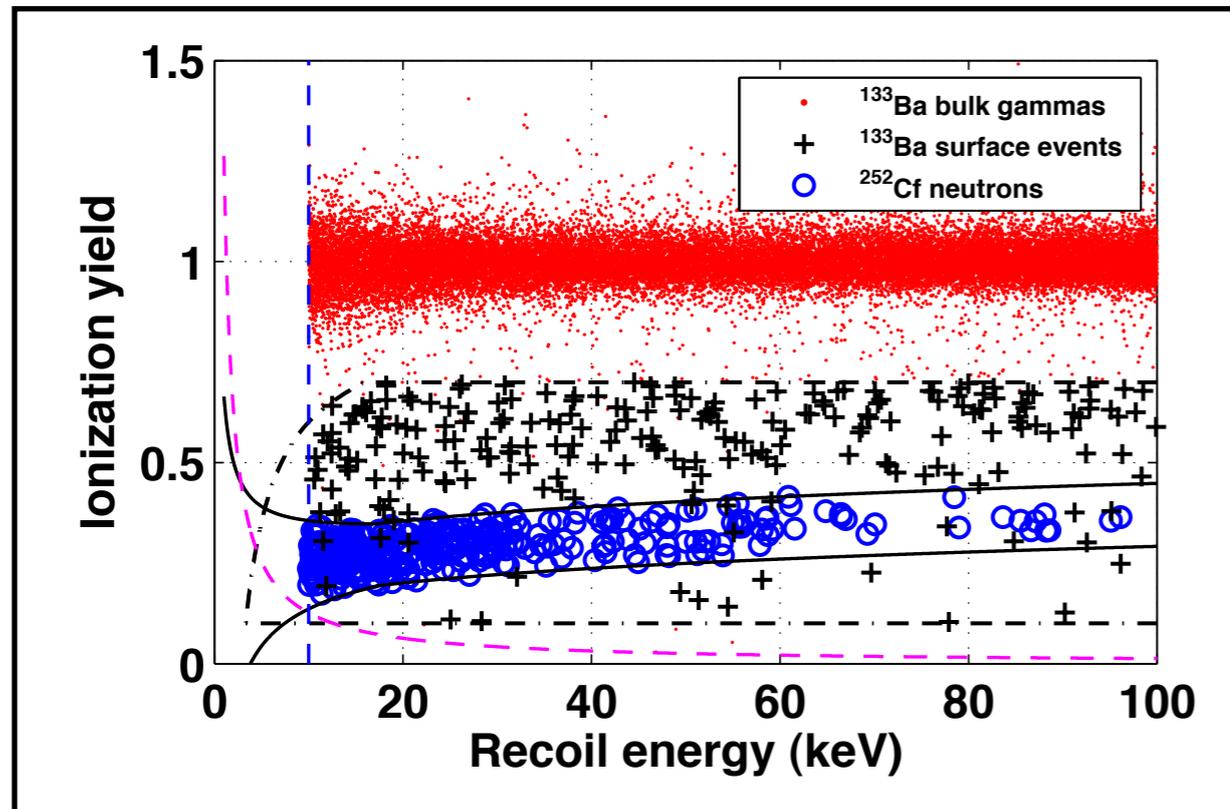
Billard - Mon. DM I  
Nelson - Mon. DM II  
Speller - Poster



- CDMS II: 30 detectors (19 Ge, 11 Si) installed and operated in the Soudan Underground Laboratory, MN, USA from Jun. 06 - Mar. 09.
- Measures ionization and phonons (read out by TES)
- Science Results: CDMS-Ge, CDMS-Si, Ge-Low Threshold, Annual Modulation

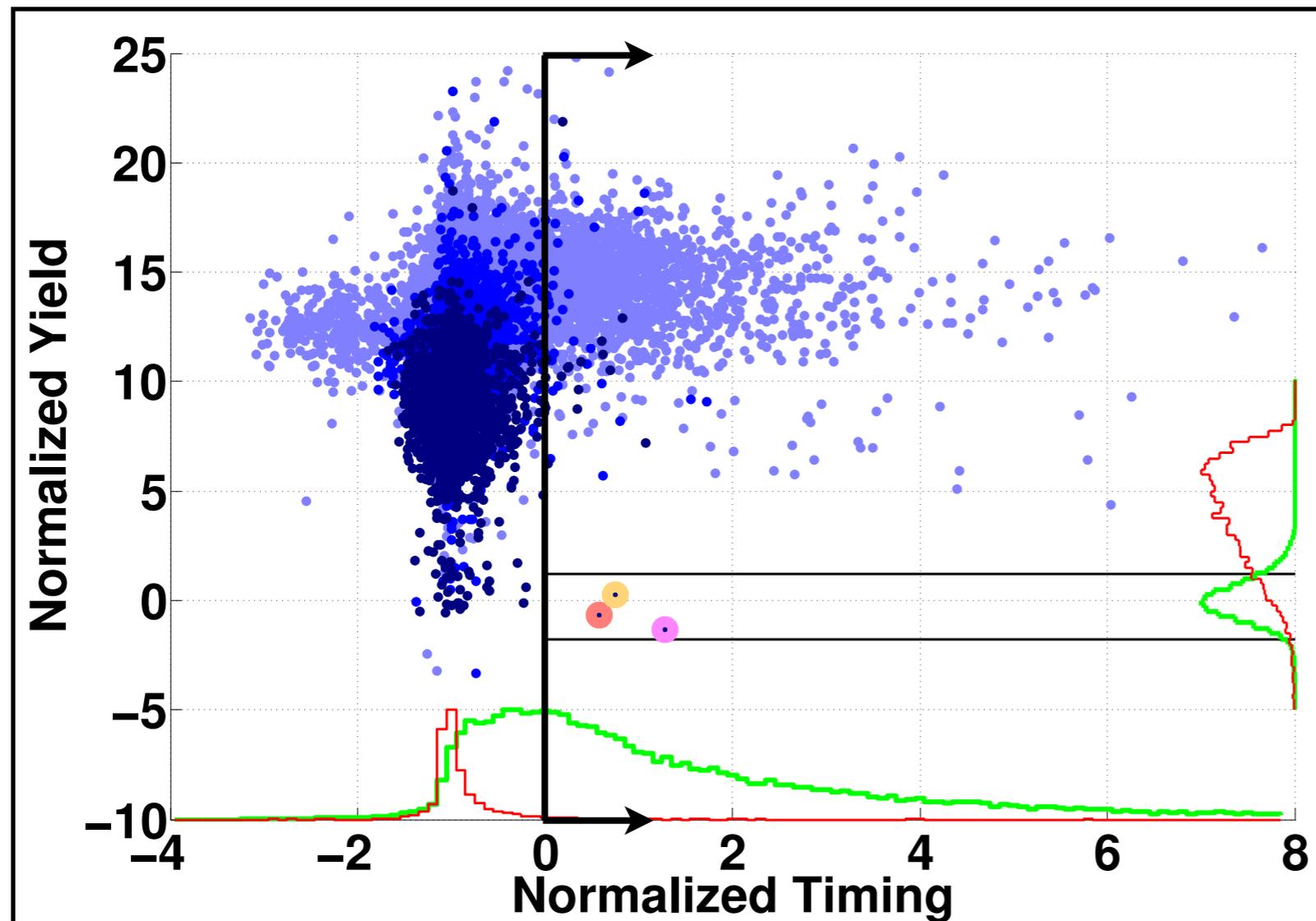


# Background Rejection



- Most backgrounds produce electron recoils and have yield (ionization/phonon energy)  $\sim 1$ .
- WIMPs and neutrons produce nuclear recoils and have yield  $\sim 0.3$ .
- Surface events can be identified using timing properties of the phonon and charge pulses.

# Recent Results: CDMS II-Si Detectors



Observed 3 events.

- Shades of blue indicate three separate timing cut energy ranges.

- 7- 20 keV

- 20 - 30 keV

- 30 - 100 keV

- Background Estimate

- Surface Events

$$0.41^{+0.20}_{-0.08}(\text{stat.})^{+0.28}_{-0.24}(\text{syst.})$$

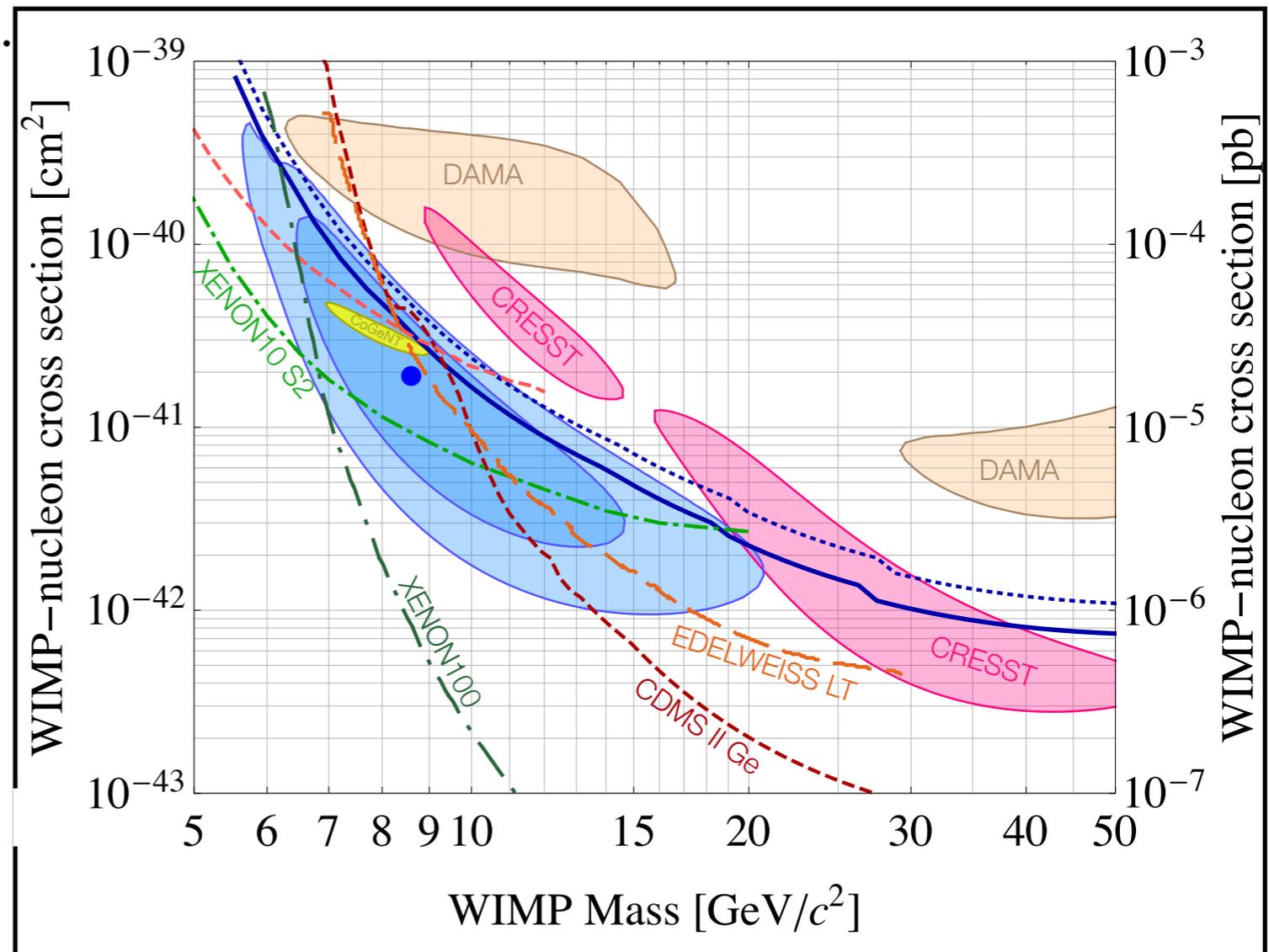
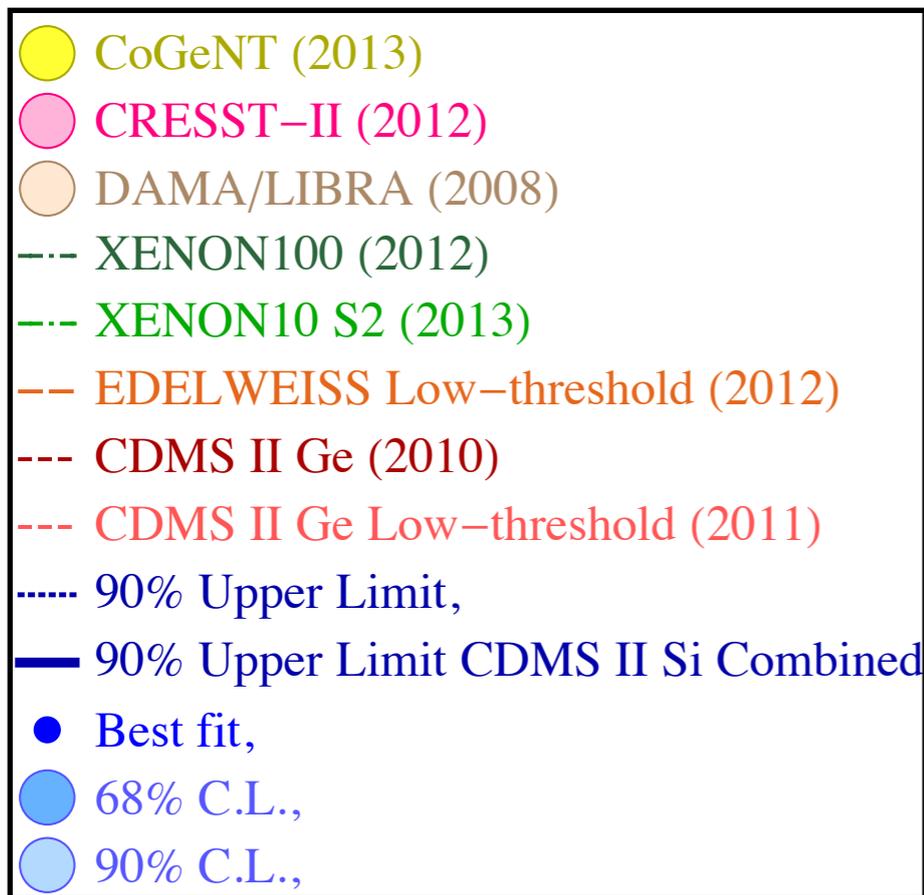
- < 0.13 neutrons from  
Cosmogenics &  
Radiogenics

- < 0.08  $^{206}\text{Pb}$  recoils from  
 $^{210}\text{Pb}$  decays

# CDMS II - Si Results

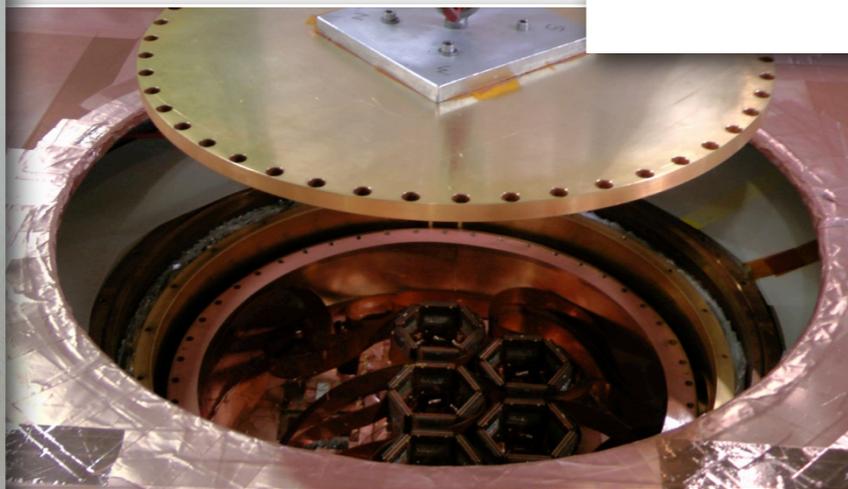
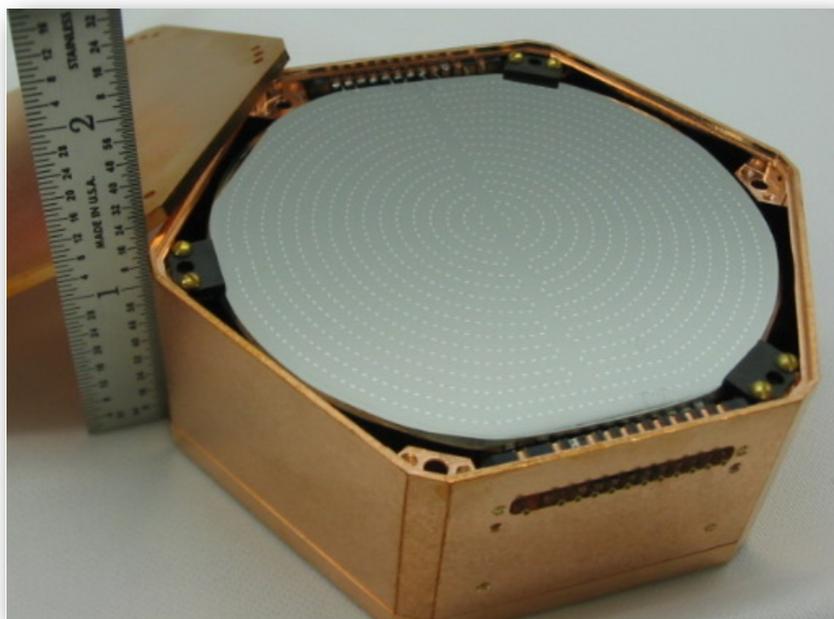
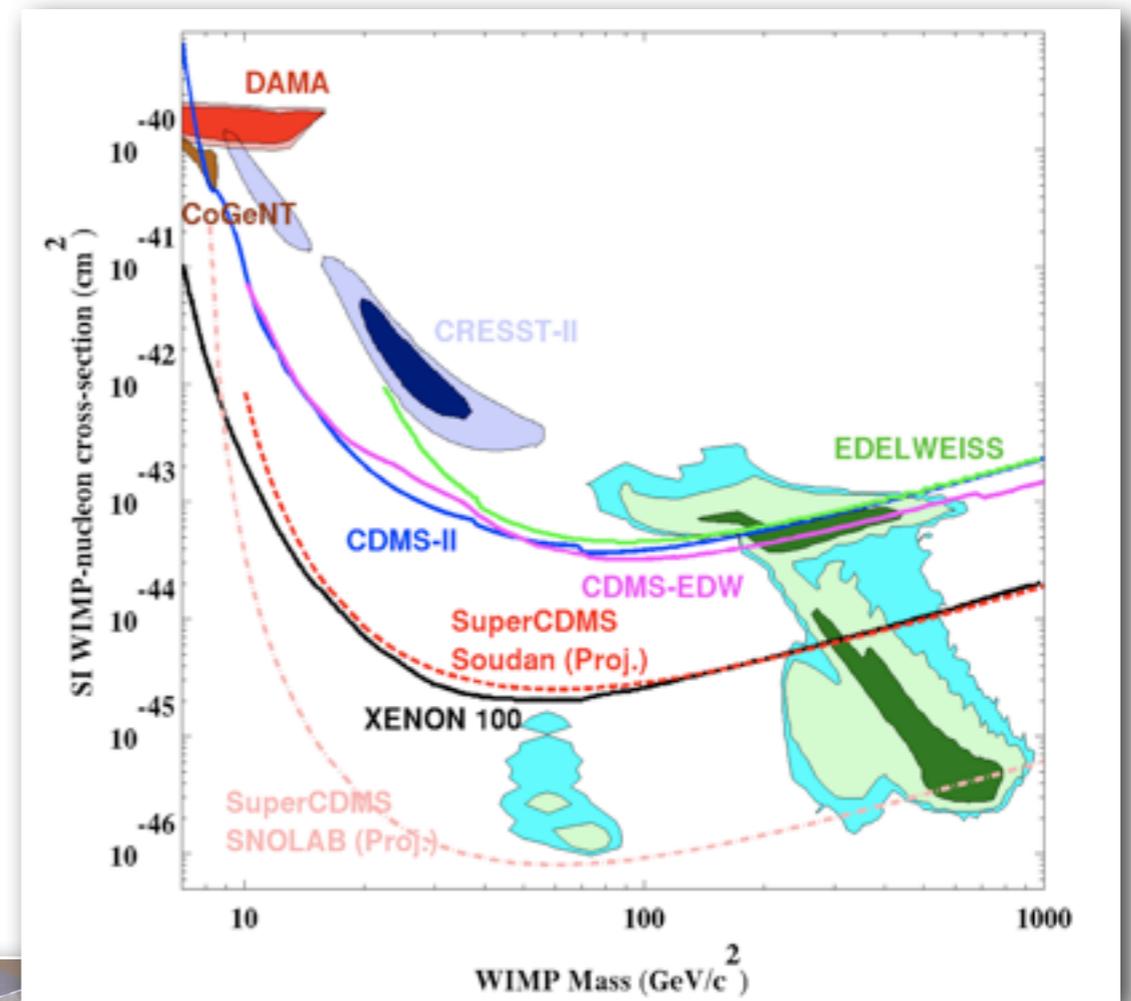
- Three events observed in the signal region.
- A profile likelihood analysis favors a WIMP+background hypothesis over the known background estimate as the source of our signal at the 99.81% C.L. ( $\sim 3\sigma$ , p-value: 0.19%)

- The maximum likelihood occurs at a WIMP mass of 8.6 GeV/c<sup>2</sup> and WIMP-nucleon cross section of  $1.9 \times 10^{-41}$ .
- Does not rise to level of discovery, but does call for further investigation.



# SuperCDMS @ Soudan

- Currently operating 5 towers of advanced iZIP detectors (~9 kg Ge) in the existing cryostat at the Soudan Underground Laboratory.
- After 3 years of operation, expected to improve sensitivity to spin-independent WIMP-nucleon interactions by a factor of ~10 over existing CDMS II results.



Installation complete Nov. 8, 2011. Operating with final detector settings since Mar. 2012.

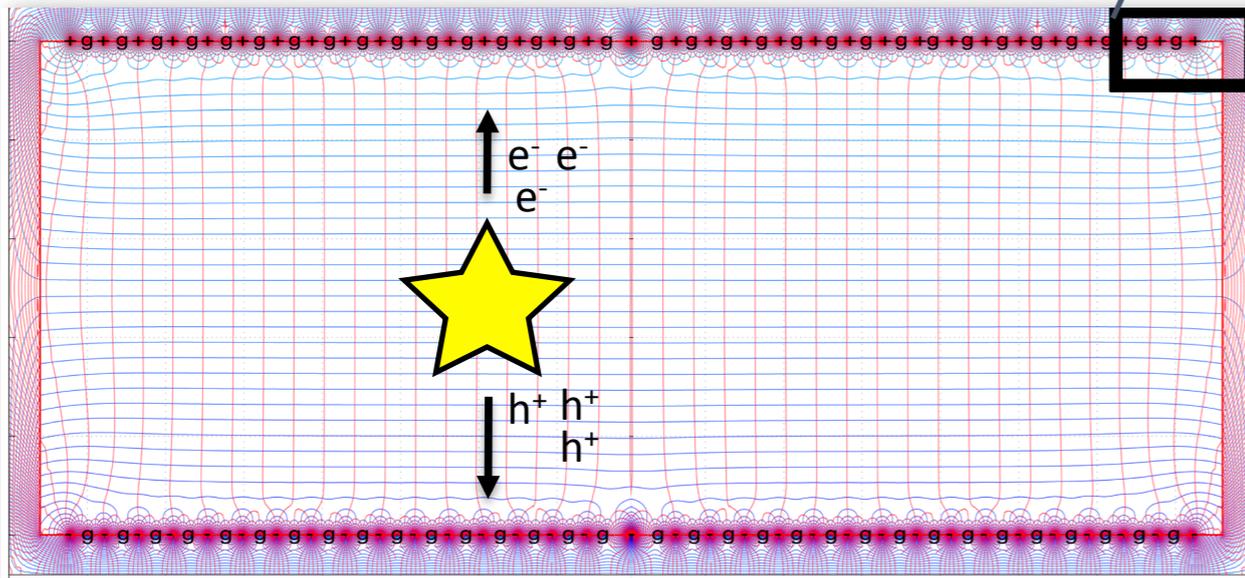
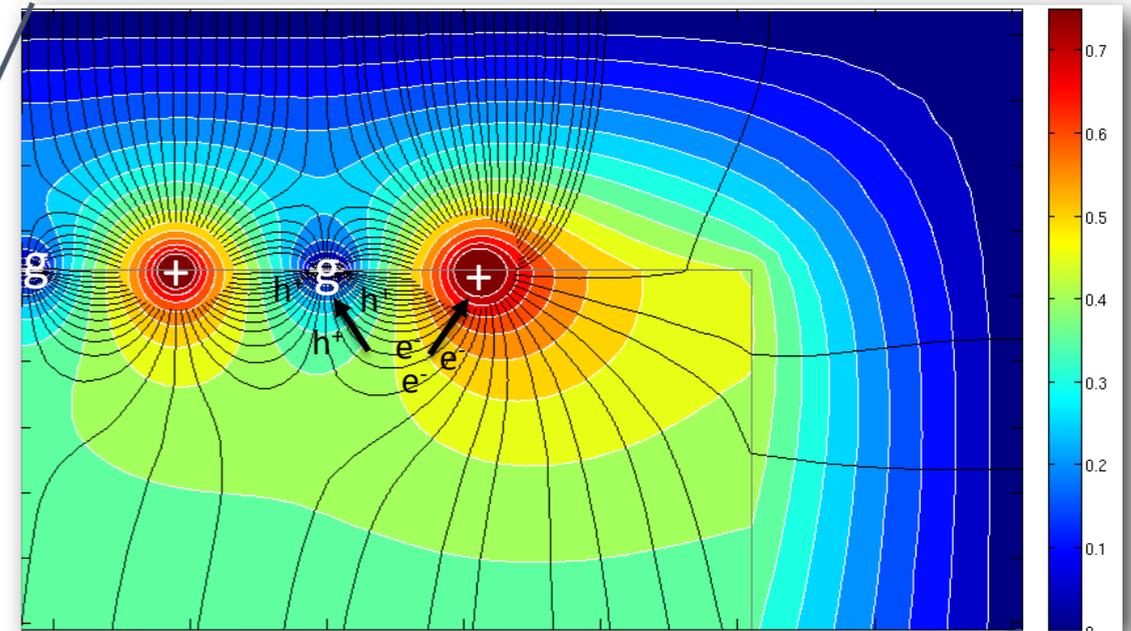
# SCDMS iZIPs: Charge Signal

## Bulk Events:

Equal but opposite ionization signal appears on both faces of detector (symmetric)

## Surface Events:

Ionization signal appears on one detector face (asymmetric)



arXiv:1305.2405

# SCDMS iZIPs: Charge Signal

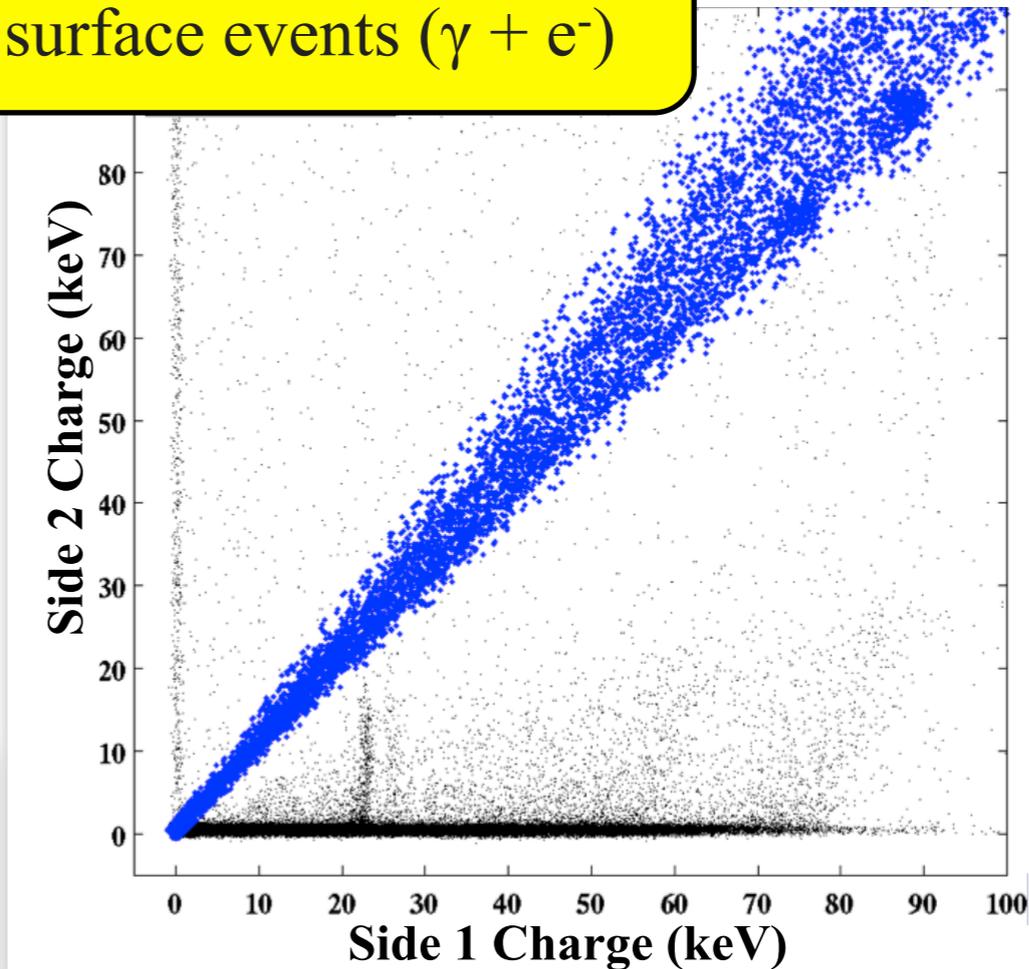
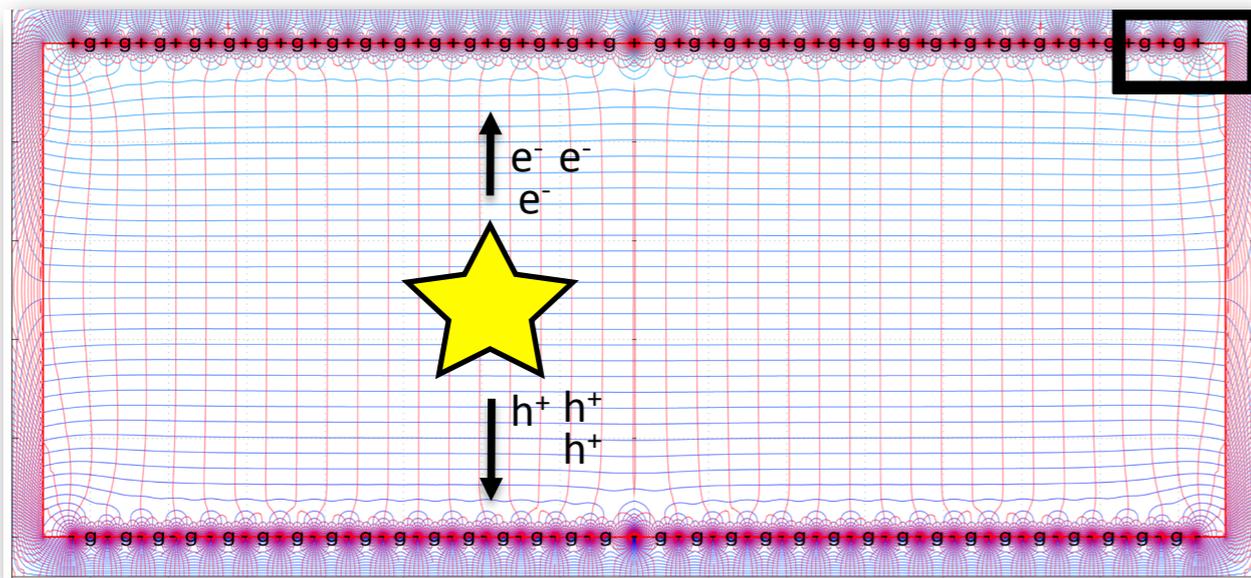
- bulk events ( $\gamma$ )
- surface events ( $\gamma + e^-$ )

## Bulk Events:

Equal but opposite ionization signal appears on both faces of detector (symmetric)

## Surface Events:

Ionization signal appears on one detector face (asymmetric)



- ~50% fiducial volume (8-115 keVr)
- < 0.6 events in 0.3 ton-years
- **Good enough for a 200 kg experiment run for 4 years at SNOLAB!**

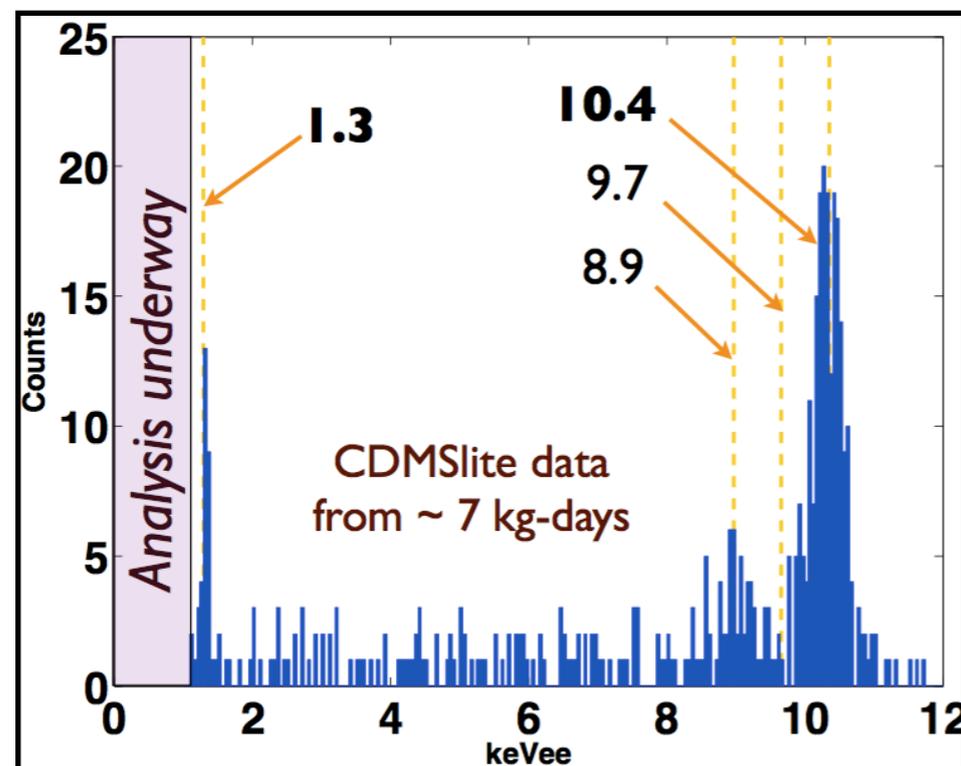
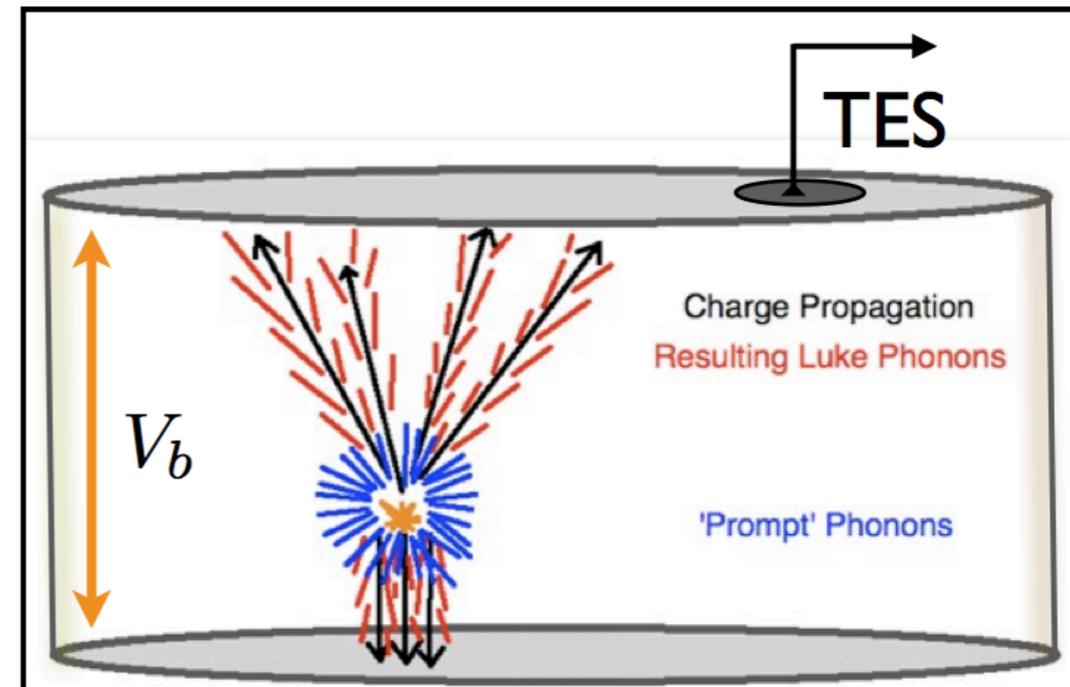
arXiv:1305.2405

# CDMSlite

- Alternate running mode to explore low mass WIMPs utilizing Luke phonons

$$E_{\text{luke}} = N_{e/h} \times eV_b$$

- Luke energy scales as bias voltage and noise remains constant until breakdown

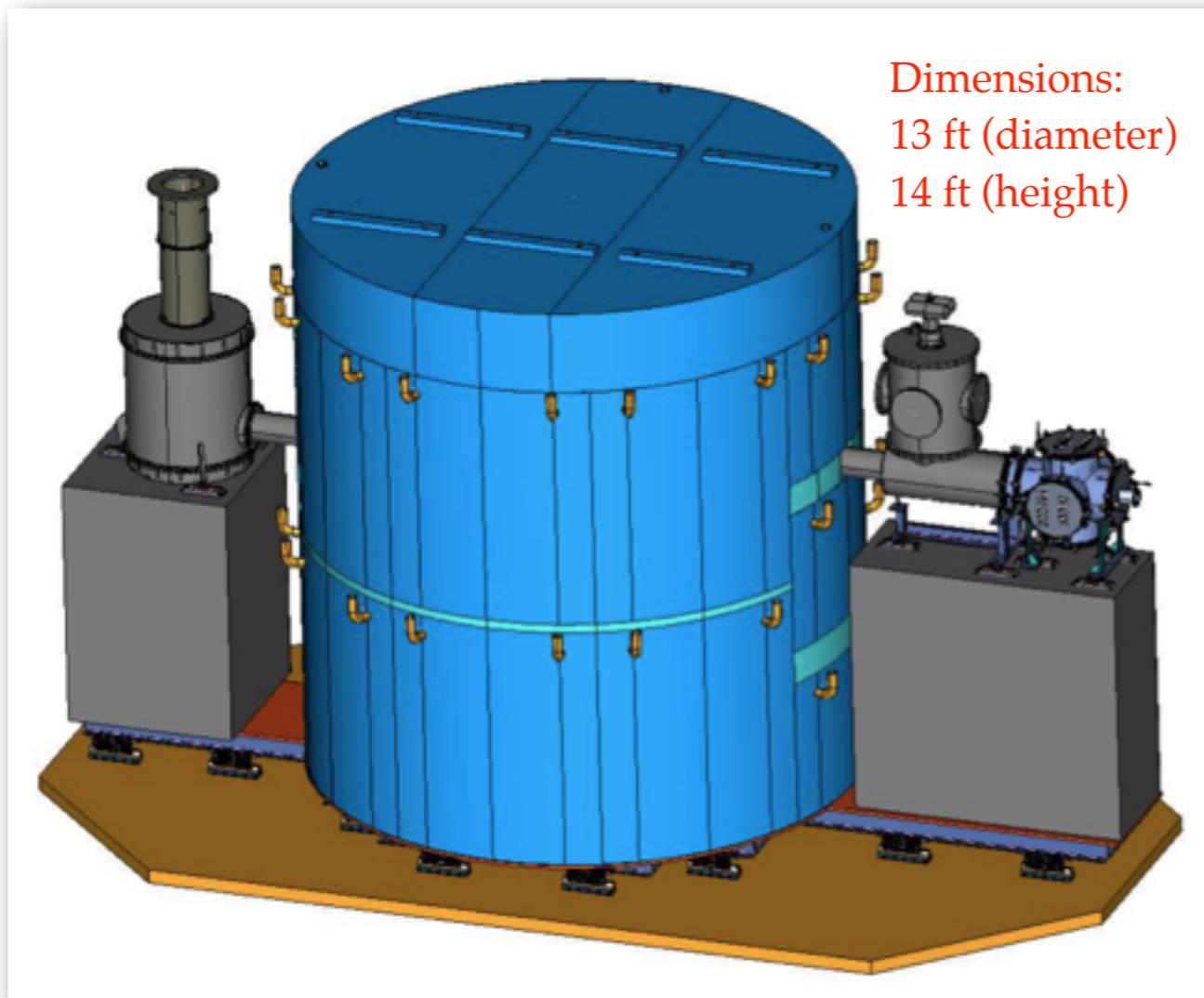


- Resulting Luke amplification has excellent energy resolution potentially down to 1.3  $eV_{ee}$ .
- Resolution of various Ge activation lines.

**New Results to be Announced Wednesday!**

# Future: SuperCDMS @ SNOLAB

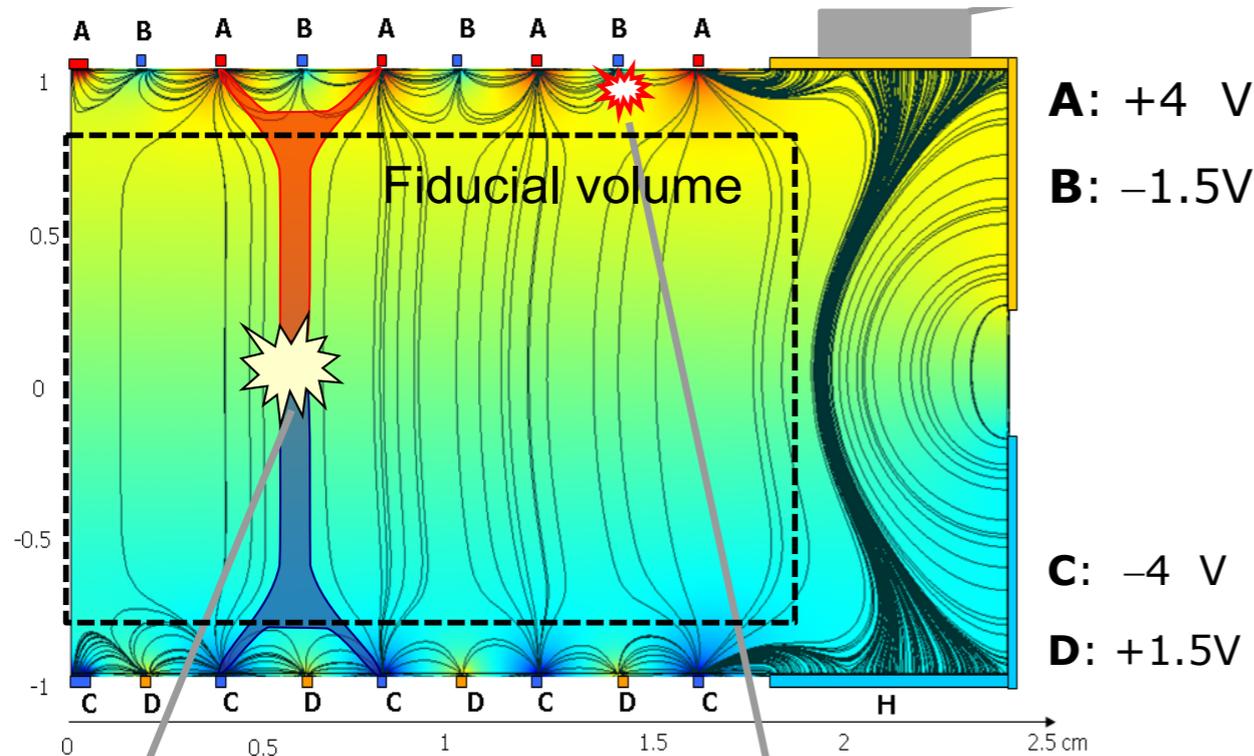
## Planned Setup



- cryostat volume of up to 400 kg target
  - 200 kg experiment with sensitivity of  $8 \times 10^{-47} \text{ cm}^2$  at  $60 \text{ GeV}/c^2$
  - Pb/Cu shielding for external radiation
  - increased PE shielding (neutrons)
  - possible neutron veto
- Calibration runs at Soudan indicate that the new iZIPs have good enough surface rejection capabilities for a 200 kg experiment at SNOLAB to run 4 years! ([arXiv:1305.2405](https://arxiv.org/abs/1305.2405))

# EDELWEISS III

NTD thermal sensor



**A:** +4 V

**B:** -1.5V

**C:** -4 V

**D:** +1.5V

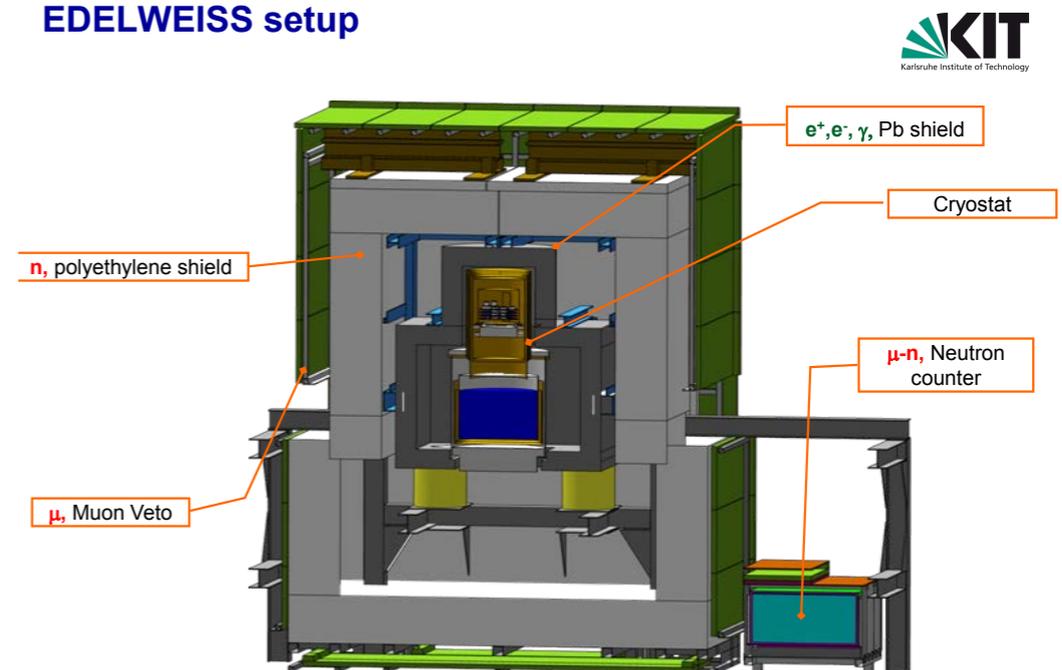
Bulk/Fiducial event  
Charge collected on  
electrodes A&C

Surface event  
Charge collected on  
electrodes A&B

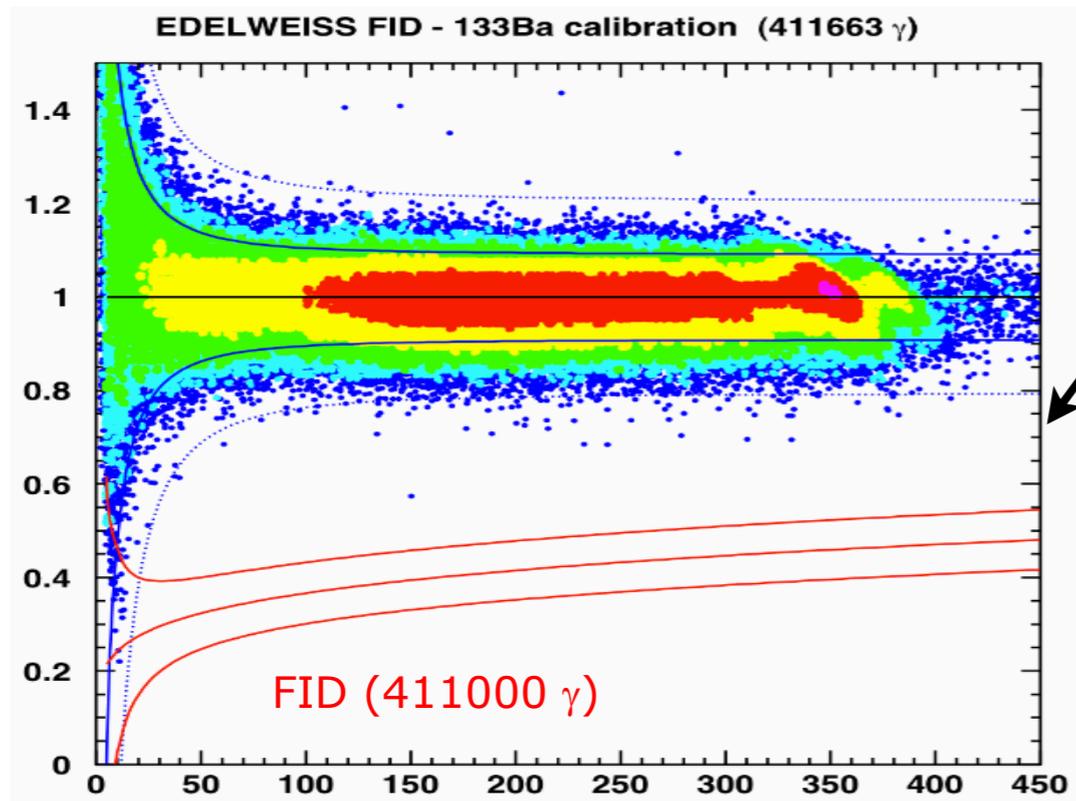
- Discrimination from ionization  
yield and charge collection  
symmetry.

- Located in the Laboratoire Souterrain de Modane (LSM) between Italy and France.
- Detectors instrumented with electrodes to measure charge and NTD thermal sensors to measure phonon signal.

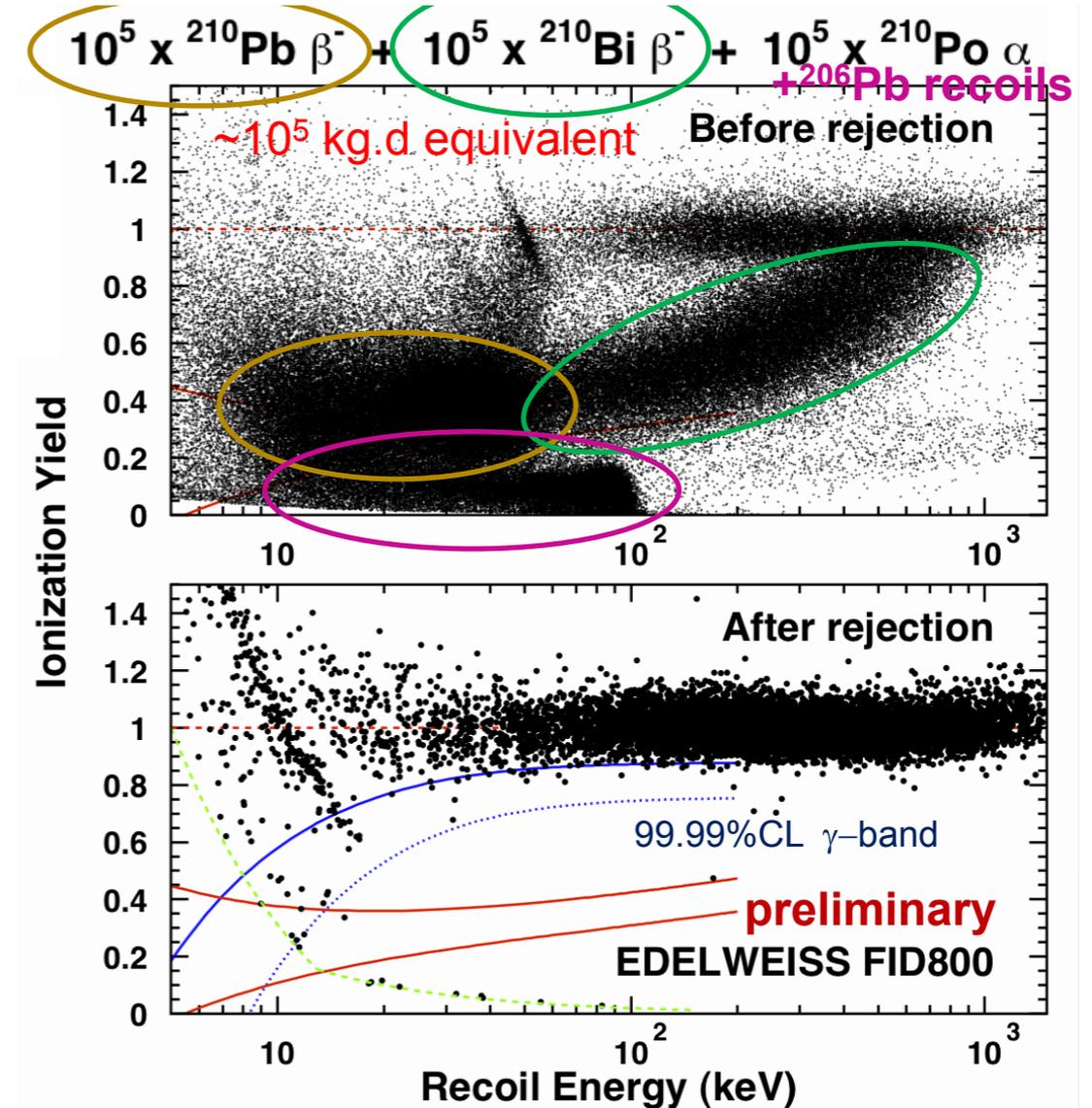
EDELWEISS setup



# EDELWEISS III



-Improvements to  $\gamma$ -discrimination



-Improvements to surface event discrimination

- $< 4 \times 10^{-5}$  misID events per kg-d ( $E_r > 15$  keV)

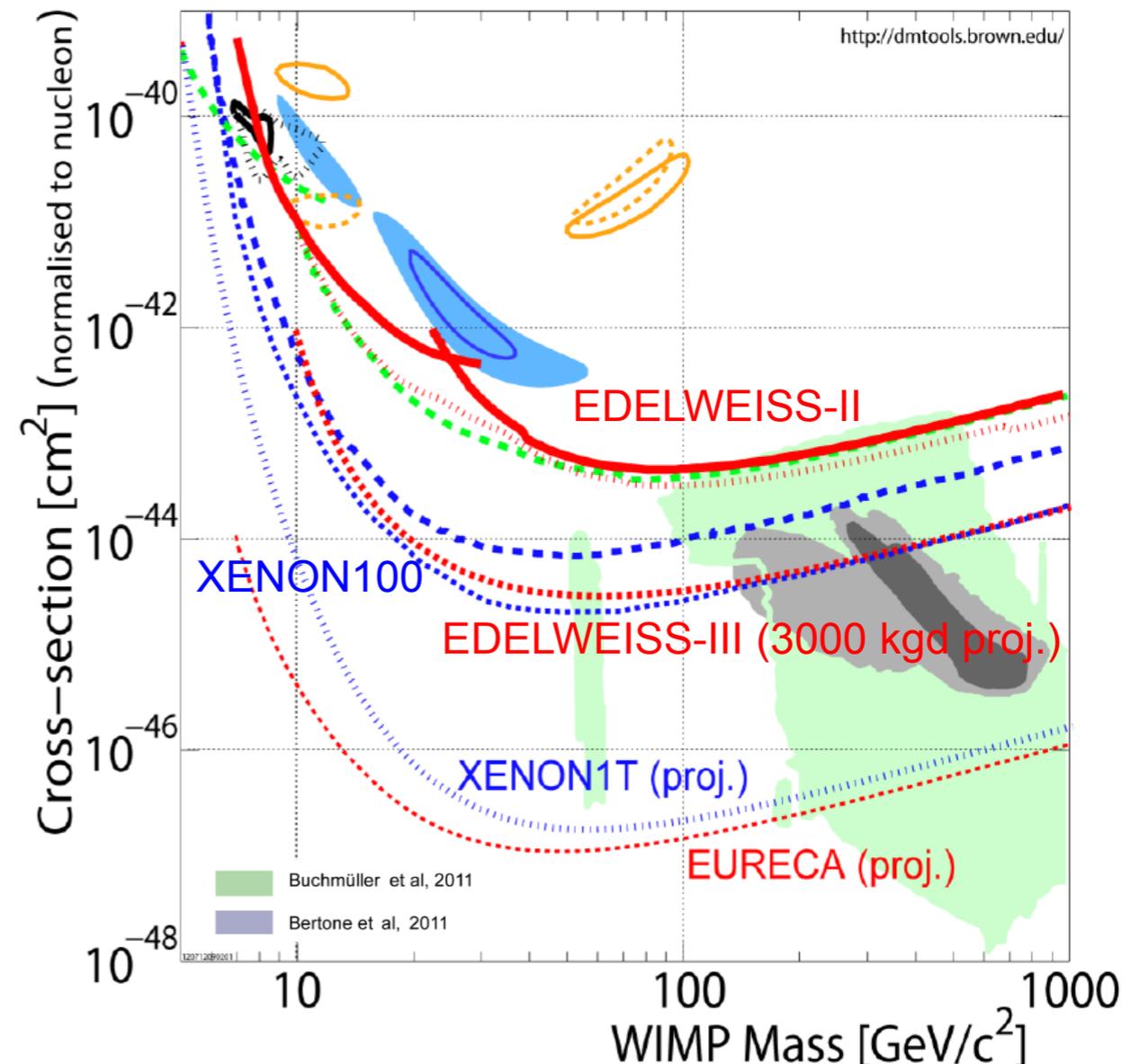
# EDELWEISS III - Projections

## Sept. 2013

- EDELWEISS III  
Commissioning runs underway
- 15 FID detectors of mass 800g each
- upgraded cryostat, readout electronics and kapton cables
- New PE shield and copper screens

## End of 2013

- Fully equipped cryostat with 40 FID detectors of 800g mass each.

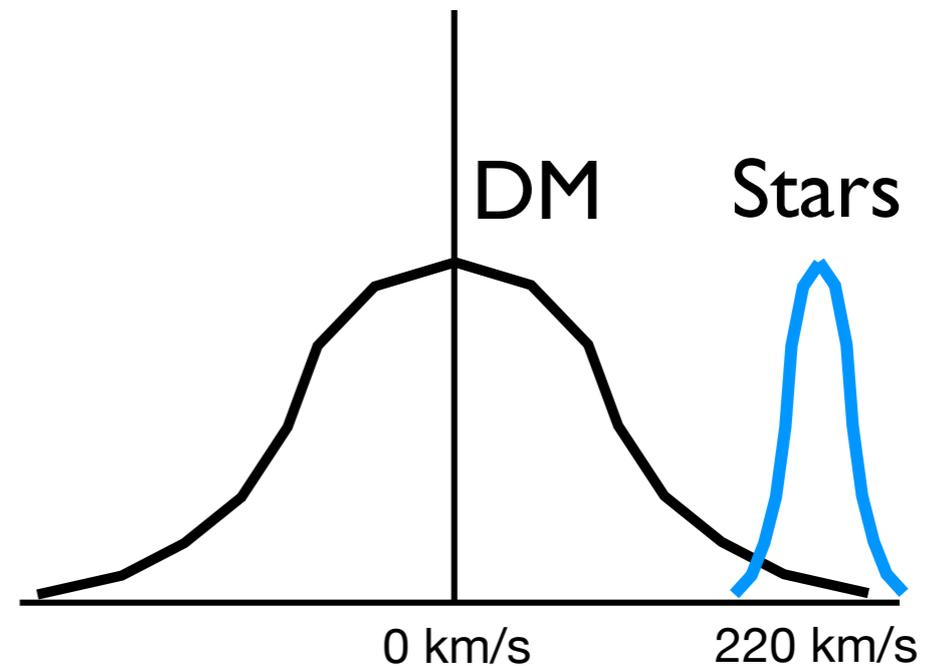


# The Experiments Part 2: Addressing a Long Standing Issue

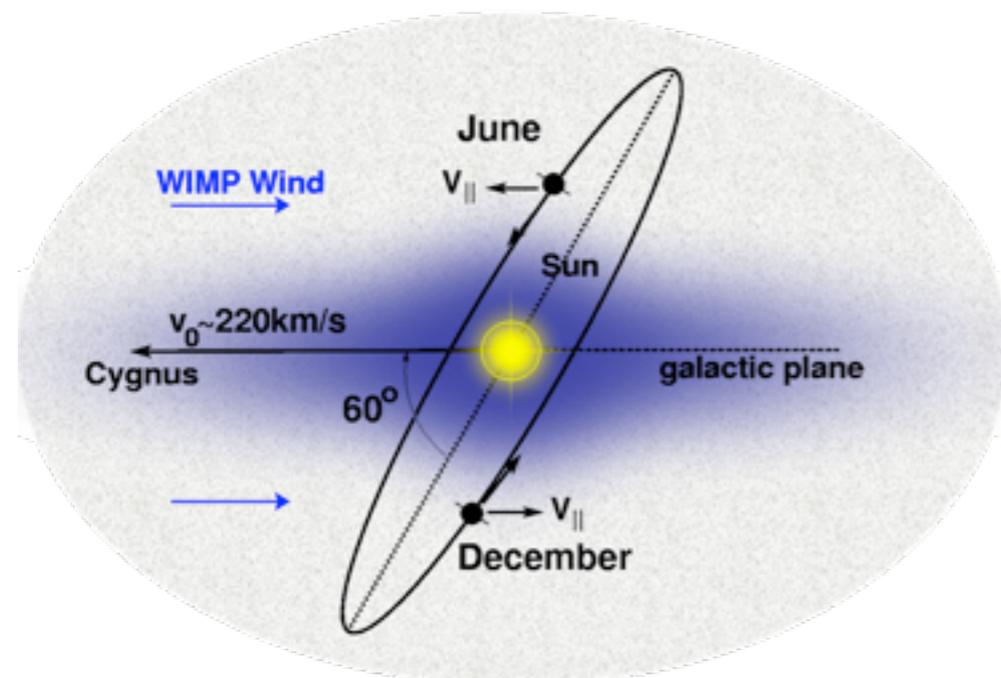
## DAMA/LIBRA Modulation Signal

# Signal Modulation

- Baryons travel together in roughly circular orbits with small velocity dispersion
- Dark matter particles travel individually with no circular dependence and large velocity dispersion



$V_\theta$  (at out galactic radius)



- As a result, the flux of WIMPs passing through Earth modulate over the course of a year as Earth rotates around the sun.